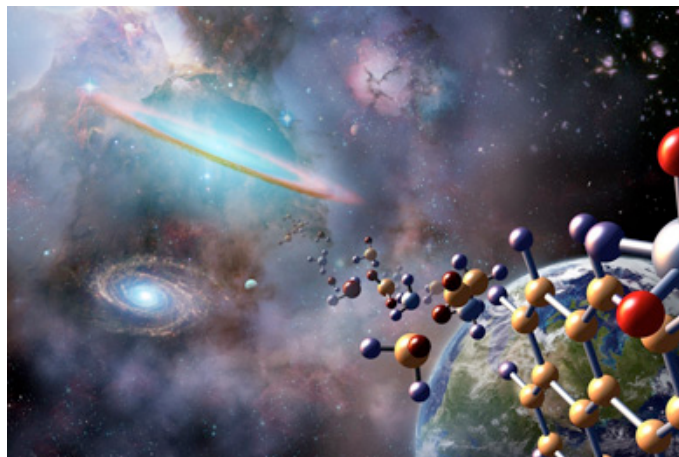


ORGANIC CHEMISTRY IN THE ATMOSPHERE OF THE EARLY EARTH

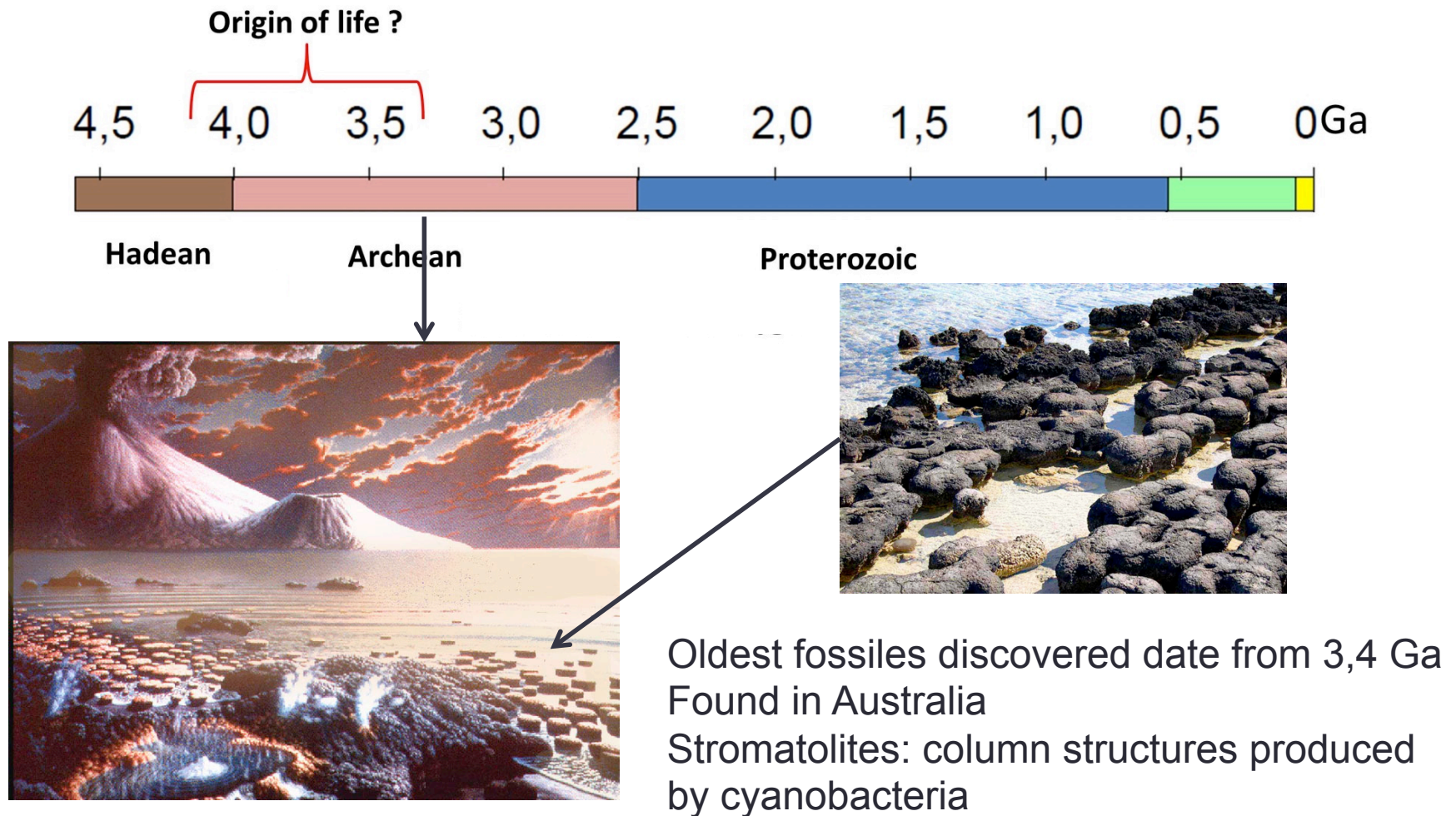


Prebiotic molecules in Space and Origins of Life on Earth
664. Wilhelm und Else Heraus-Seminar. Bad Honnef



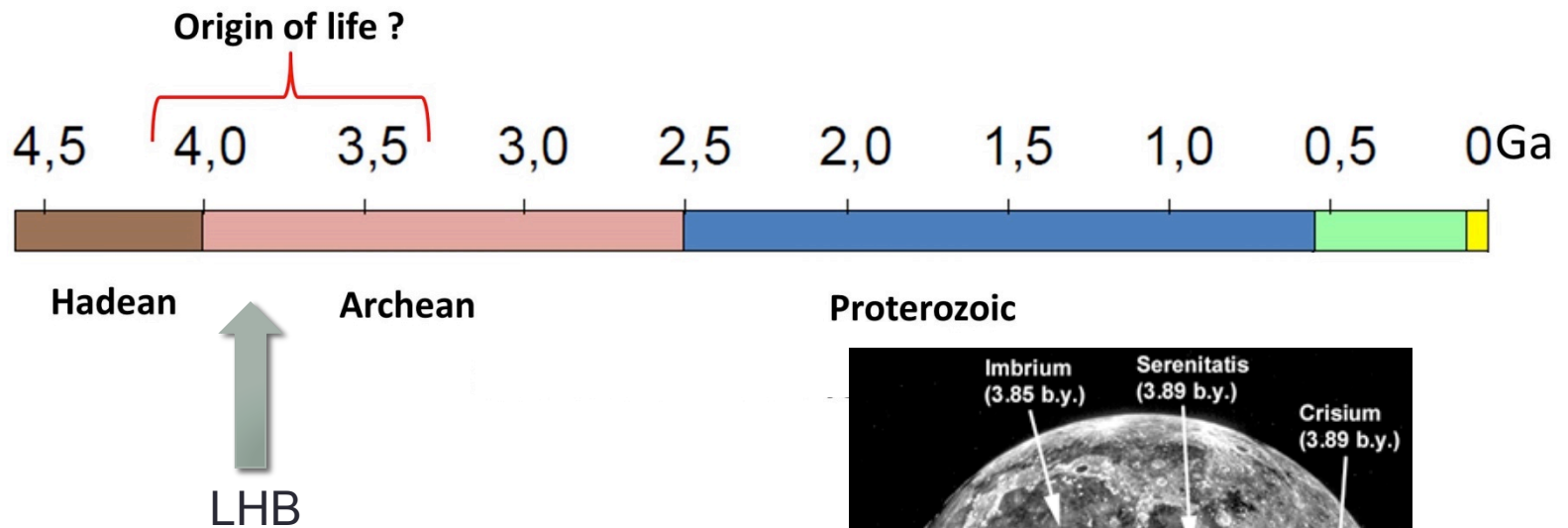
When life appears on Earth

Key for success: ocean + organic molecules



When life appears on Earth

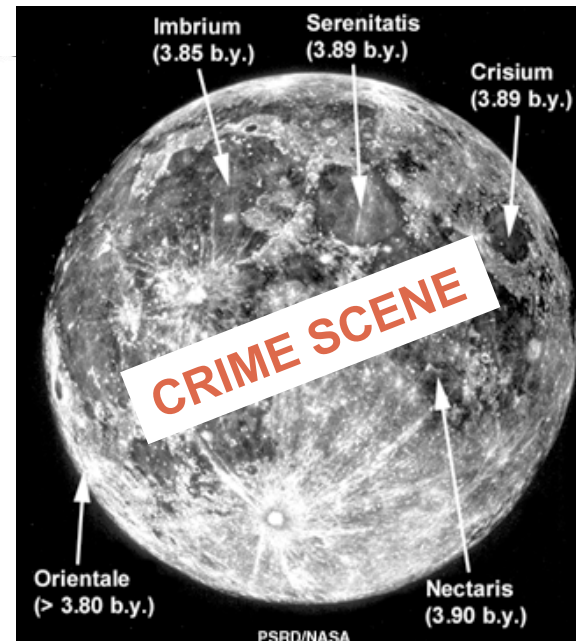
Key for success: ocean + organic molecules



