

Interstellar ices as a source of complex organic molecules of interplanetary solar system objects

Grégoire DANGER

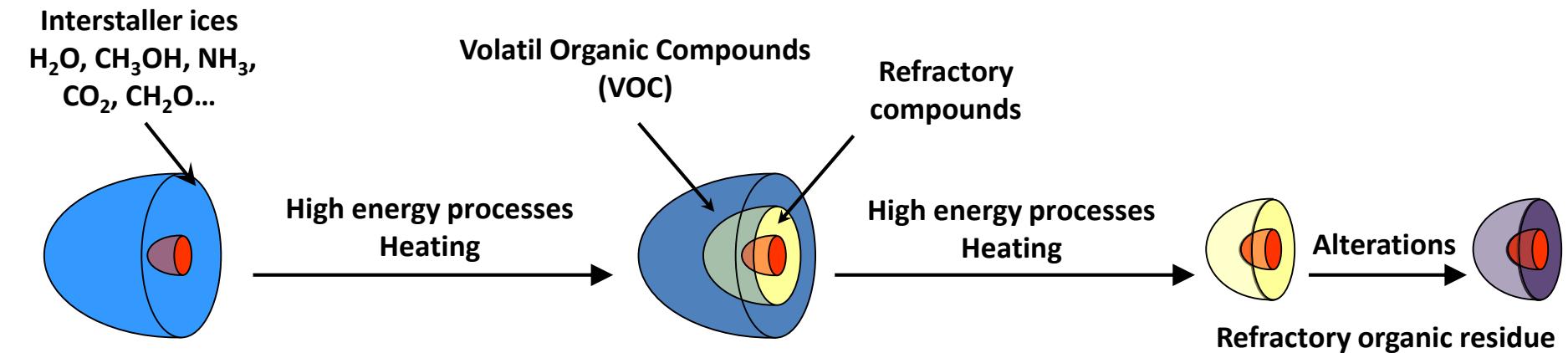
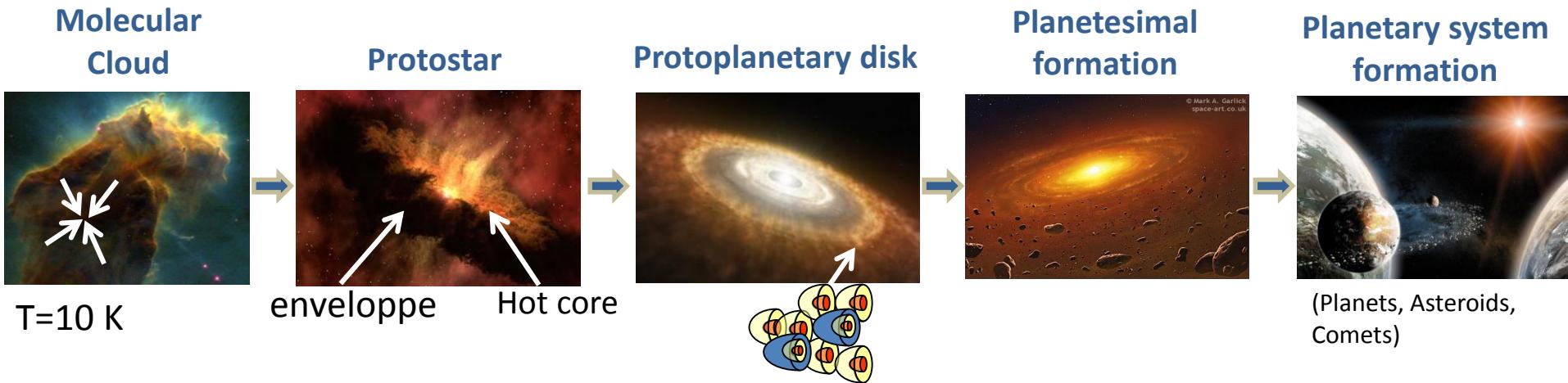
N. Abou Mrad, A. Fresneau, F. Duvernay, L. d'Hendecourt

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<http://piim.univ-amu.fr/Equipe-ASTRO>



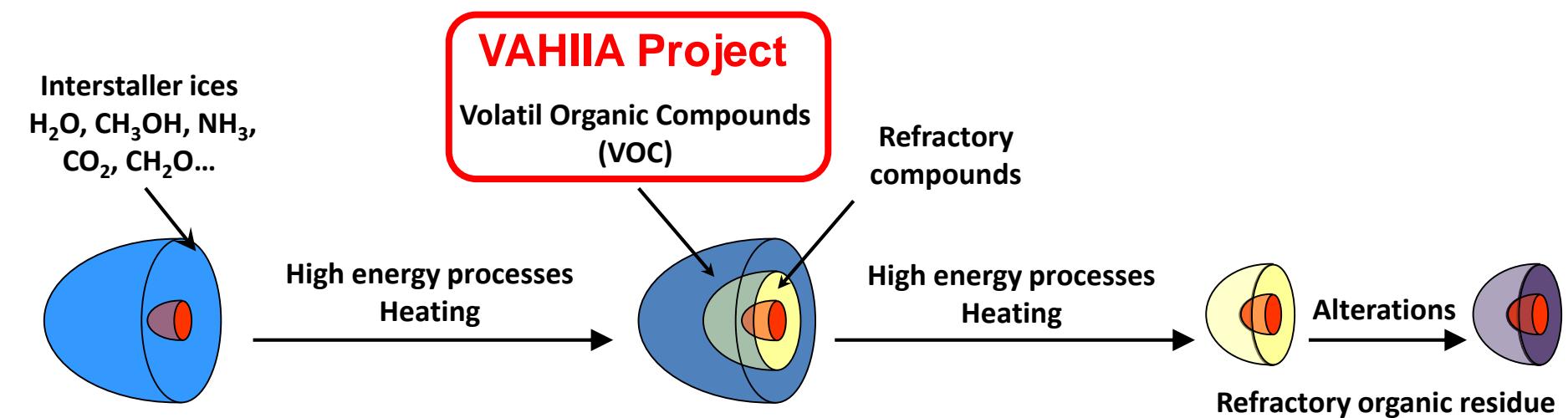
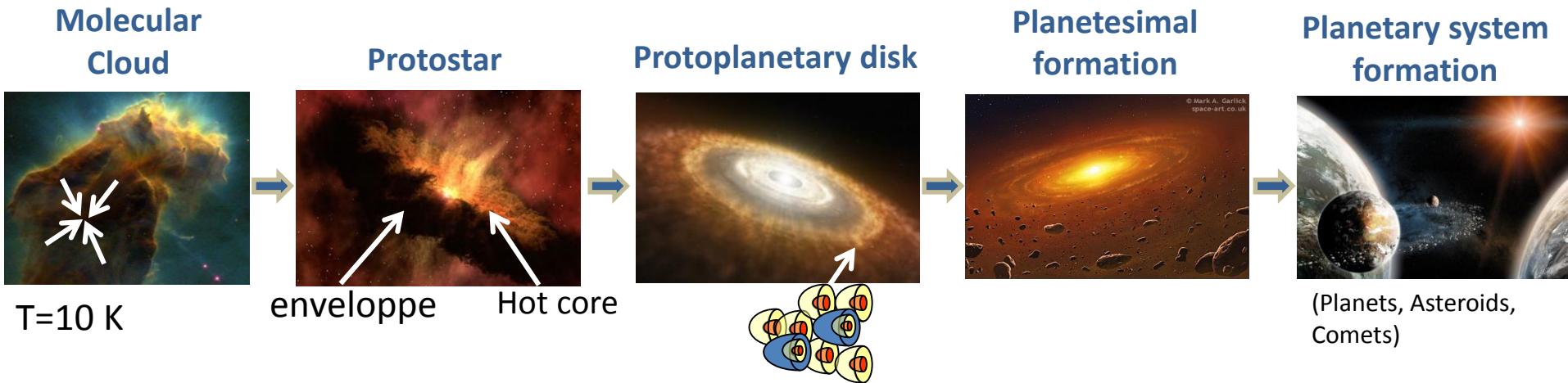
Evolution of interstellar icy grains

Toward the formation of complex organic matter in interplanetary bodies

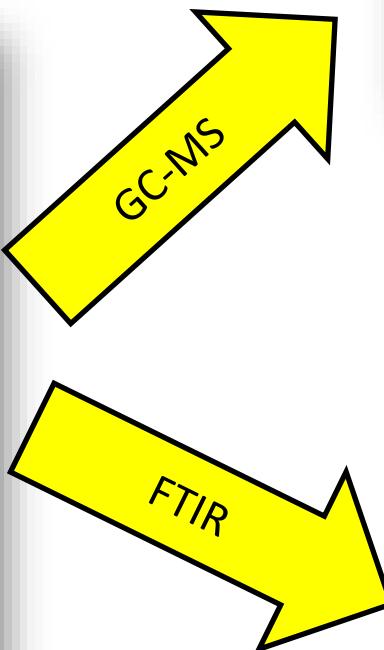
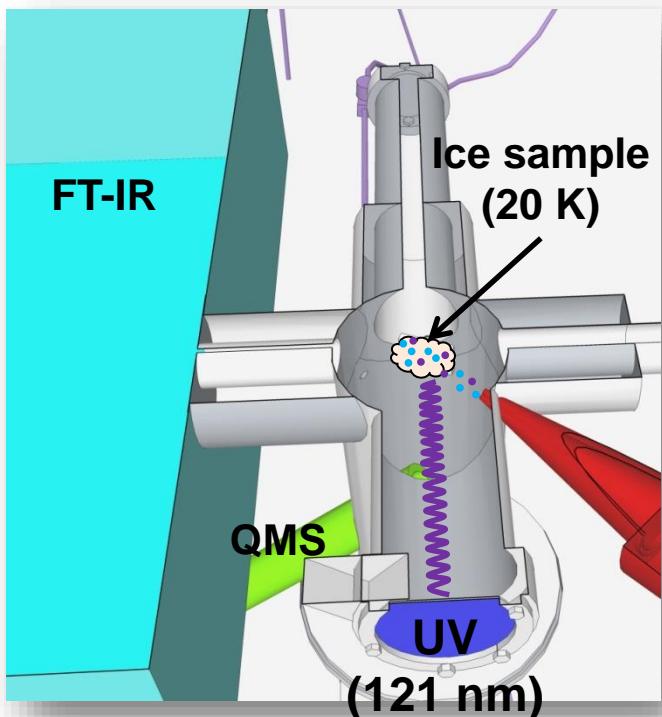


