The SRG/eROSITA All-Sky Survey

View of the Fornax galaxy cluster

T.H. Reiprich¹, A. Veronica¹, F. Pacaud¹, P. Stöcker¹, V. Nazaretyan¹, A. Srivastava¹, A. Pandya¹, J. Sanders², M. Yeung², A. Chaturvedi³, M. Hilker⁴, B. Seidel⁵, K. Dolag⁵, E. Hernández-Martínez⁵, J. Comparat², V. Ghirardini², M. Kluge², A. Liu² et al., to be subm.



Thomas Reiprich Argelander Institute for Astronomy University of Bonn <u>http://dark-energy.net</u>













Motivation



- Fornax cluster (M<10¹⁴ M_☉, central galaxy NGC 1399) is nearby and bright. → Resolve details.
- Dynamically active system; e.g., from infall/merger of NGC 1404 shock front is predicted at outskirts (Sheardown+18).
 → Search for this and other features with eROSITA's unlimited fieldof-view.
- Fornax A group is bound to the main cluster (Drinkwater+01) but no gaseous emission bridge has been found in the past.
 → Revisit this, taking advantage of eROSITA's excellent low surface brightness sensitivity in the soft band.

SRG/eROSITA X-ray All-Sky Survey

> eRASS1 Western Galactic half

0.3-0.63 keV 0.63-1.04 keV 1.04-2.3 keV

~52" Healpix



Credit: MPE, J. Sanders for the eROSITA consortium

Fornax is very close (z = 0.005) and huge $(D_{100} \sim 4.4 \text{ deg})!$



R₂₀₀s on 2MASS galaxy distribution

Fornax is very close (z = 0.005) and huge $(D_{100} \sim 4.4 \text{ deg})!$



• Several slides removed here, sorry. ☺ But the paper will soon be submitted, and appear on arXiv. ☺

Summary

- Essentially all the extended structures seen in the wavelet-filtered image are due to real emission variations (as expected).
- Much (but not all) of it is due to structure in the outskirts of the Fornax cluster.
- No obvious large scale bow shock, as predicted from NGC 1404 merger scenario (Sheardown+18), is seen (TBC); instead, emission "fingers" beyond R_{500} that fade out into enhanced emission regions well beyond R_{100} are discovered.
- Galaxy and GC populations roughly trace the outer X-ray emission regions.
- Indications for an apparent bridge of emission connecting the Fornax A group with the Fornax cluster are found.
- Overall structure, including bridge to Fornax A, is very similar to that of the Fornax digital twin in constrained local volume simulation (SLOW, Dolag+23).