Late Heavy Bombardment

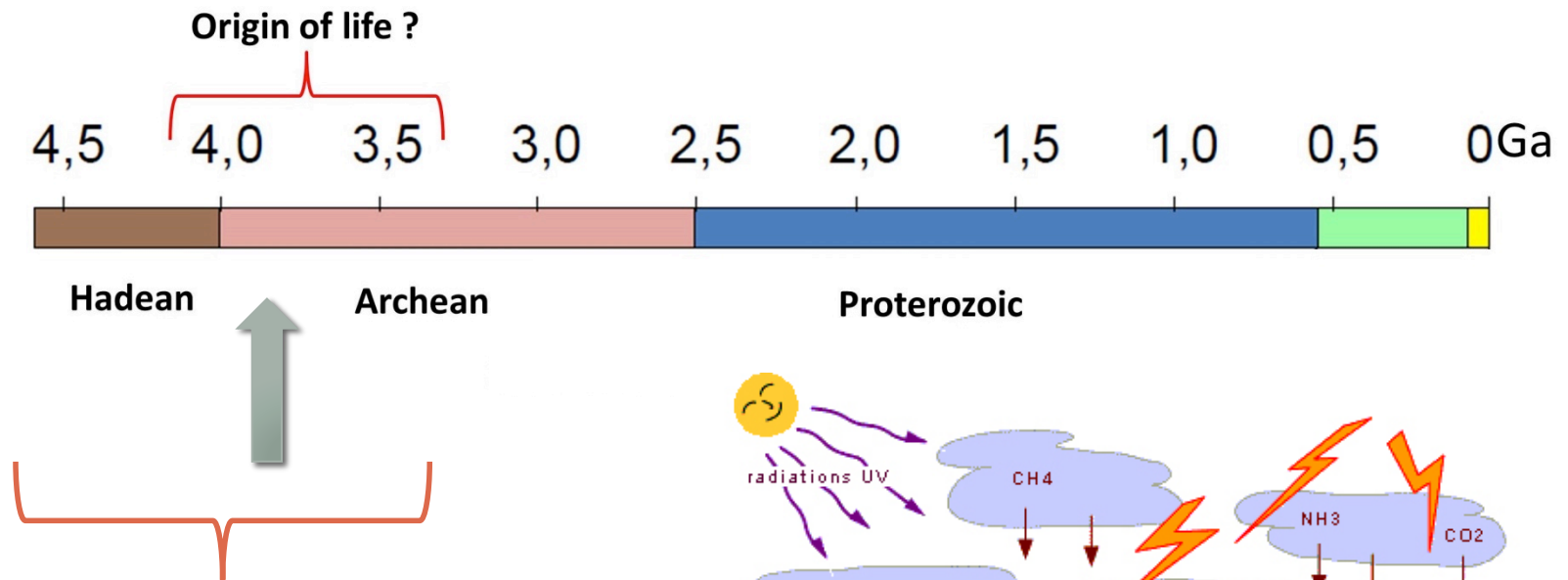
→ Late Veneer?

Volatiles are possibly added during this period

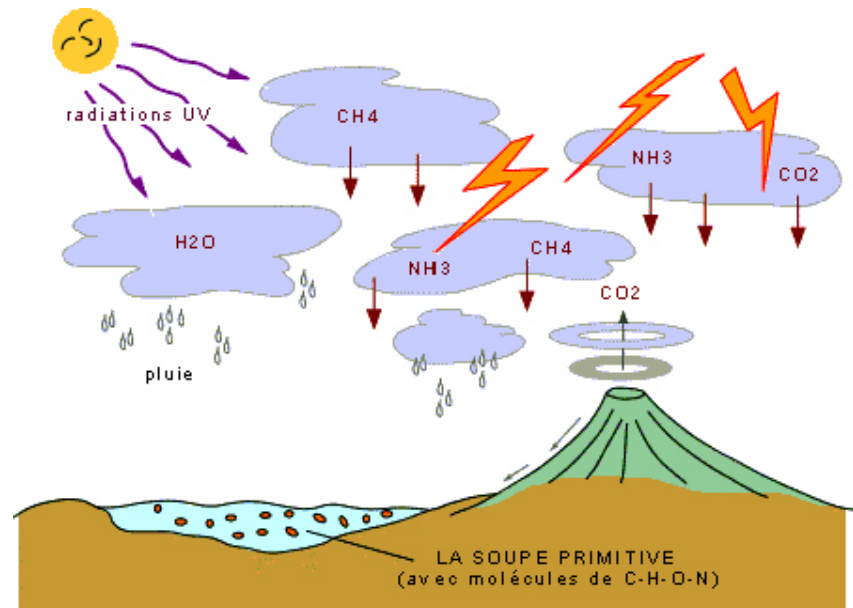


When life appears on Earth

Key for success: ocean + organic molecules



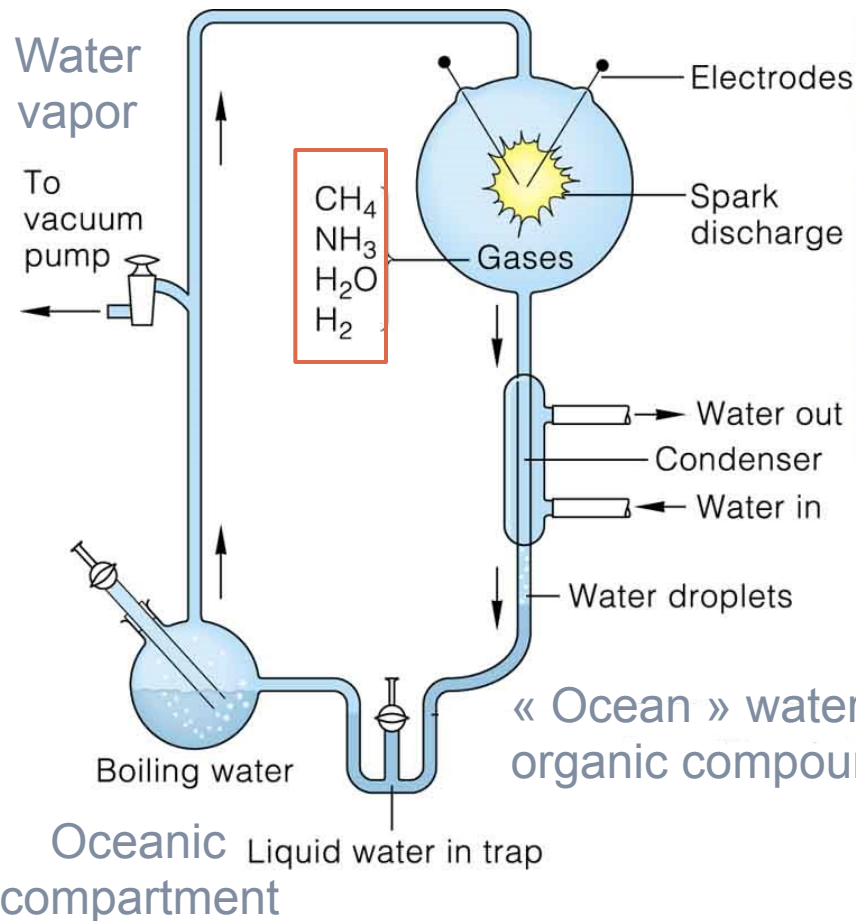
- Environmental conditions at this epoch?
- Role of the atmospheric chemistry & the primitive soup theory?