Evolution of interstellar icy grains

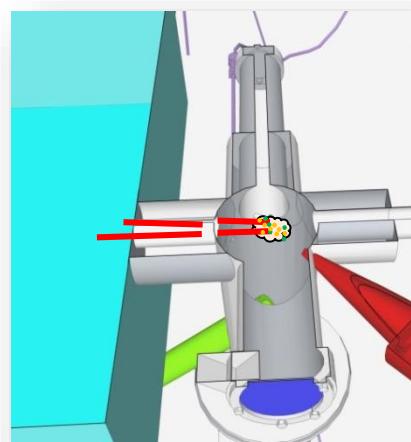
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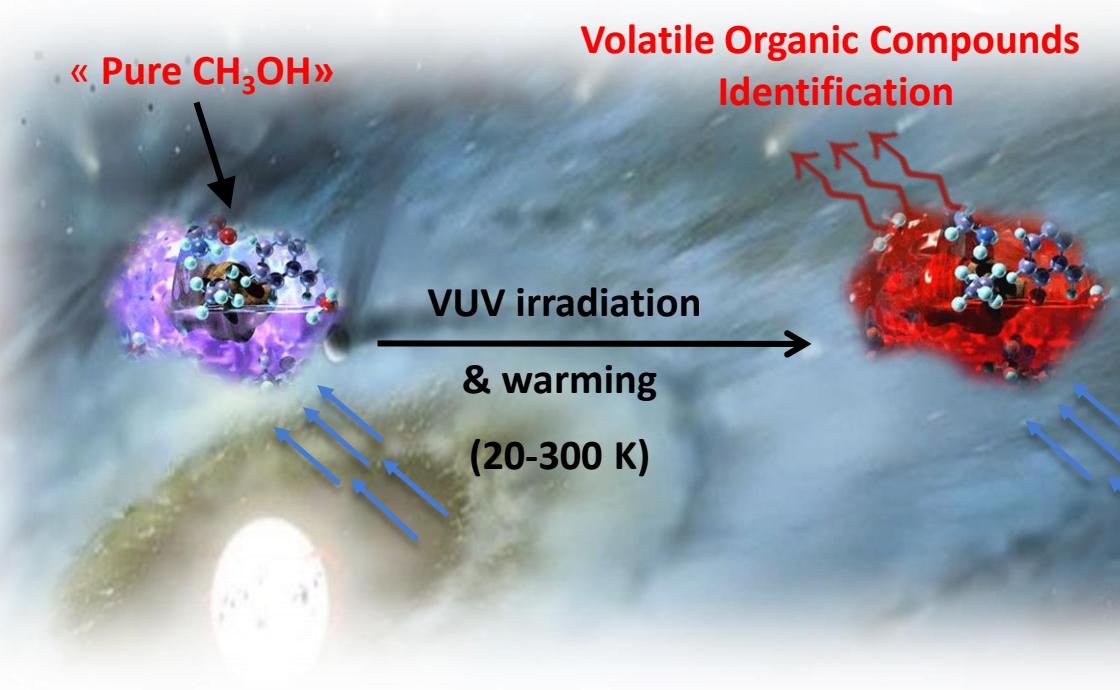
Vacuum chamber (10^{-9} mbar)



Ice formation in simulated conditions



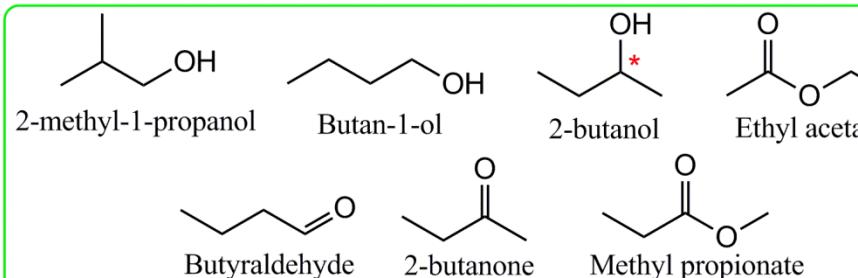
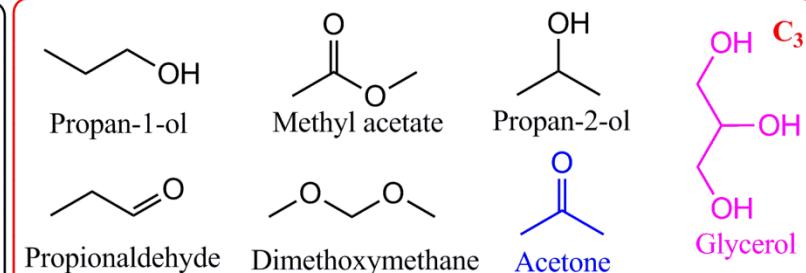
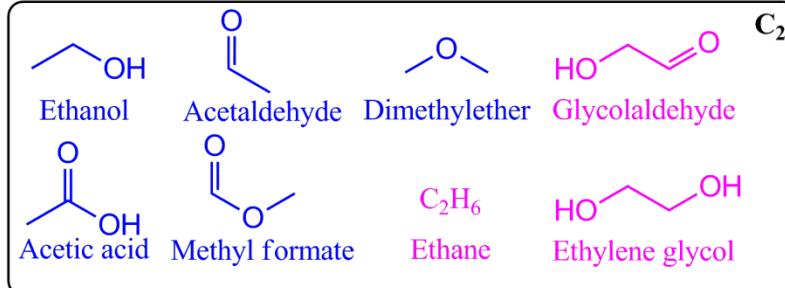
Methanol CH₃OH
An abundant source of reduced carbon in
interstellar and cometary ices



Methanol CH₃OH

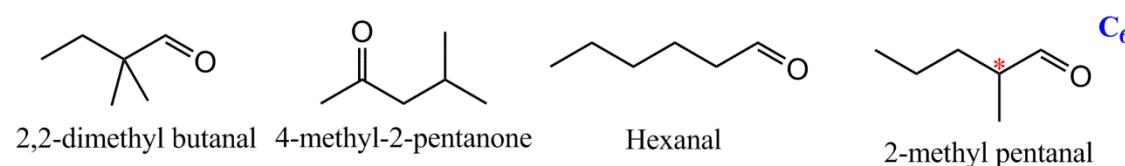
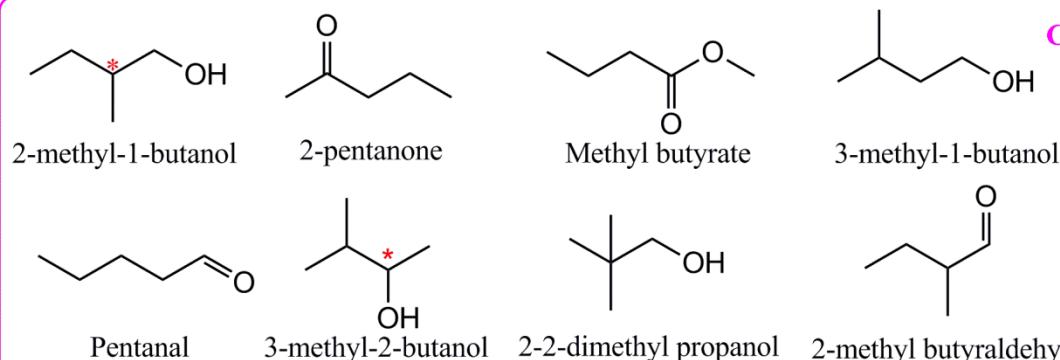
Analysis with the VAHIIA system – products identified

CO	C ₁
Carbon monoxide	
CO ₂	
Carbon dioxide	
CH ₂ O	
Formaldehyde	
HCOOH	
Formic acid	
CH ₄	
Methane	

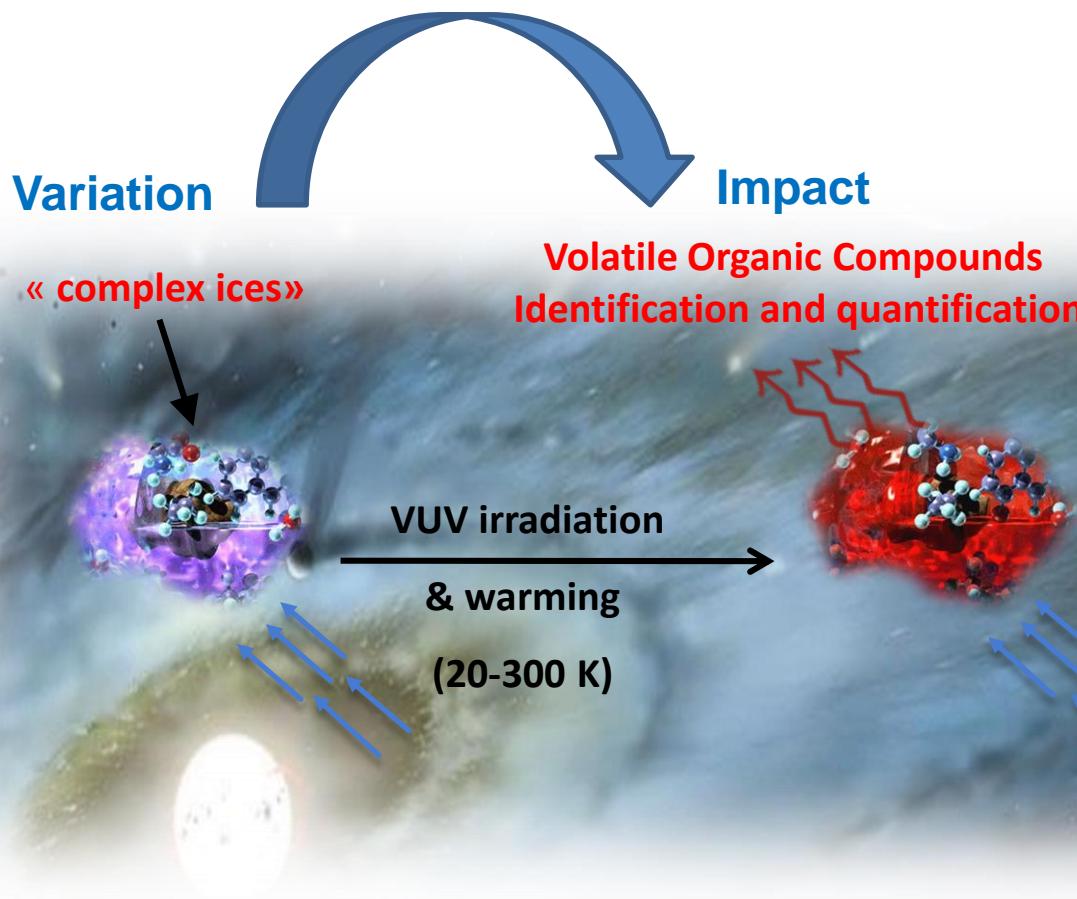


35 molecules identified

- Gerakines et al., A&A, 1996, 312, 289
 Oberg et al., A&A, 2009, 504, 891
 Henderson et al., ApJ, 2015, 800, 66
 Maity et al., PCCP, 2015, 17, 3081
 Kaiser, Angew. Chem., 2015, 54, 195



Complex ices
**Impact of the ice composition on VOC
abundances detected in the gas phase**