Oparin, 1924

Experimental simulation of the early Earth atmosphere

- Experiment by Miller&Urey (1953)

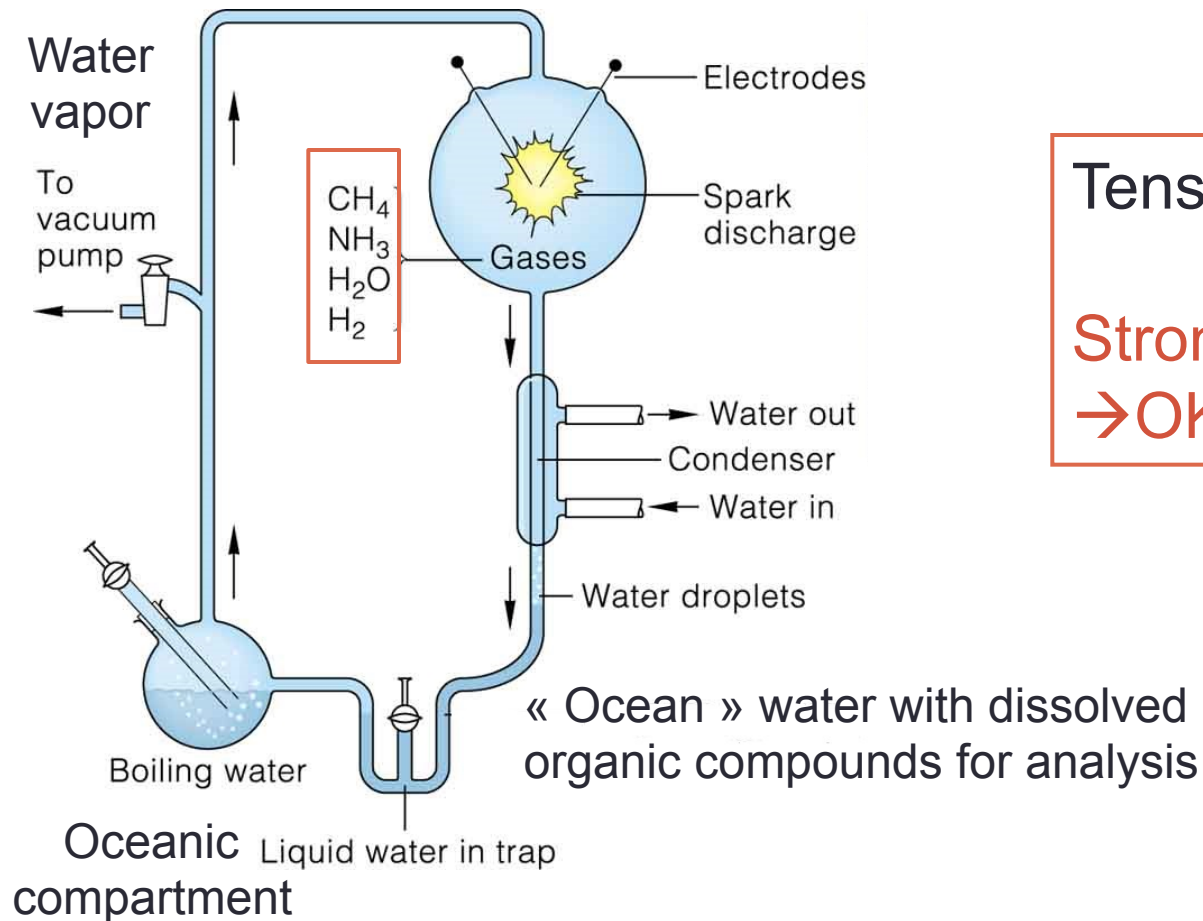


Primitive atmosphere
expected to be made of
 CH_4 , NH_3 , H_2O , H_2

=Analogy with giant planets
@time before H_2 escape

Experimental simulation of the early Earth atmosphere

- Experiment by Miller&Urey (1953)



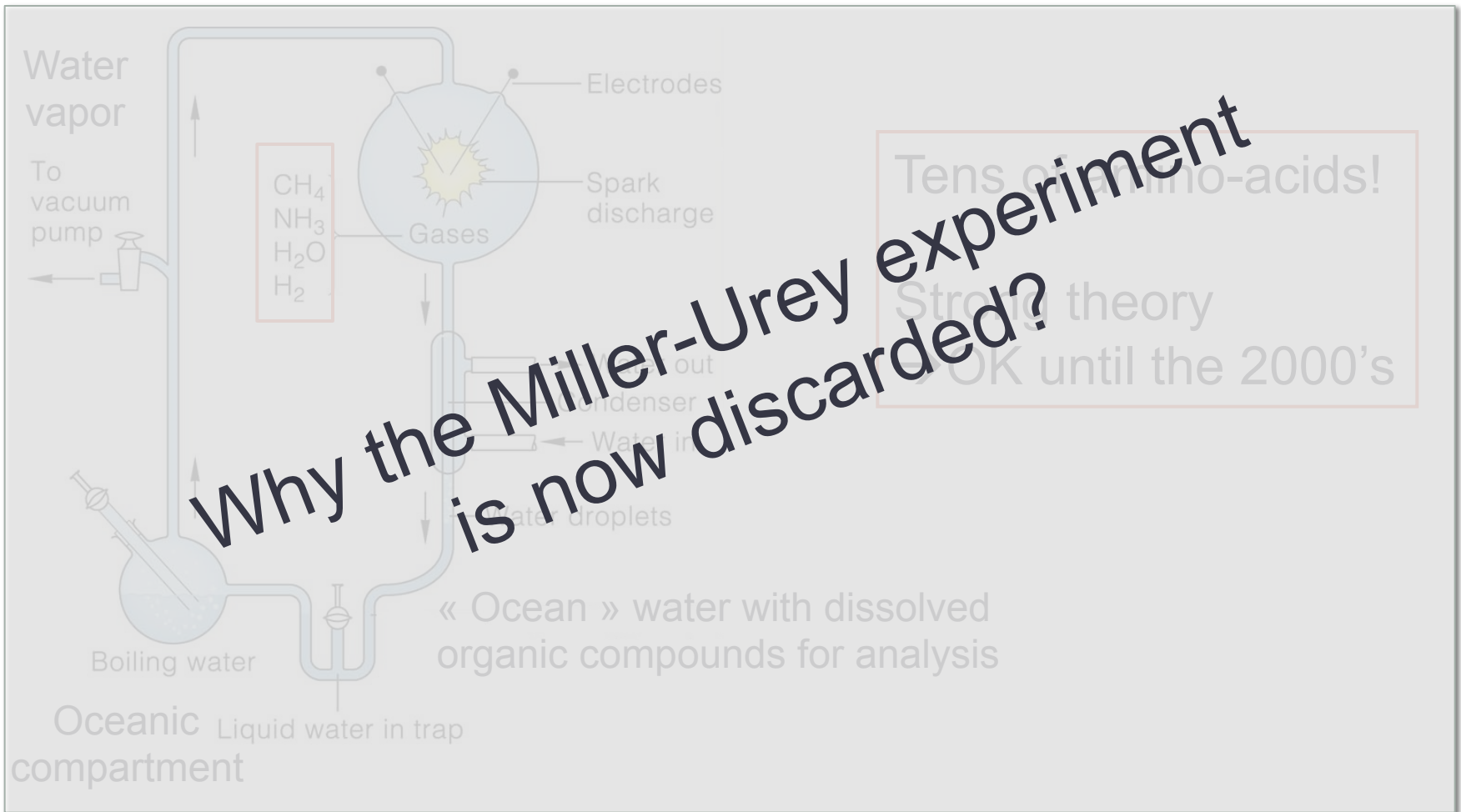
Tens of amino-acids!

Strong theory

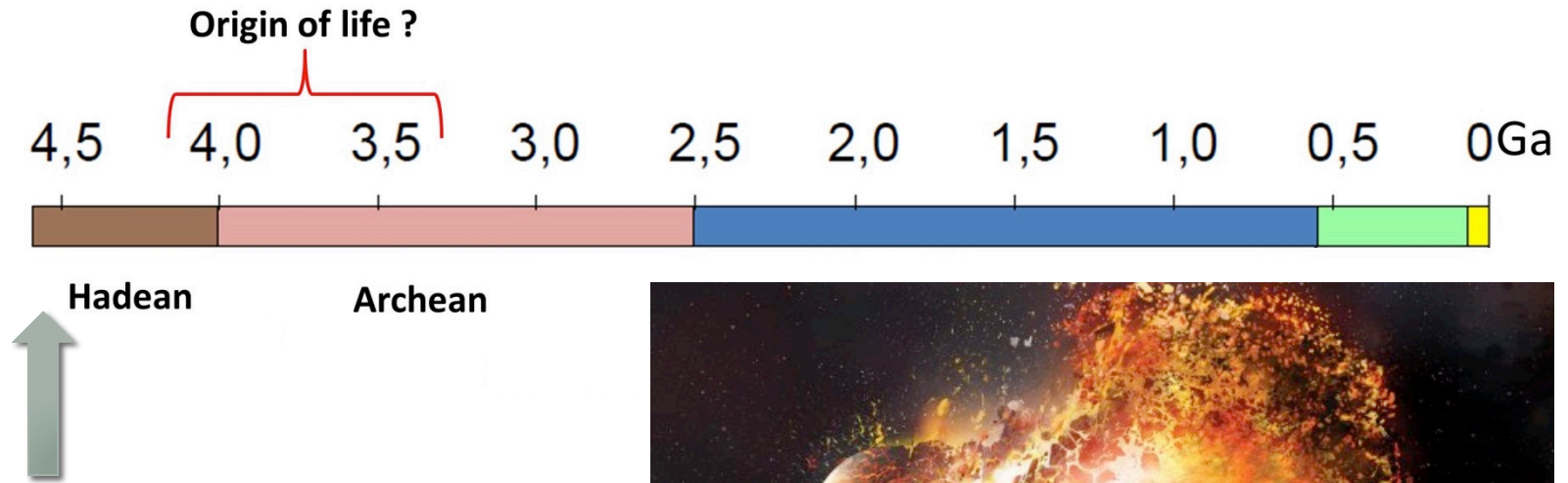
→ OK until the 2000's

Experimental simulation of the early Earth atmosphere

- Experiment by Miller&Urey (1953)



Massive escape of the proto-atmosphere



Large impacts
Degassed steam

→ No more primitive
reduced atmosphere



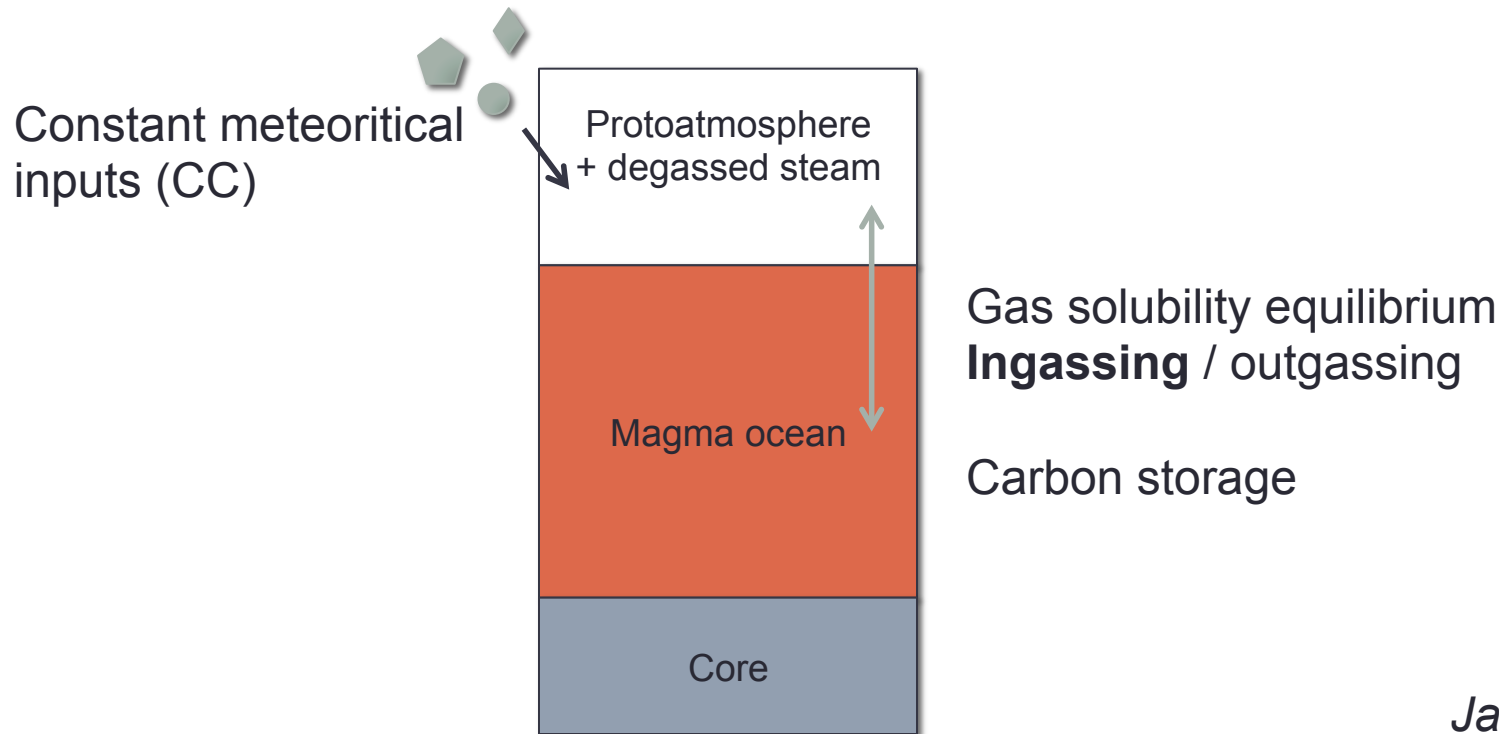
Artist view of the cataclysmic impact that
formed the Moon

Reconstruction of a 2ndary atmosphere

1. Magma ocean:

Ingassing process through gas-solubility equilibrium

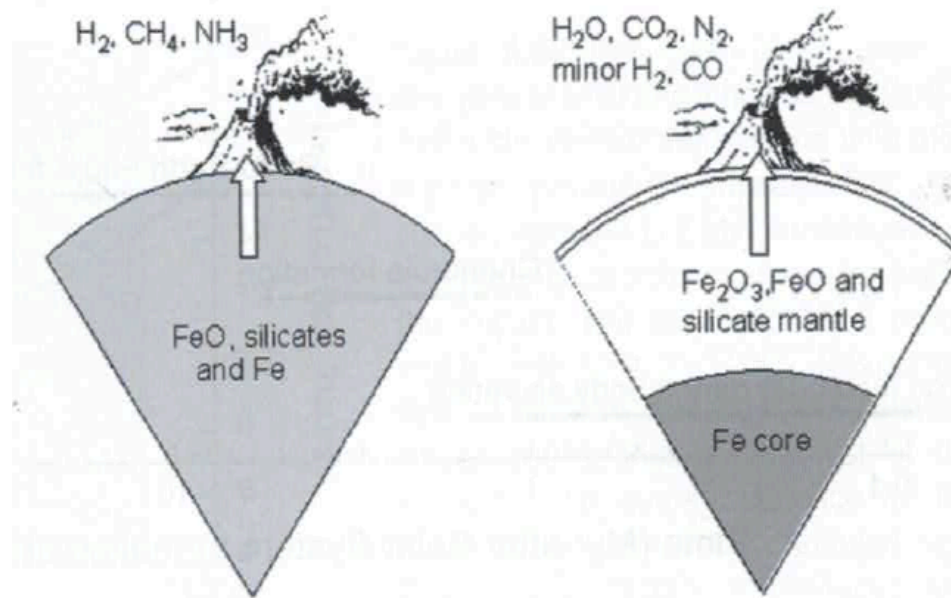
→ capture of a small portion of the primitive atmosphere



Jacobsen+ 2008
Dasgupta 2013

Reconstruction of a 2ndary atmosphere

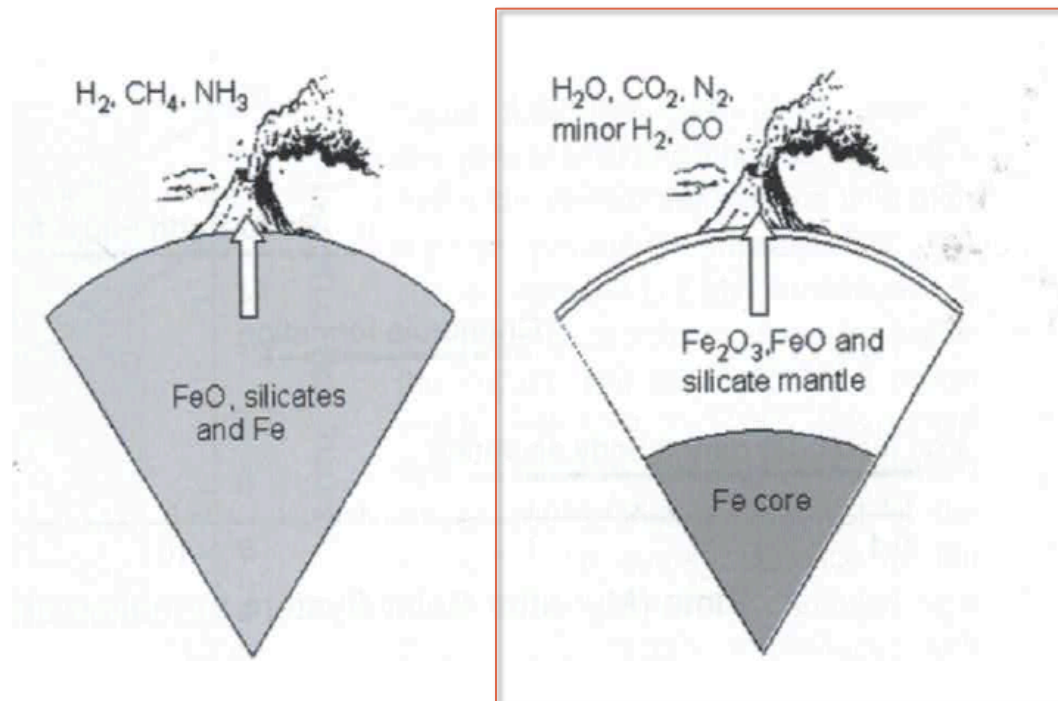
2. Outgassing during Hadean eon: formation of the ocean and the atmosphere



Depending on the oxydation state of the mantle, the degassed atmosphere can be either reduced or oxydized