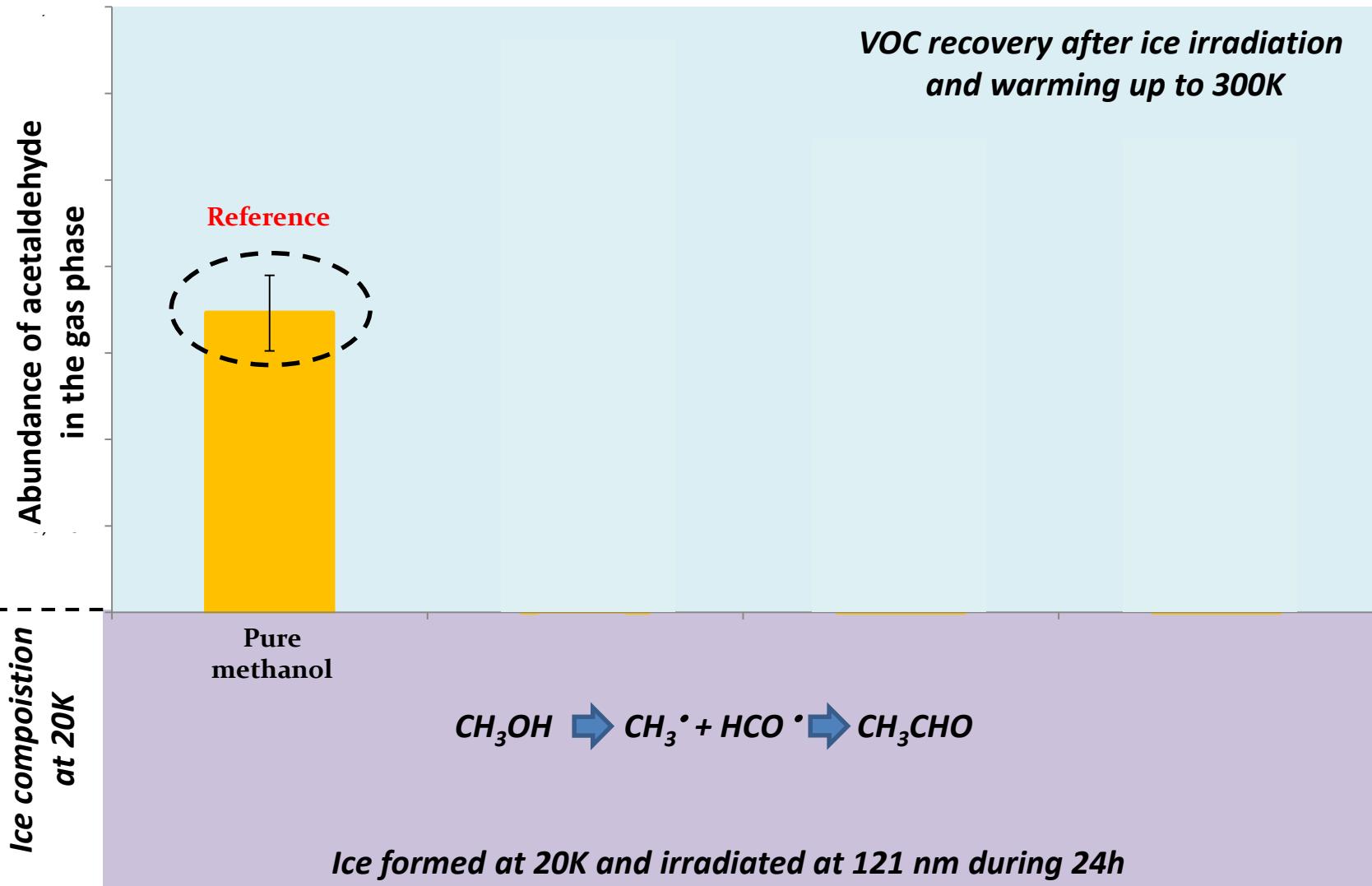


Complex ices
Impact of the ice composition on VOC
abundances detected in the gas phase

Abou Mrad et al., ApJ, 2017, 846, 124

VOC example: Acetaldehyde

*VOC recovery after ice irradiation
and warming up to 300K*

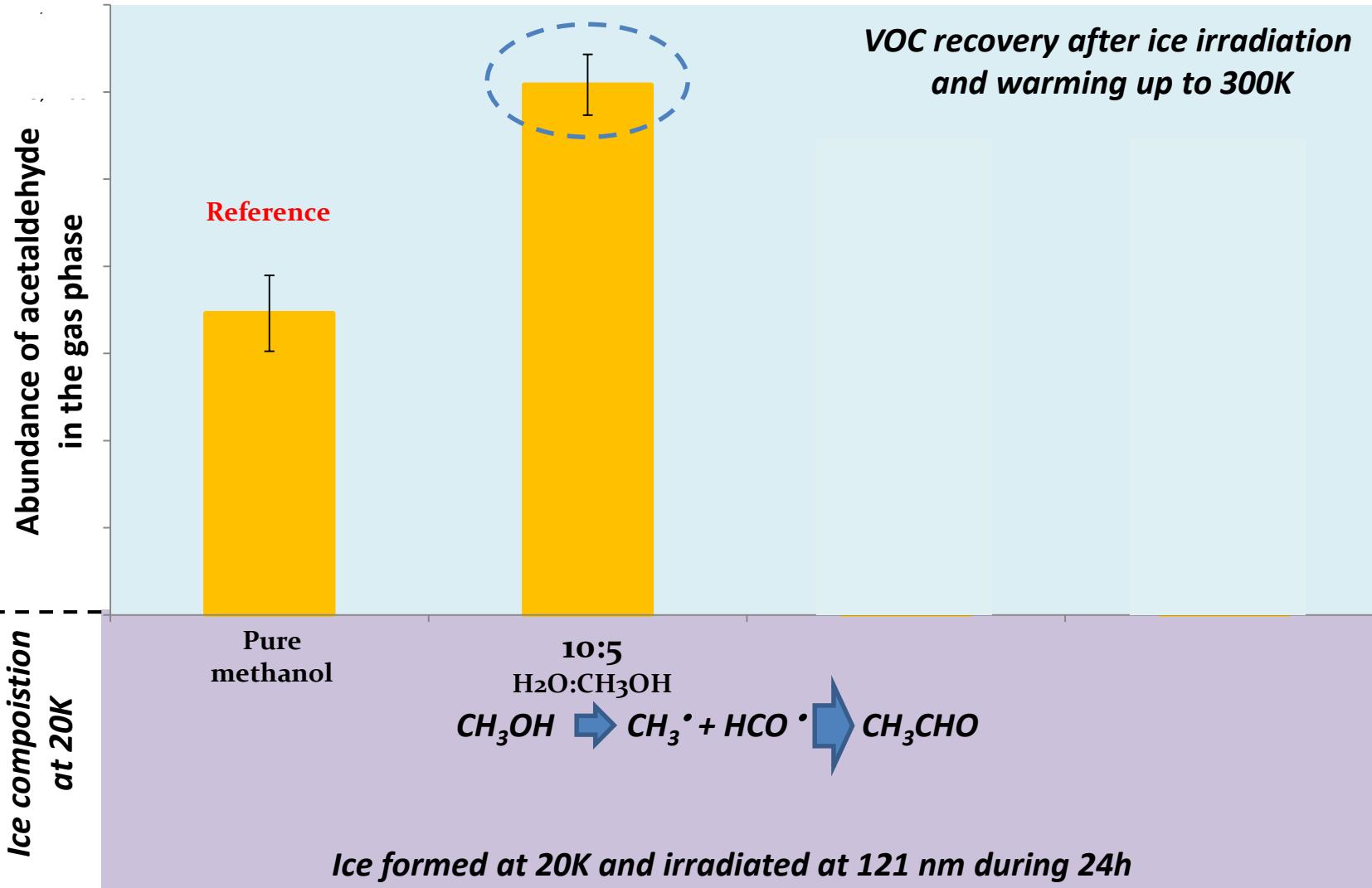


Complex ices

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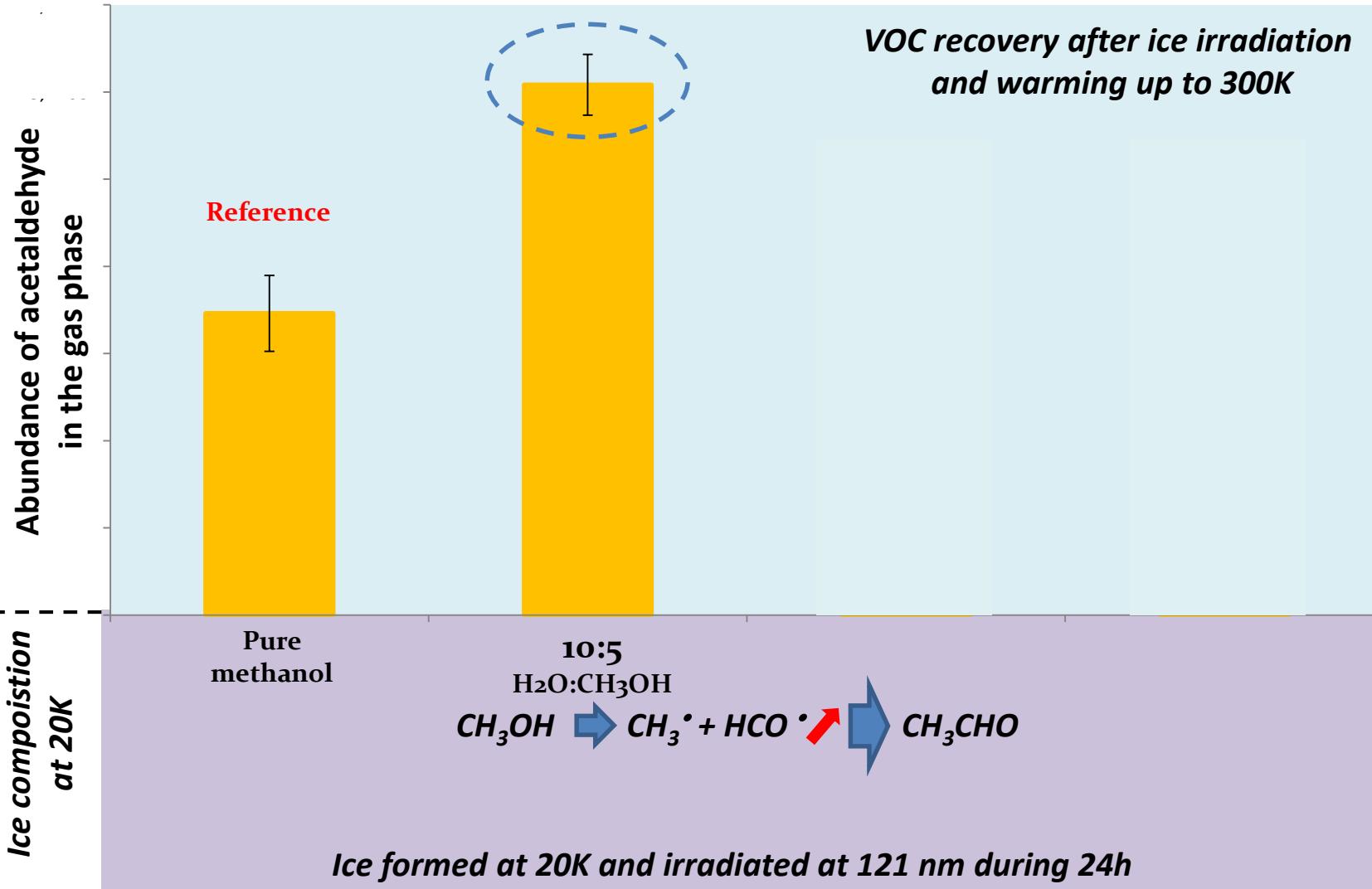


Complex ices

Impact of the ice composition on VOC abundances detected in the gas phase

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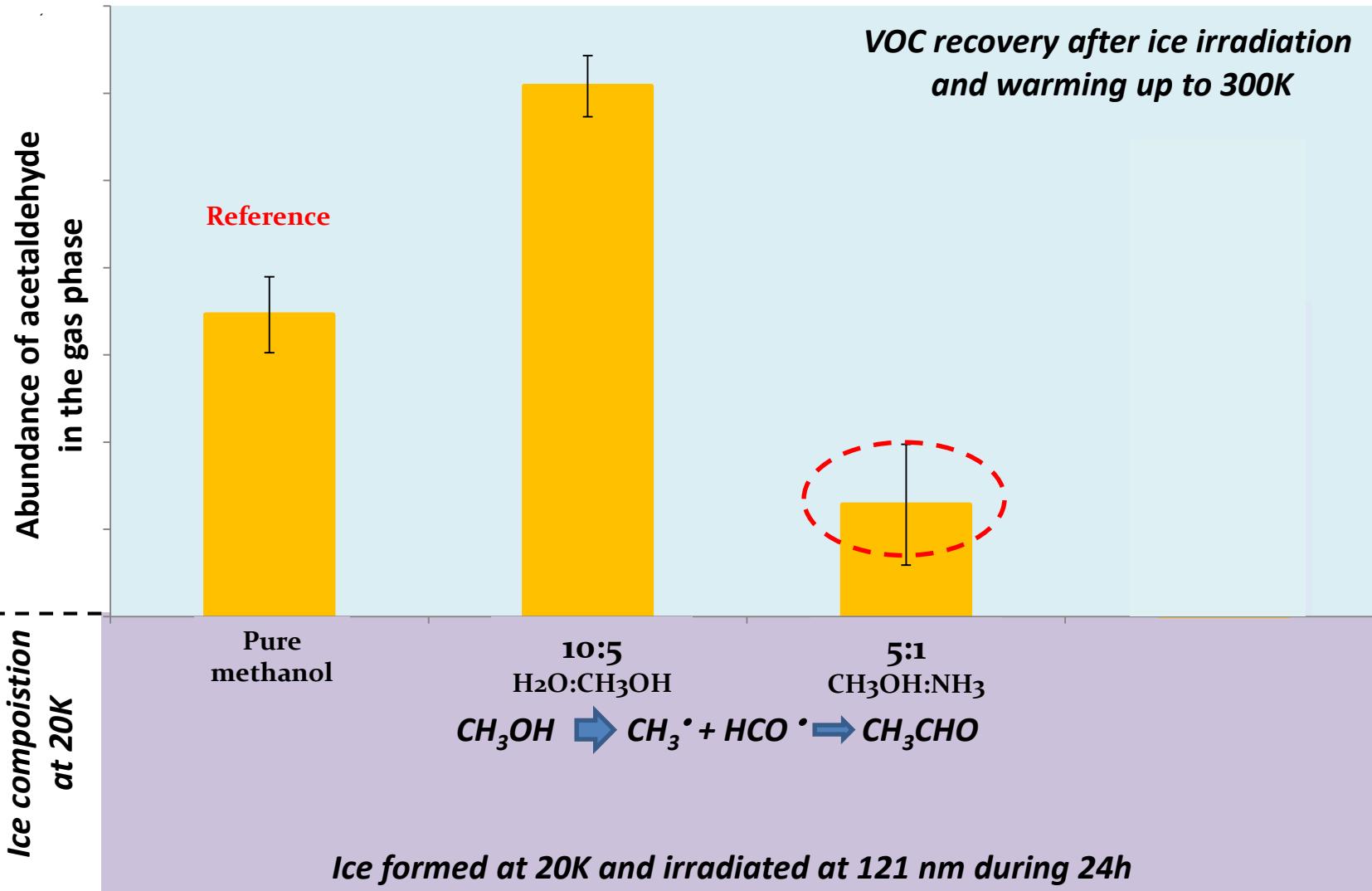
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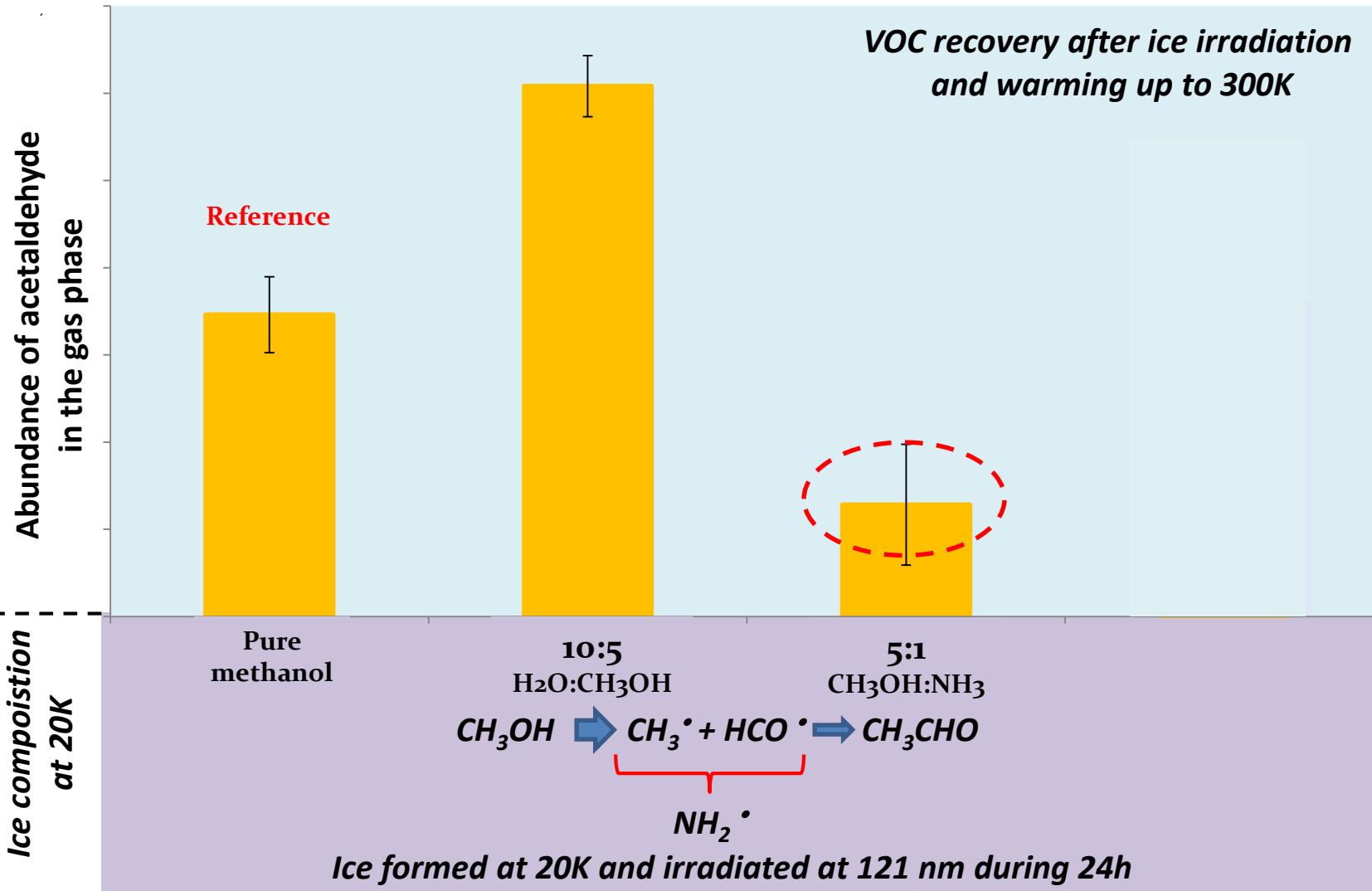
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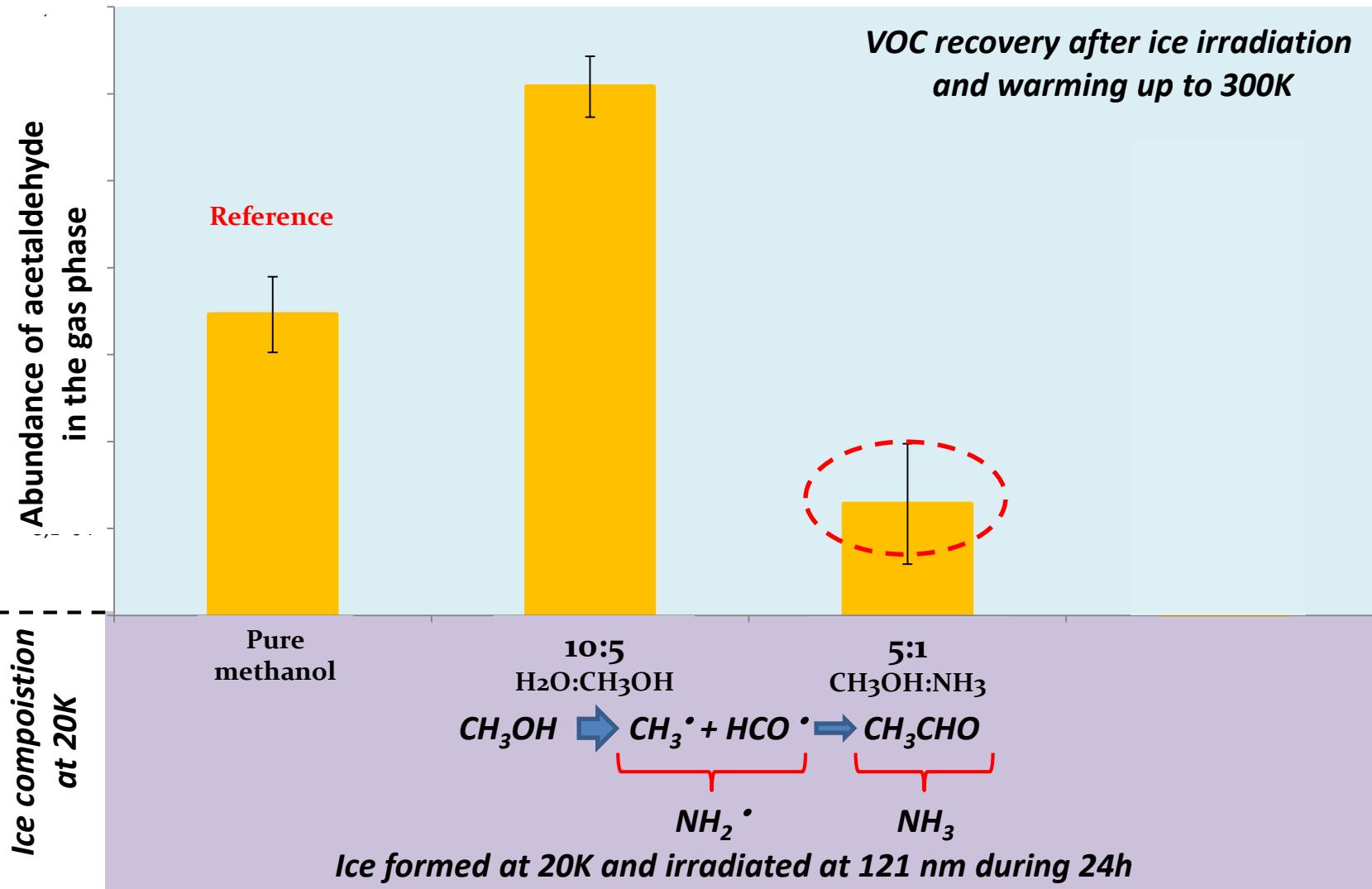
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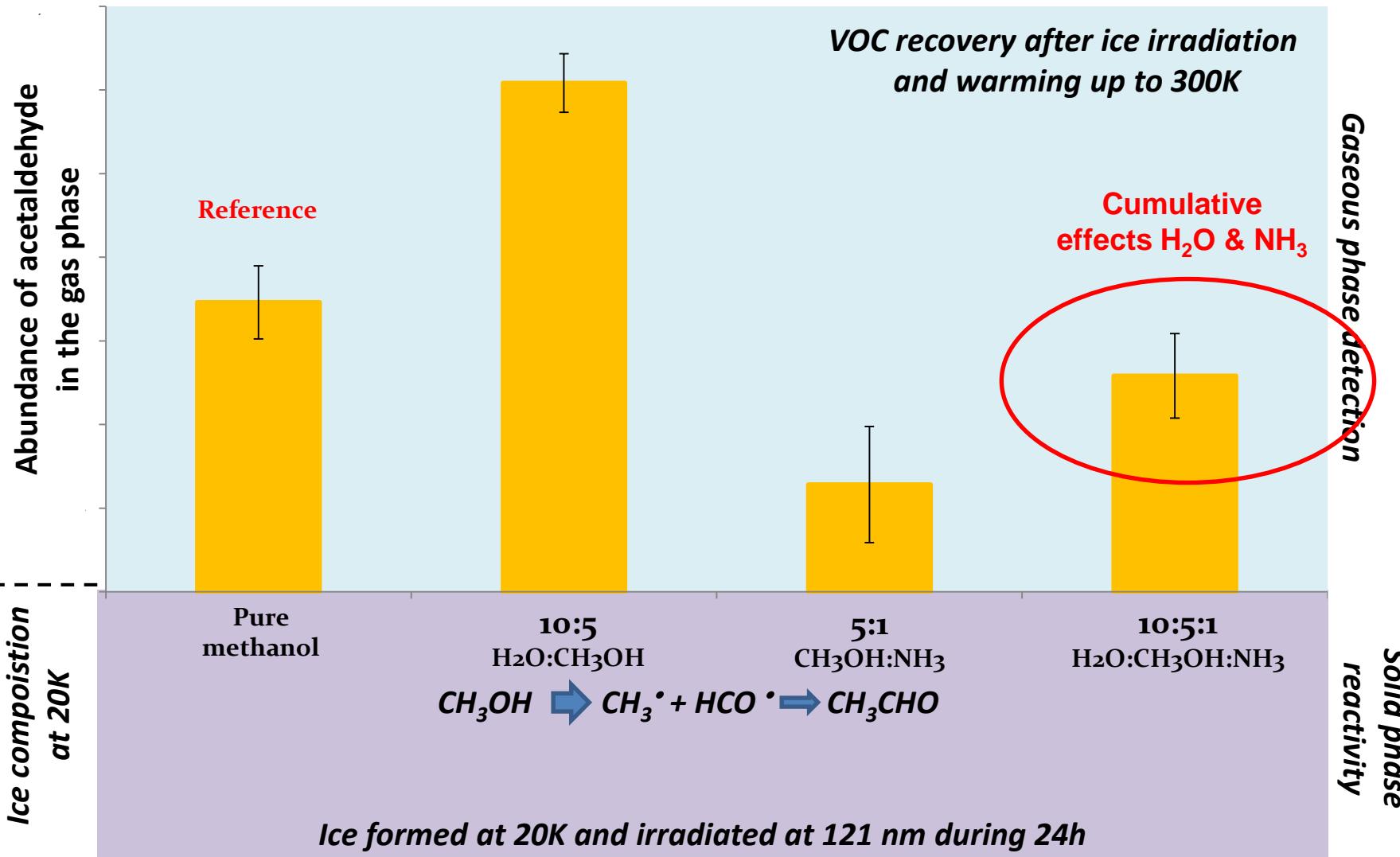
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Complex ices
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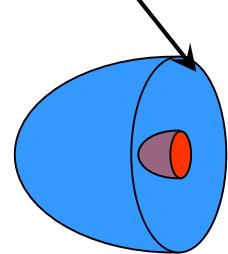
VOC example: Acetaldehyde