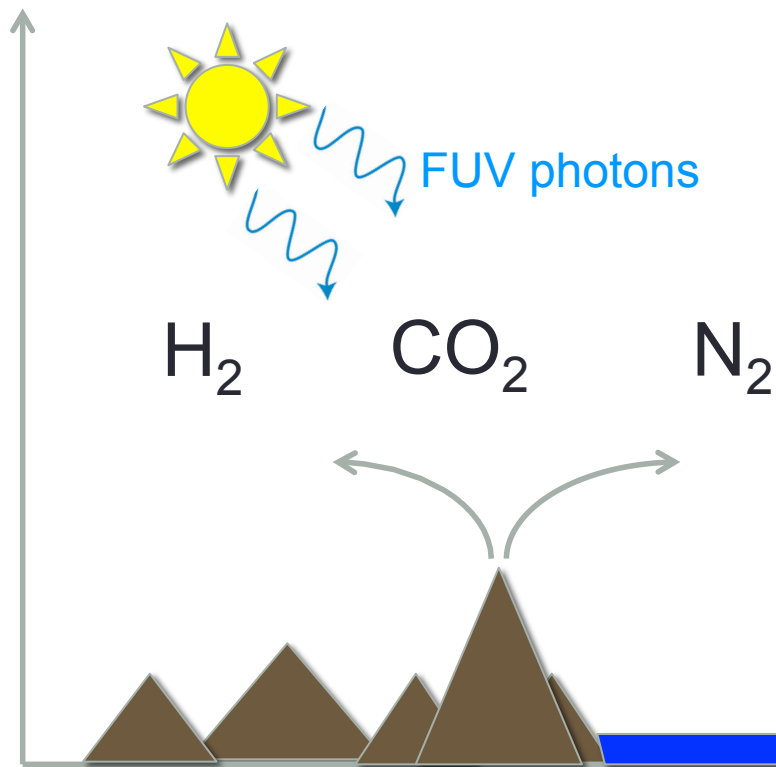
Reconstruction of a 2ndary atmosphere

2. Outgassing during Hadean eon: formation of the ocean and the atmosphere



Hadean oxygen fugacity similar as in present day mantle
(Tail et al. 2011)

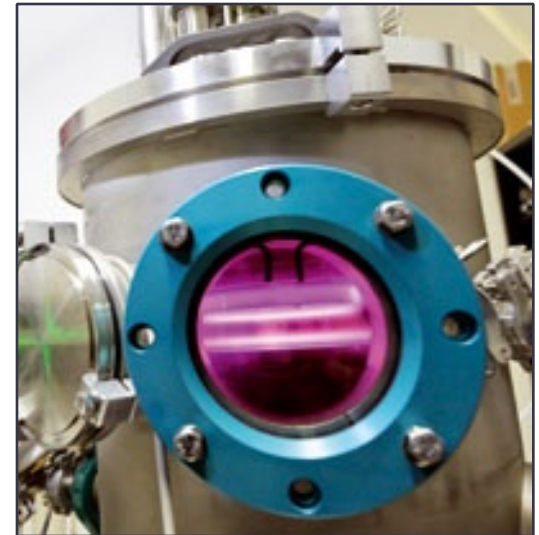
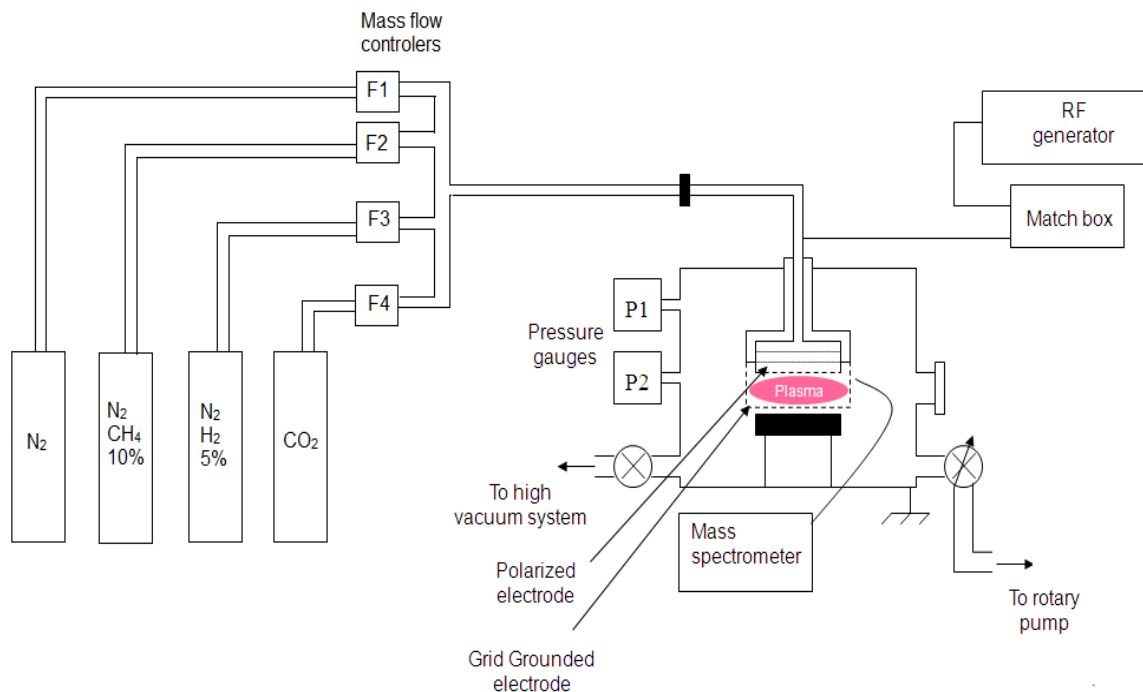
Organic chemistry in an oxydized atmosphere in the Hadean eon?



- High atmosphere submitted to harsh UV photons
- Example of unsuspected chemical growth in Titan thermosphere (talk by N. Balucani)

Experimental simulation of the chemistry in the early Earth high atmosphere

Benjamin Fleury, PhD thesis, 2015



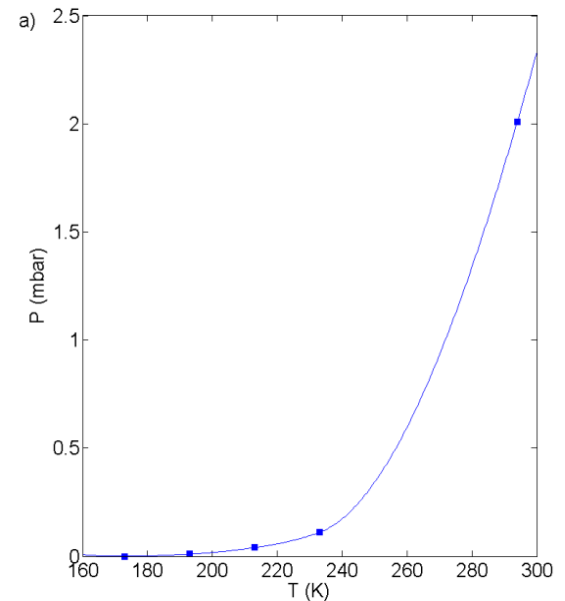
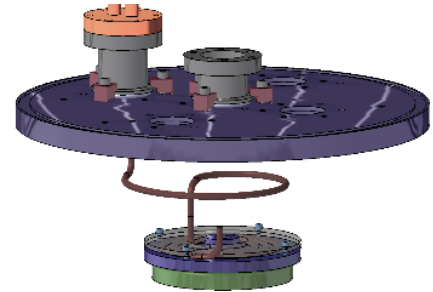
RF Plasma discharge: e⁻ as proxys for FUV photons

Gas mixture: $N_2 - CO_2 - H_2$ 91 - 5 - 4%

P=1 mbar, neutrals and ions at room temperature

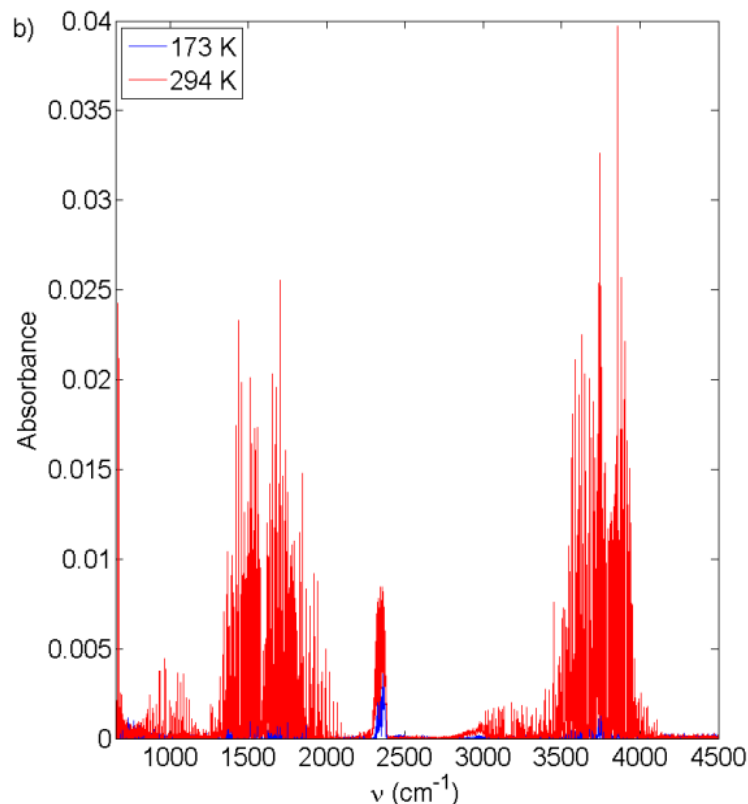
Cryogenic trapping of the gas products

- Cooling of the electrodes by a regulated LN₂ circulation:
T=173K
 - Accumulation of gas products
 - After 2 hrs of running experiment
 - Vacuum pumping
 - isolation of the reactor
 - Slow warming
- 2 mbar of products



Main product : Water !