Evolution of interstellar icy grains

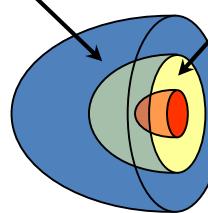
Toward the formation of complex organic matter in interplanetary bodies

Intersteller ices
 H_2O , CH_3OH , NH_3 ,
 CO_2 , CH_2O ...



Projet VAHIIA
Volatile Organic Compounds (VOC)

High energy processes
Heating



Refractory
compounds

High energy processes
Heating



Refractory organic residue

Chemistry in diluted
environment
20 K – 150 K

Radical and Thermal
reactivities in water ice

Formation of small
complex organic molecules

Water matrix
restructuration

and water
desorption

150 K – 185 K

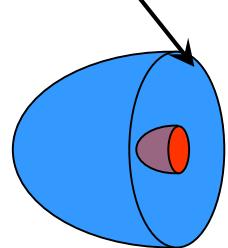
① Main desorption
of VOCs



Evolution of interstellar icy grains

Toward the formation of complex organic matter in interplanetary bodies

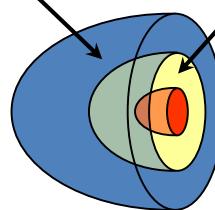
Intersteller ices
 H_2O , CH_3OH , NH_3 ,
 CO_2 , CH_2O ...



Projet VAHIIA

Volatile Organic Compounds (VOC)

High energy processes
Heating



Refractory
compounds

High energy processes
Heating

Projet RAHIIA

Refractory organic residue

ANR RAHIIA_SSOM (2016-2021)

Chemistry in diluted environment
 $20 \text{ K} - 150 \text{ K}$

Radical and Thermal reactivities in water ice

Formation of small complex organic molecules

Water matrix restructuration and water desorption
 $150 \text{ K} - 185 \text{ K}$

- ① Main desorption of VOCs
- ② Some VOCs trapped and react in the water matrix

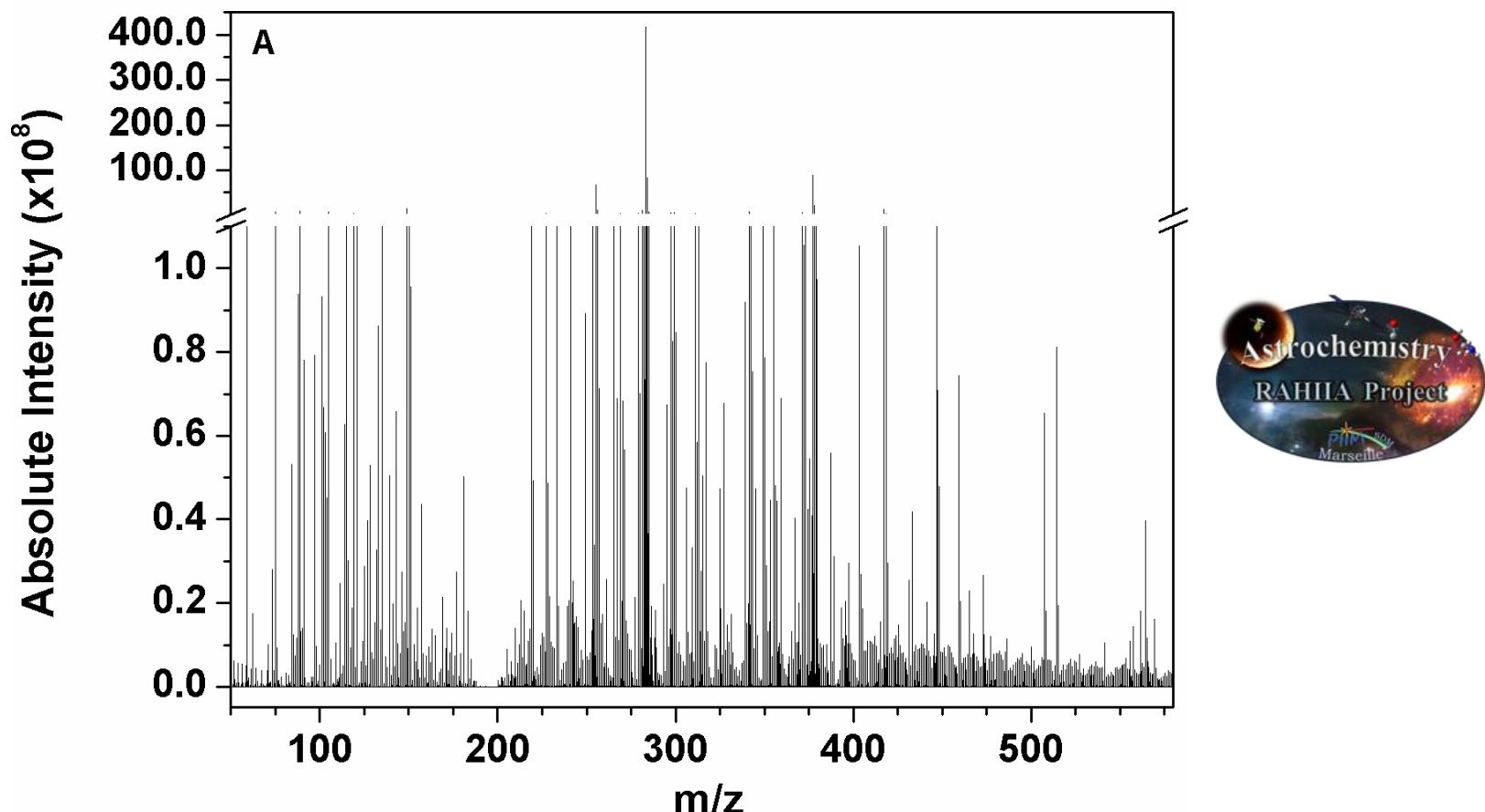
Chemistry in concentrated environment
 $> 185 \text{ K}$

Reactivity in absence of water

Formation of « macromolecules »

VHRMS (orbitrap) analyses of organic residues ($\text{H}_2\text{O}/\text{NH}_3/\text{CH}_3\text{OH} = 3/1/1$)

Collaboration with R. Thissen (IPAG, Grenoble, France) and P. de Marcellus, L. d'Hendecourt (IAS, Paris, France)

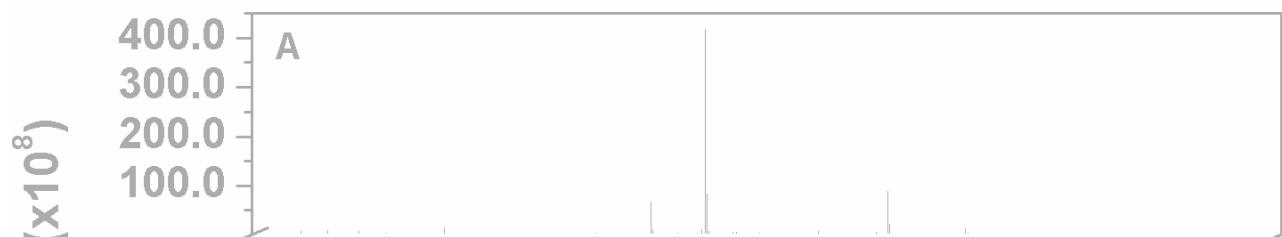


Negative ESI mode = $[\text{M}-\text{H}]^-$ analysis

Molecules with proton donor chemical functions (e.g. carboxylic acid -COOH)

VHRMS (orbitrap) analyses of organic residues ($\text{H}_2\text{O}/\text{NH}_3/\text{CH}_3\text{OH} = 3/1/1$)

Collaboration with R. Thissen (IPAG, Grenoble, France) and P. de Marcellus, L. d'Hendecourt (IAS, Paris, France)



From an ice uniquely formed of $\text{H}_2\text{O}:\text{CH}_3\text{OH}:\text{NH}_3$

VUV + warming

=

**Thousand of molecules with masses
up to 4,000 Da -> macromolecules**

=

Complex and rich chemistry

=

Important molecular diversity

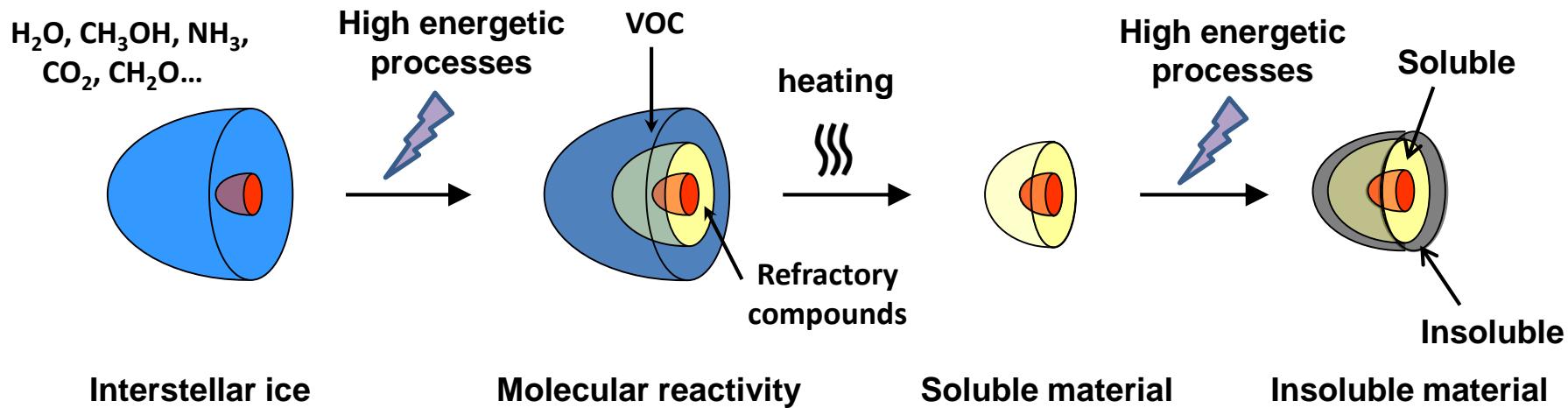
m/z

Negative ESI mode = $[\text{M}-\text{H}]^-$ analysis

Molecules with proton donor chemical functions (e.g. carboxylic acid -COOH)

Evolution of interstellar icy grains toward the formation of complex organic molecules in interplanetary objects

Primary processing



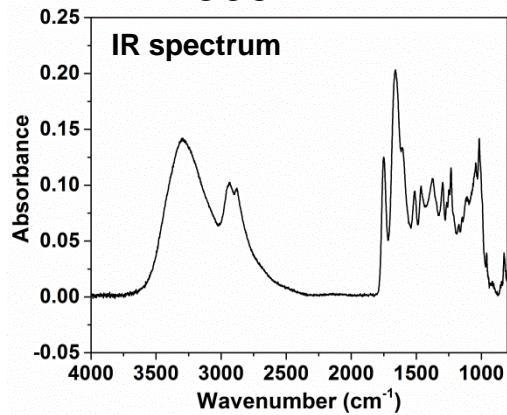
Evolution of interstellar ices: Formation of complex organic molecules

Collaboration with P. de Marcellus, L. d'Hendecourt and R. Brunetto (IAS, Paris, France)

A scenario from extraterrestrial ices to soluble and insoluble materials

Residue (soluble)

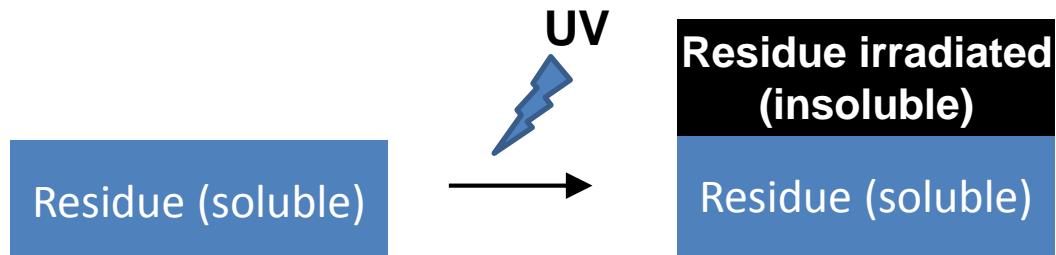
300 K



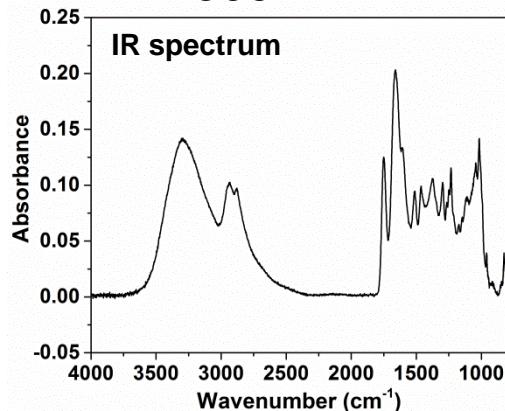
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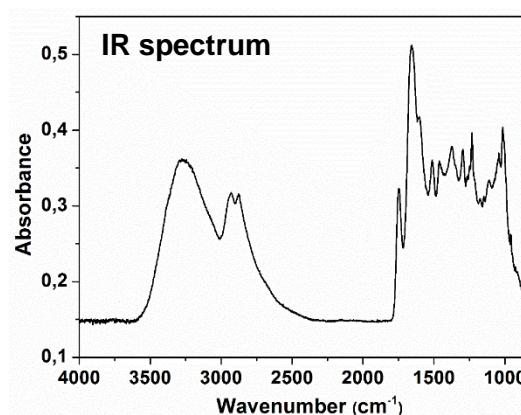
A scenario from extraterrestrial ices to soluble and insoluble materials



300 K



300 K



Irradiated zone

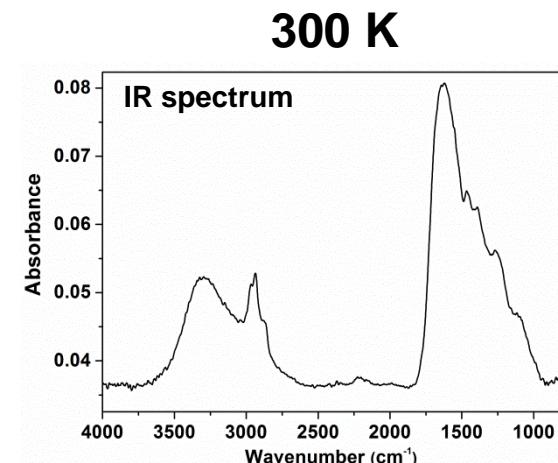
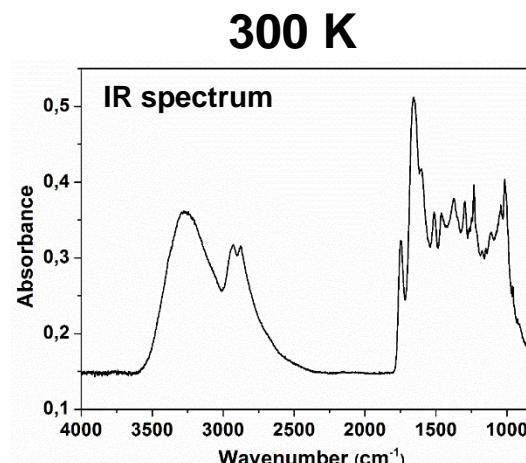
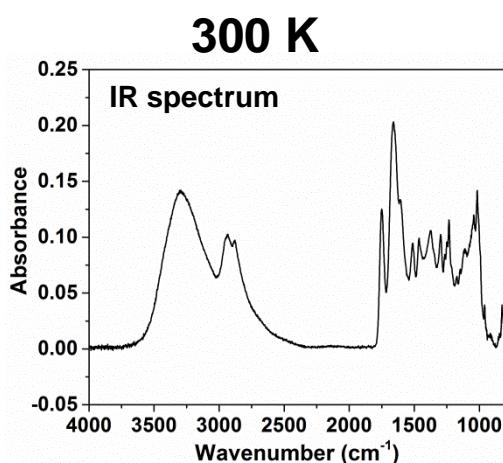
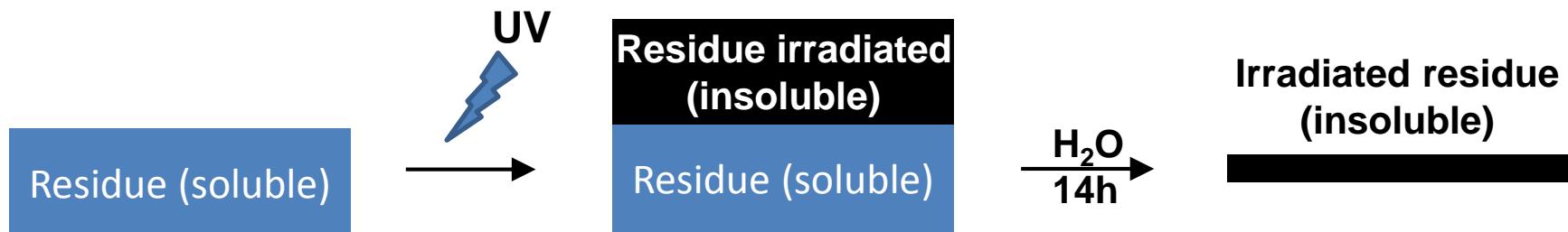


unirradiated zone

Evolution of interstellar ices: Formation of complex organic molecules

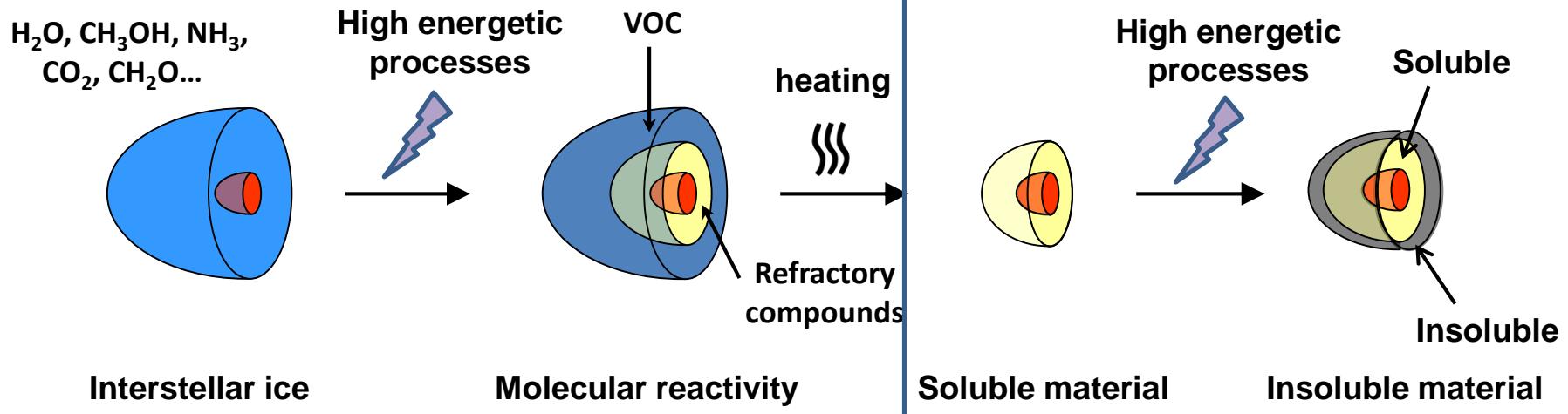
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A scenario from extraterrestrial ices to soluble and insoluble materials