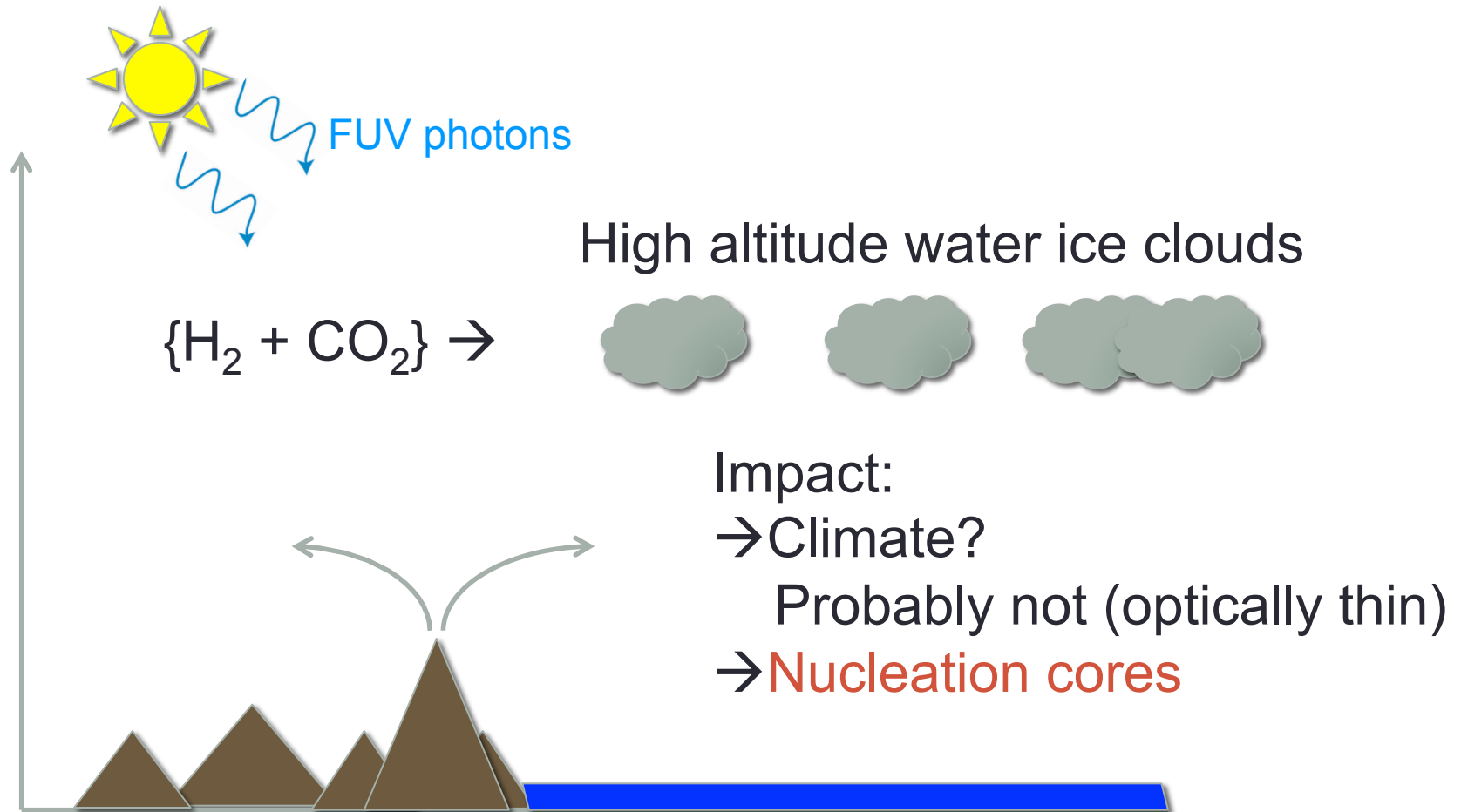
- Analysis by mid-IR spectroscopy



- Formation of O(¹D) from CO₂
 $\text{CO}_2 + \text{e}^- \rightarrow \text{CO} + \text{O}(\text{}^1\text{D}) + \text{e}^-$
 $\text{CO}_2 + \text{e}^- \rightarrow \text{CO}_2^+ + 2 \text{e}^-$
 $\text{CO}_2^+ + \text{e}^- \rightarrow \text{CO} + \text{O}(\text{}^1\text{D})$
- Formation of OH radicals
 $\text{H}_2 + \text{O}(\text{}^1\text{D}) \rightarrow \text{OH} + \text{H}$
- Production of water
 $\text{H}_2 + \text{OH} \rightarrow \text{H}_2\text{O} + \text{H}$
 $2 \text{OH} \rightarrow \text{H}_2\text{O} + \text{O}$