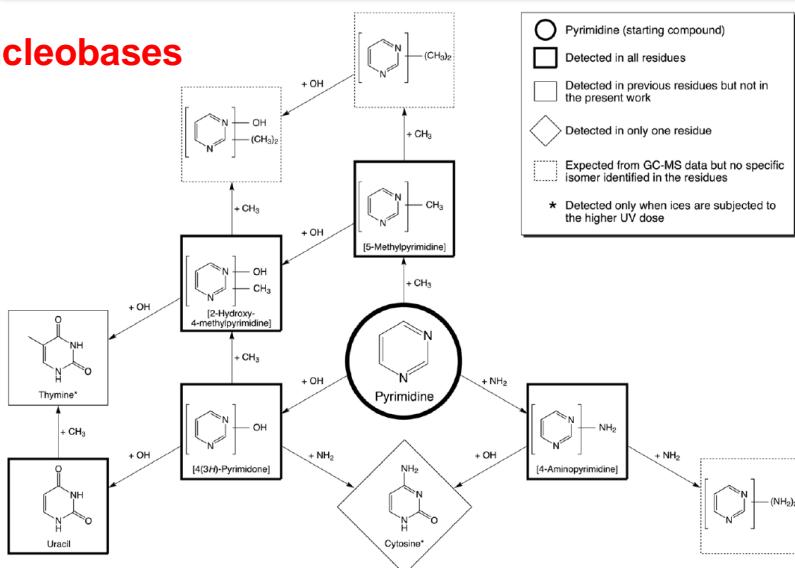
Evolution of interstellar icy grains toward the formation of complex organic molecules in interplanetary objects

Primary processing

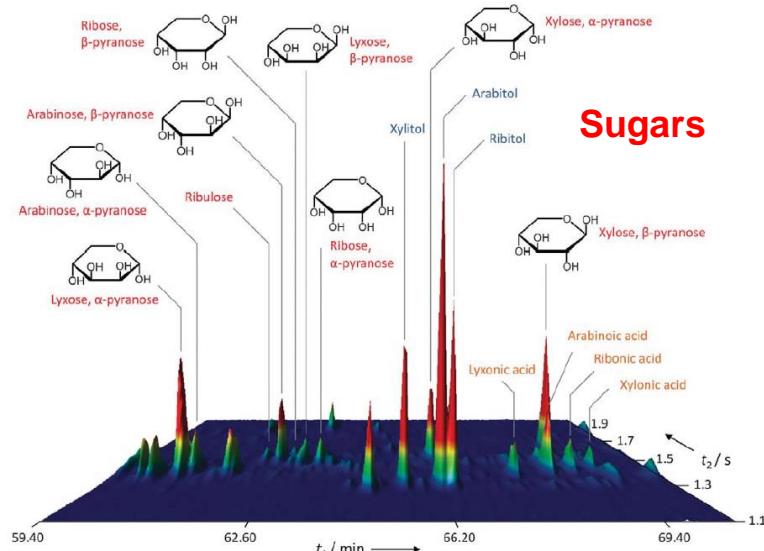


Evolution of interstellar icy grains toward the formation of molecules for prebiotic chemistry

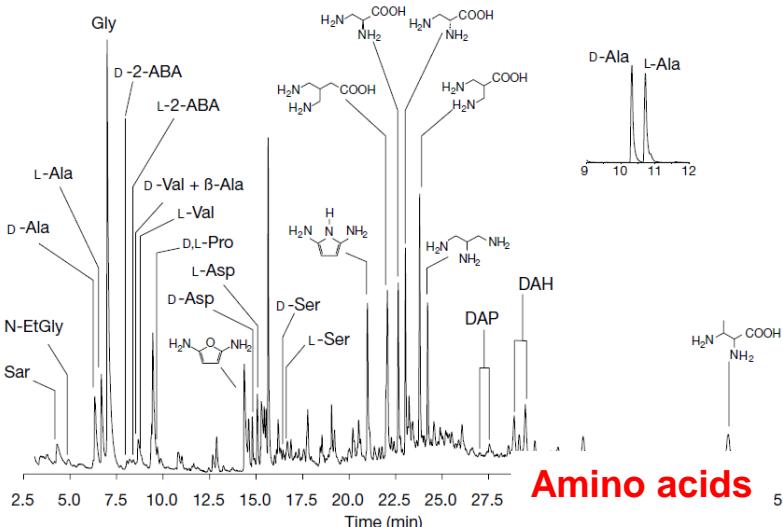
Nucleobases



Nuevo et al., 2014, *ApJ*, 793, 125



Meinert et al., 2016, *Science*, 352, 208



5

Munoz-Caro et al., 2002, *Nature*, 416, 403

Kaiser et al., 2013, *ApJ*, 765, 111

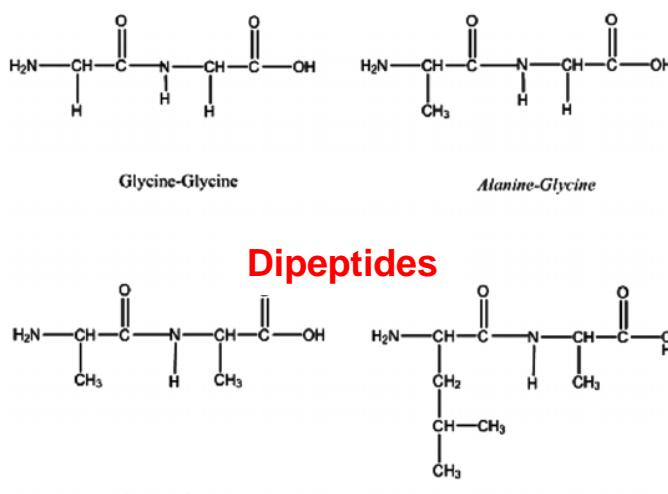
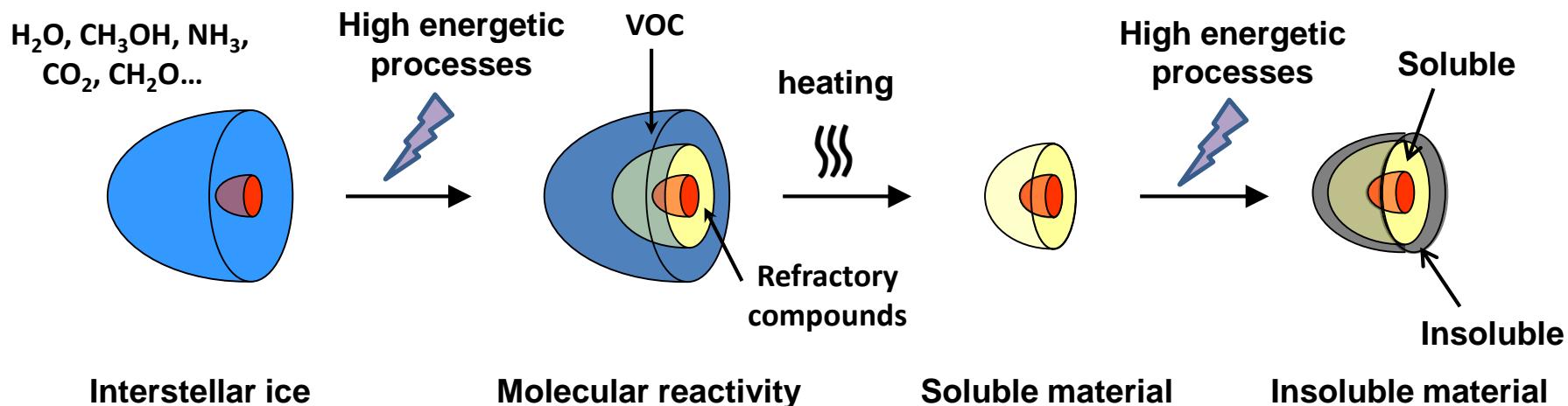


Figure 3. Dipeptides characterized via microchip capillary electrophoresis.

Evolution of interstellar icy grains toward the formation of complex organic molecules in interplanetary objects