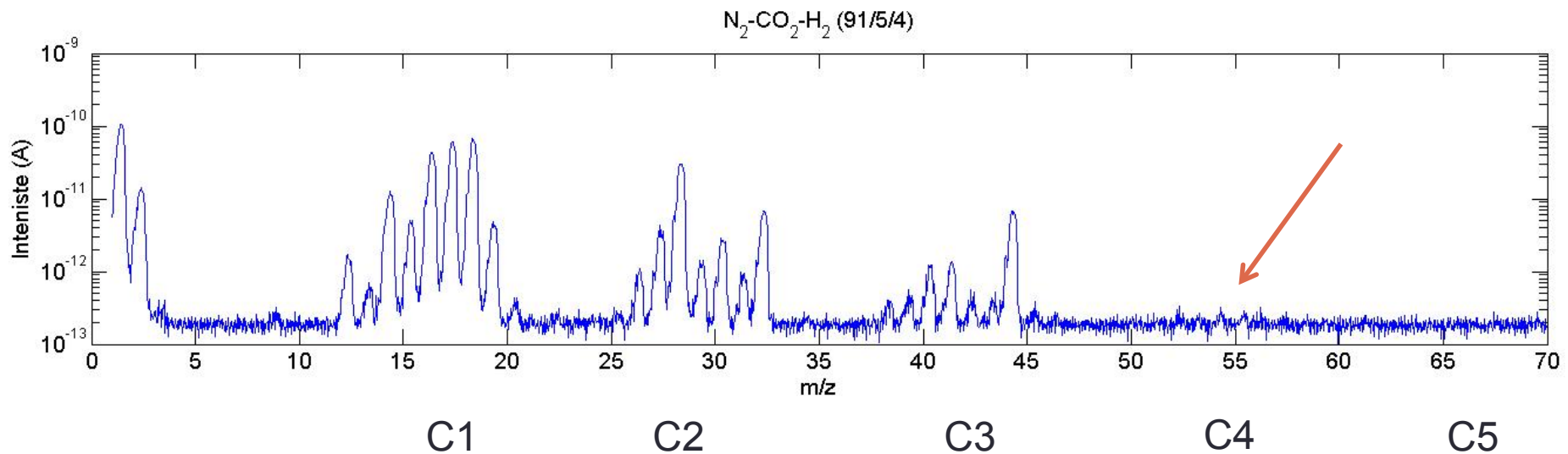
Fleury, Carrasco et al., ApJL, 2015

Unsuspected water formation at high altitude



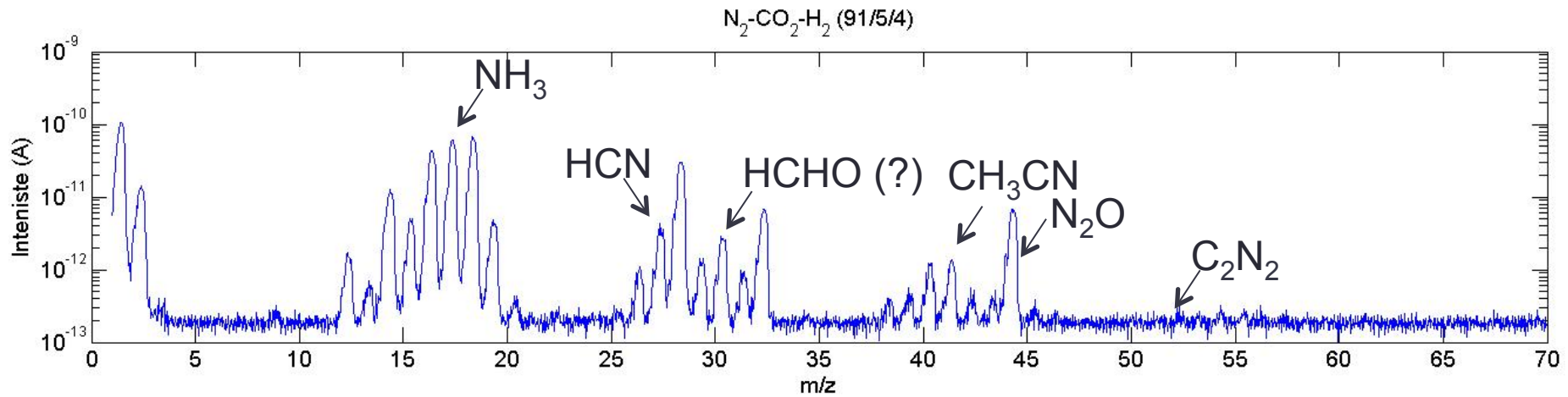
Fleury, Carrasco et al., ApJL, 2015

Analysis by *in situ* mass spectrometry



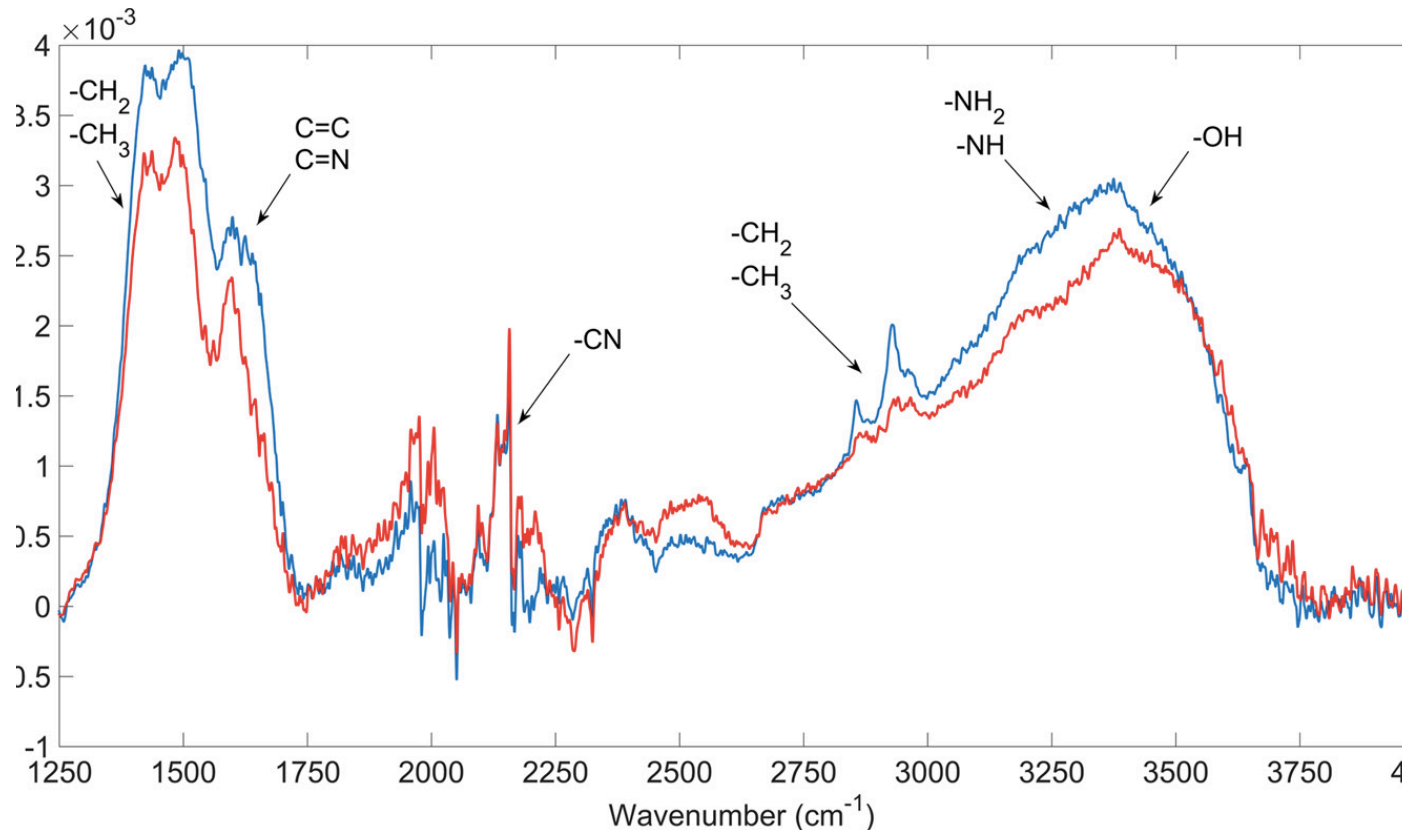
Detection of gas phase products up to C4
→CO₂ enables organic growth

Analysis by *in situ* mass spectrometry



- Identifications by mid-IR spectroscopy & GCMS
- Importance for prebiotic chemistry
 - Strecker synthesis (amino acids): NH_3 , HCN , HCHO
 - Formation of adenine (nucleic basis): HCN

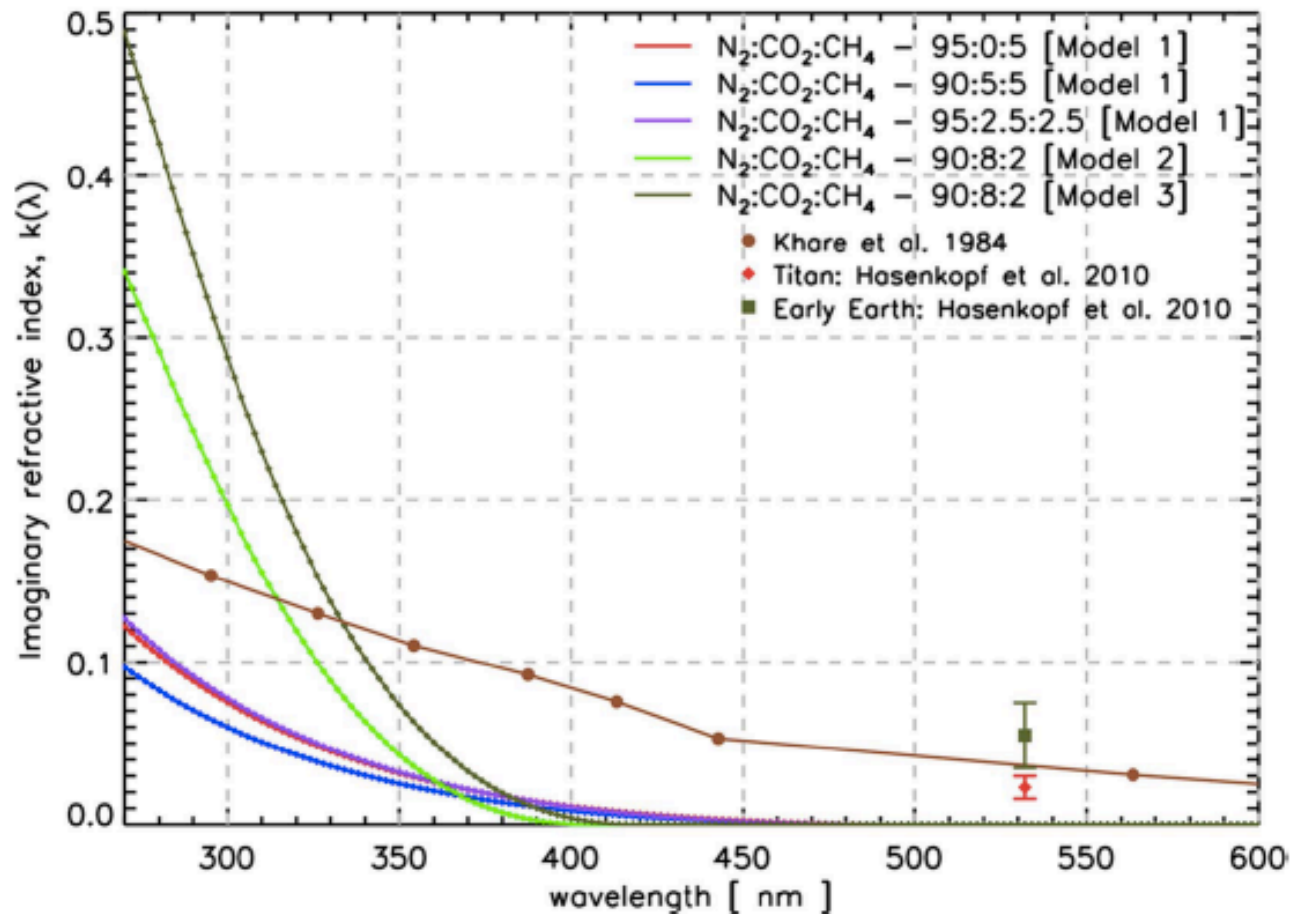
Production of solid organic material!



IR absorption spectra of the films deposited on 2 substrates in a $\text{N}_2 / \text{CO}_2 / \text{H}_2$ plasma (86 / 10 / 4%). Plasma duration = 40h

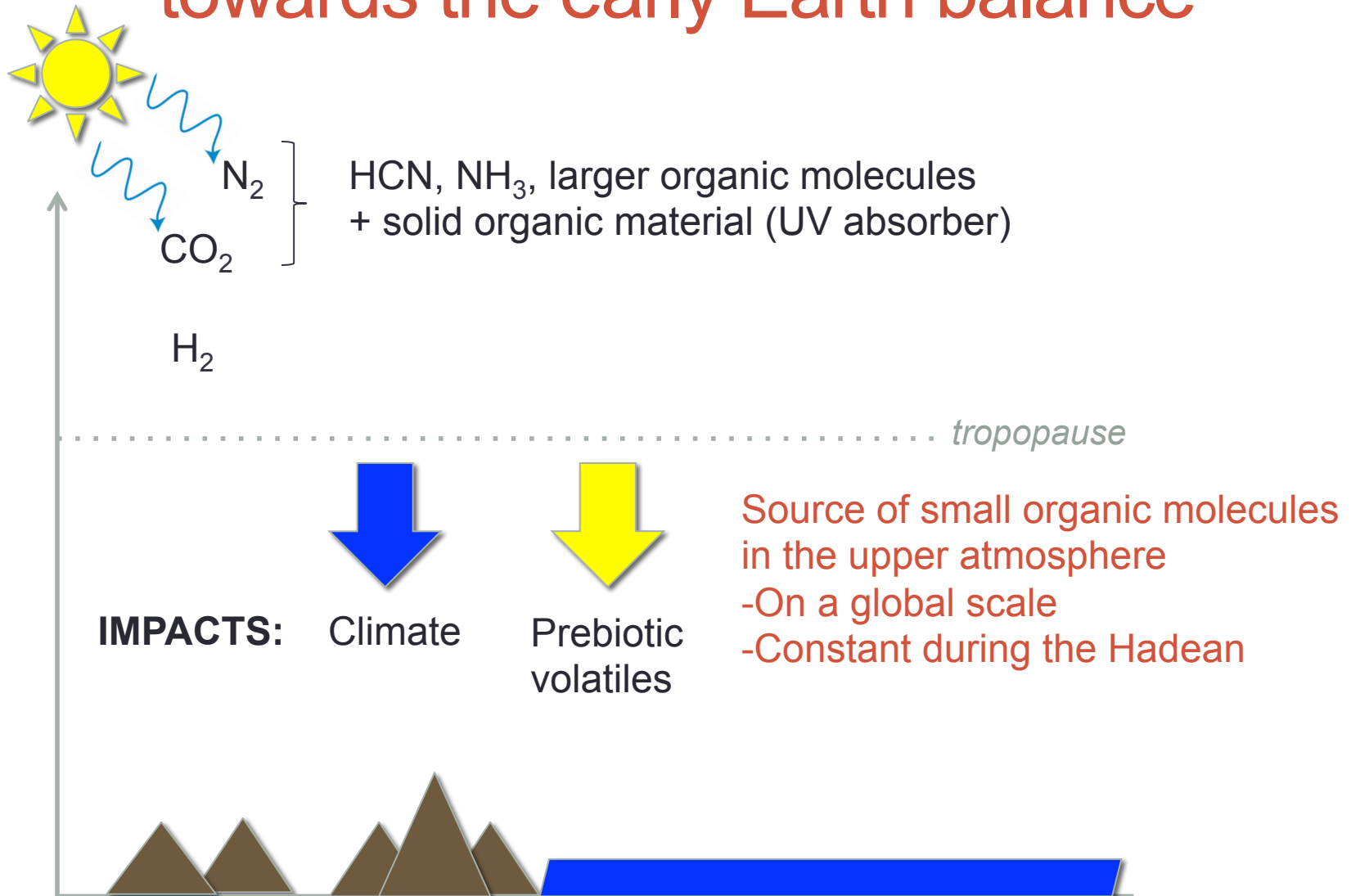
Fleury, Carrasco et al., EPSL 2017

Oxydized organic films: UV absorbers

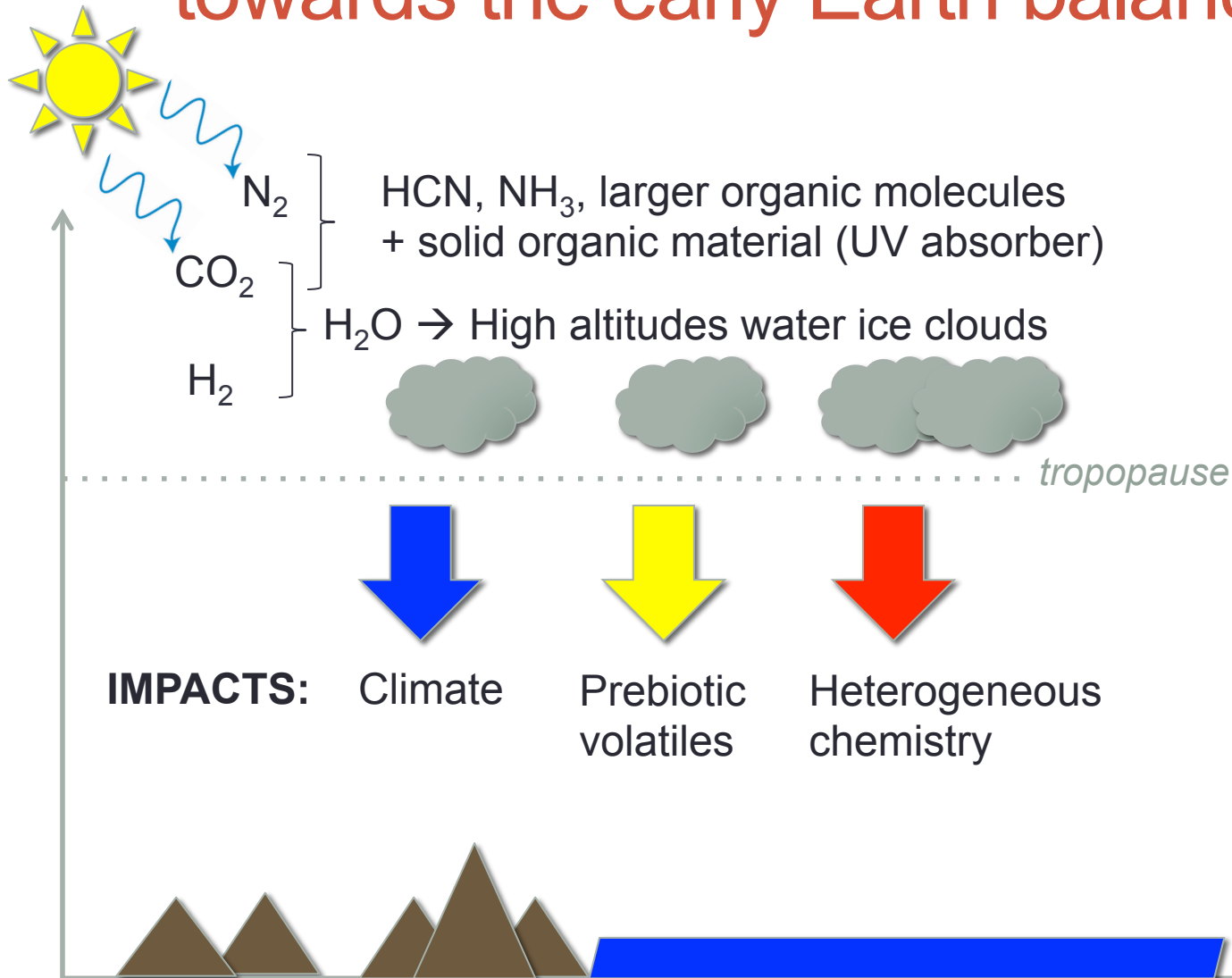


Gavilan, Broch, Carrasco et al. ApJL, 2017

Impacts of a CO₂-rich atmosphere towards the early Earth balance



Impacts of a CO₂-rich atmosphere towards the early Earth balance



Conclusion :

The upper atmosphere of the CO₂-rich early Earth as an unsuspected source for water ice and organic molecules

- Small molecules in the gas phase: NH₃, HCN
- Ubiquitous and permanent during the Hadean eon
- Not to be neglected compared to exogeneous sources