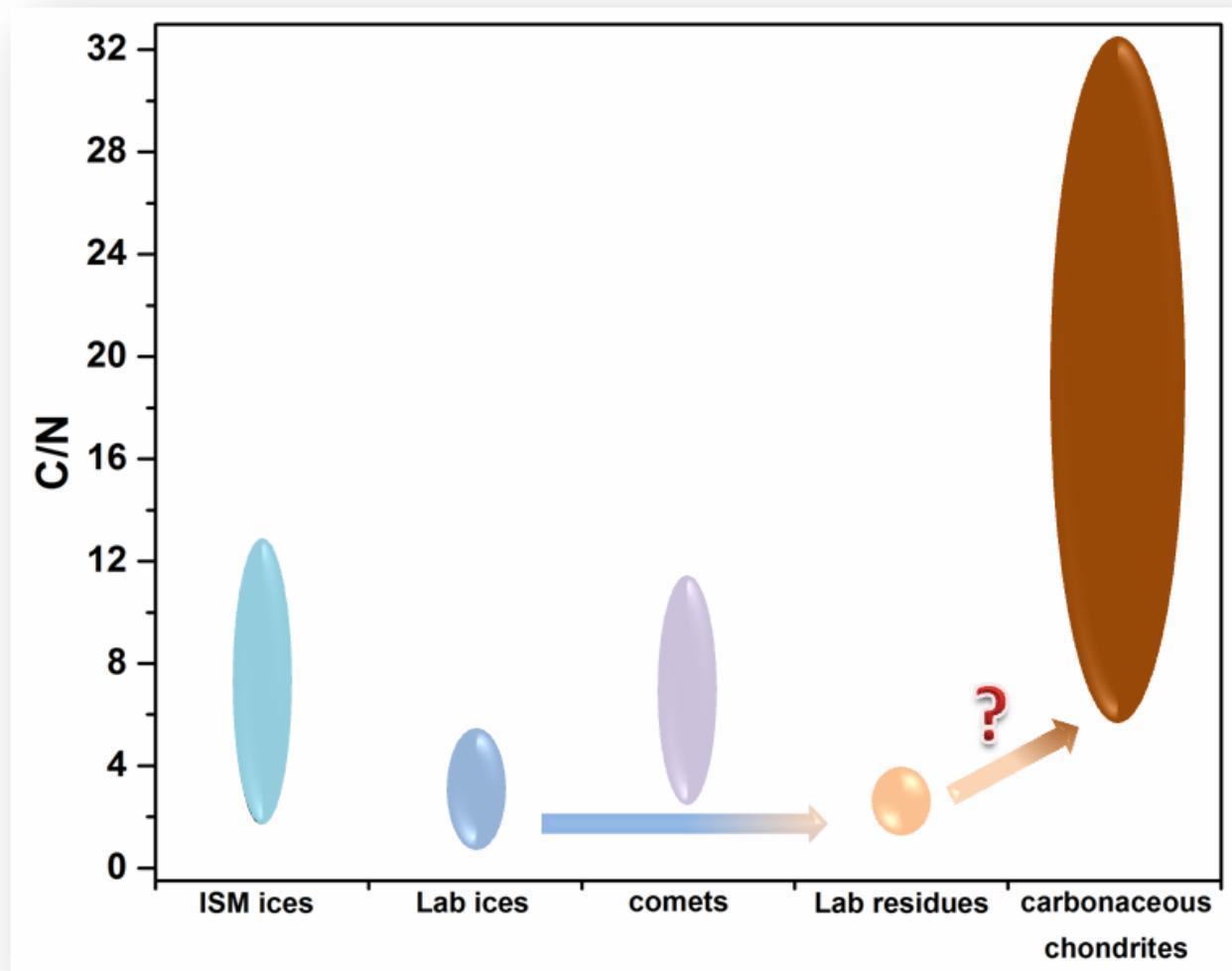
Primary processing



Formation of interplanetary bodies

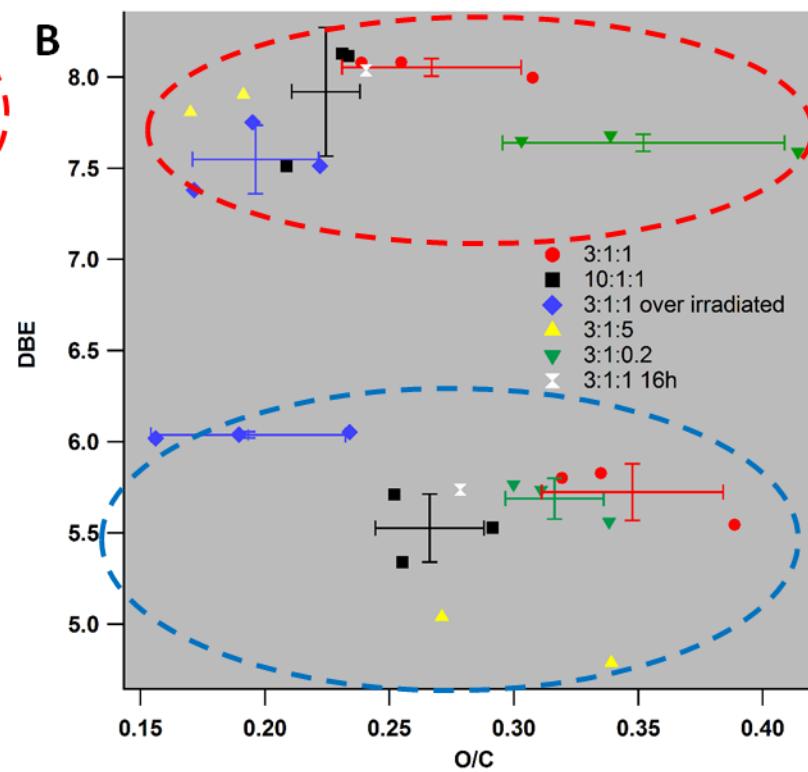
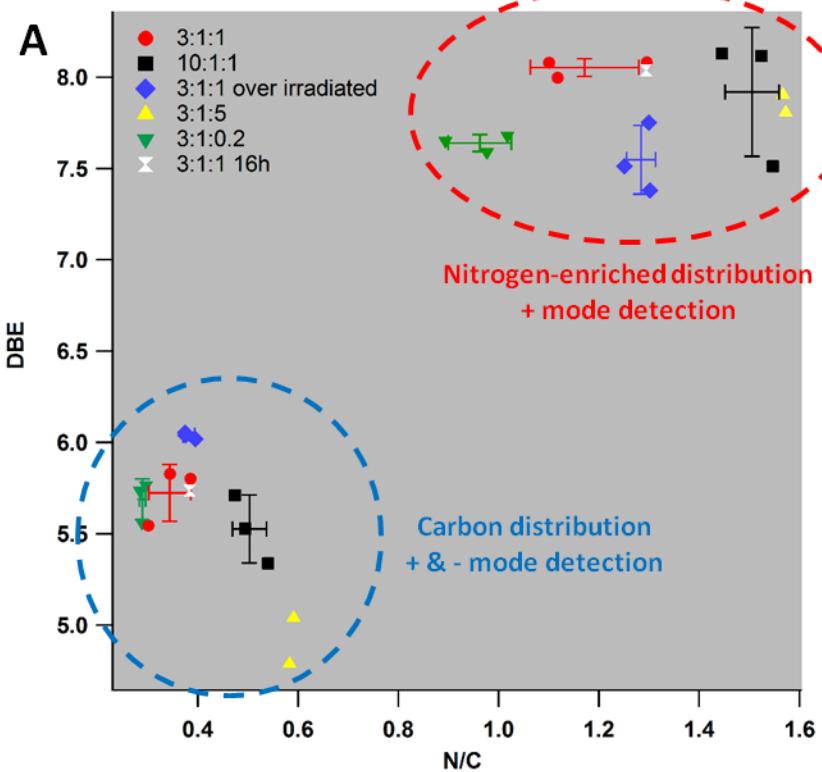
Secondary processing

Evolution of interstellar icy grains toward the formation of complex organic molecules in interplanetary objects

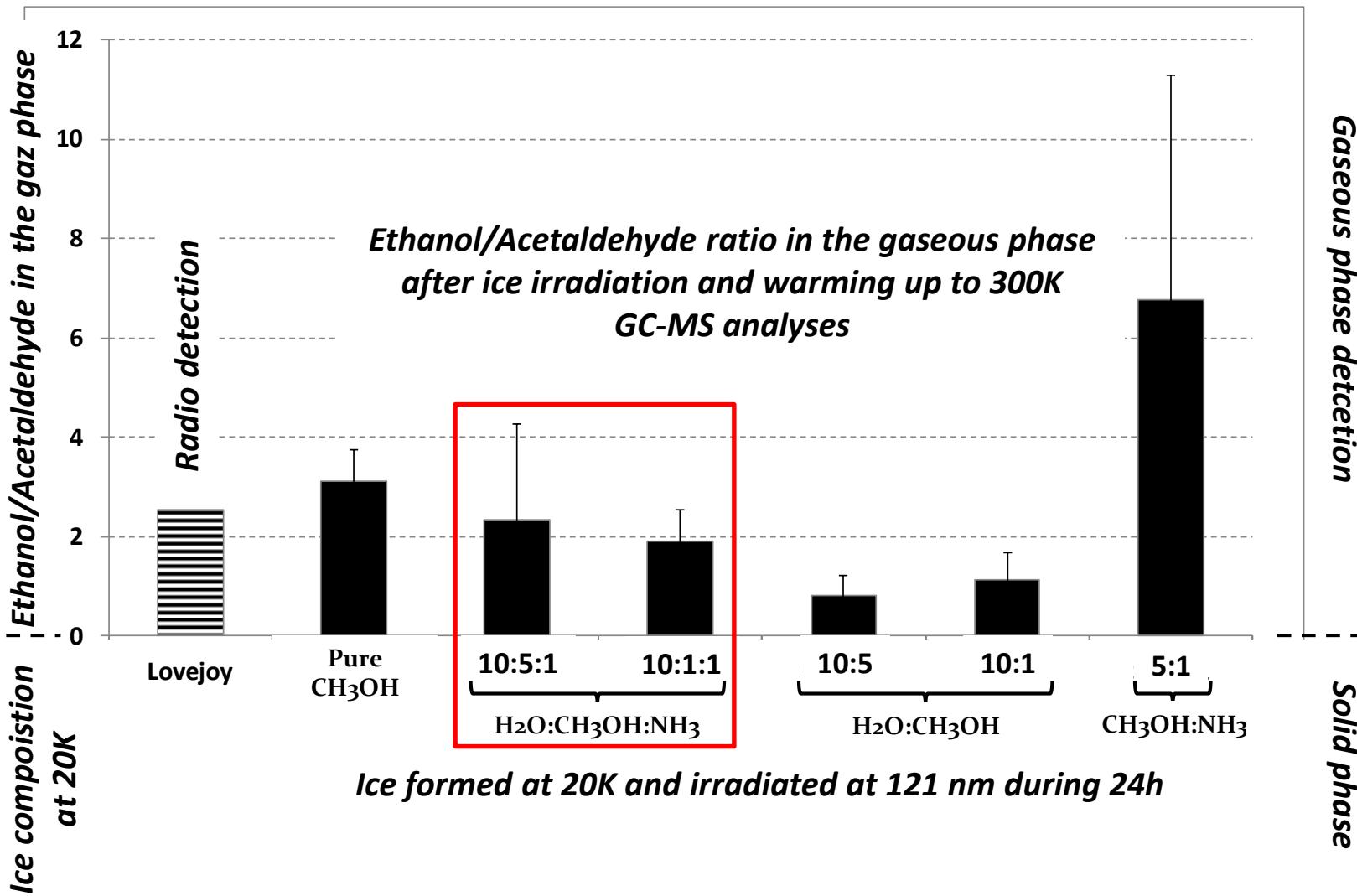


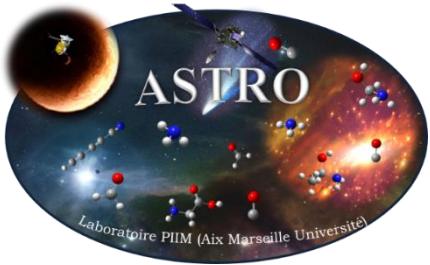
Complex ices

Variability of residues



Complex ices Comparison to observations





Team Astrochemistry, Spectrometry, Theory, Reactivity and Origin (ASTRO)
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Fabien Borget



Ninette Abou Mrad
(Post-Doc)



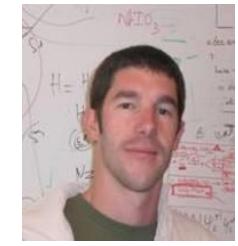
Thierry Chiavassa



Isabelle Couturier



Grégoire Danger



Fabrice Duvernay



Aurélien Fresneau
(PhD student)



Teddy Butsher
(PhD student)



Nathalie Pietri



Patrice Theulé

Collaborations on this project

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R. Thissen, V. Vuitton, F. Orthous-Daunay, IPAG, Grenoble

Funding: CNRS (PCMI/EPOV/**PNP**), **CNES (exobiologie)**,
région PACA, Fondation AMU and **ANR (Peptisystem,**
VAHIIA & RAHIIA_SSOM)