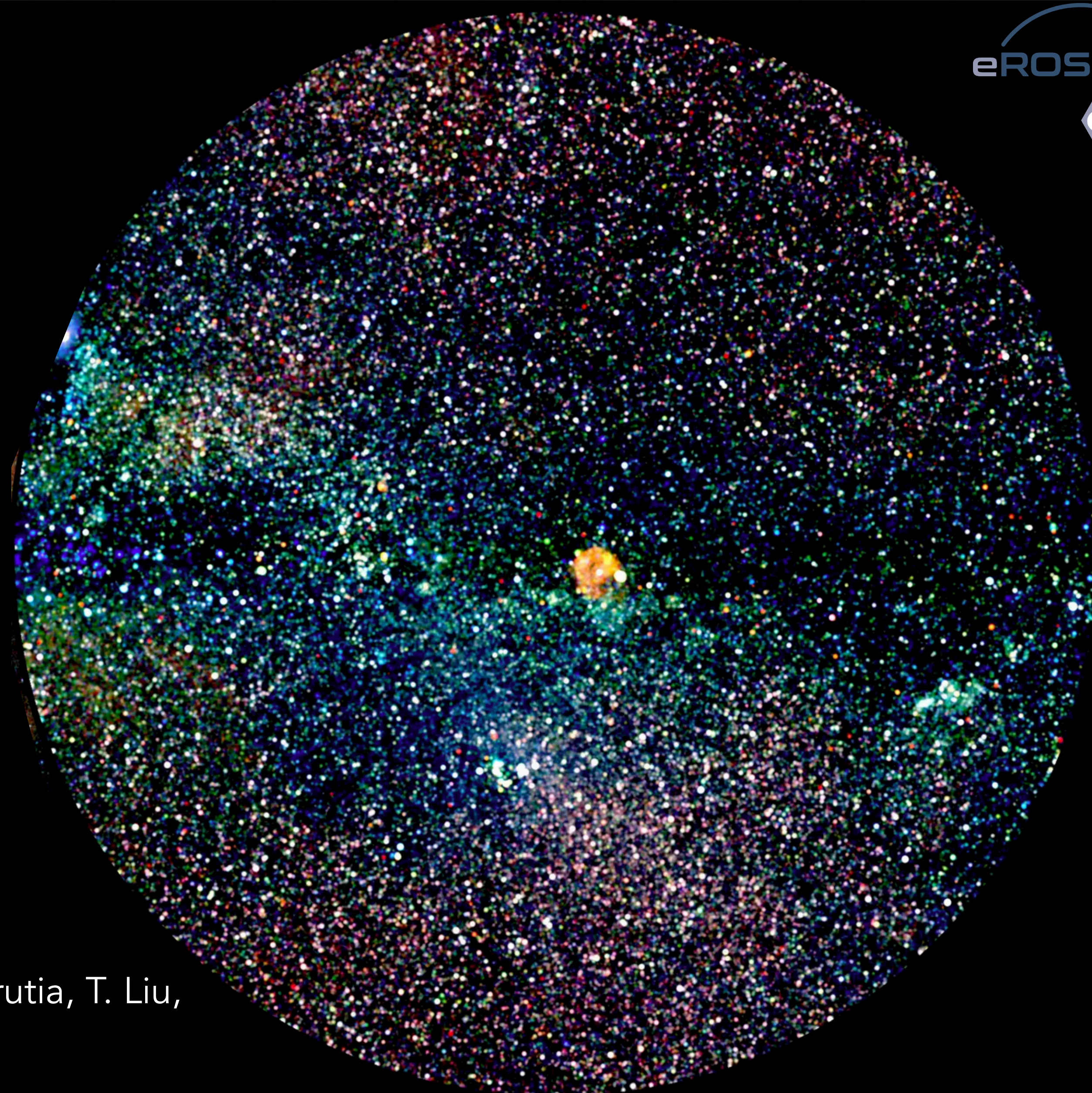


# The first eROSITA All-Sky Survey: AGN content



**Mara Salvato (MPE)**

On behalf of the eROSITA-DE/eroAGN team:

J. Wolf, J. Buchner, H. Starck, T. Dwelly, , M. Brusa, R. Shirley  
A. Merloni, K. Nandra, S. Waddell, G. Lamer, W. Roster, T. Urrutia, T. Liu,  
and many, many more





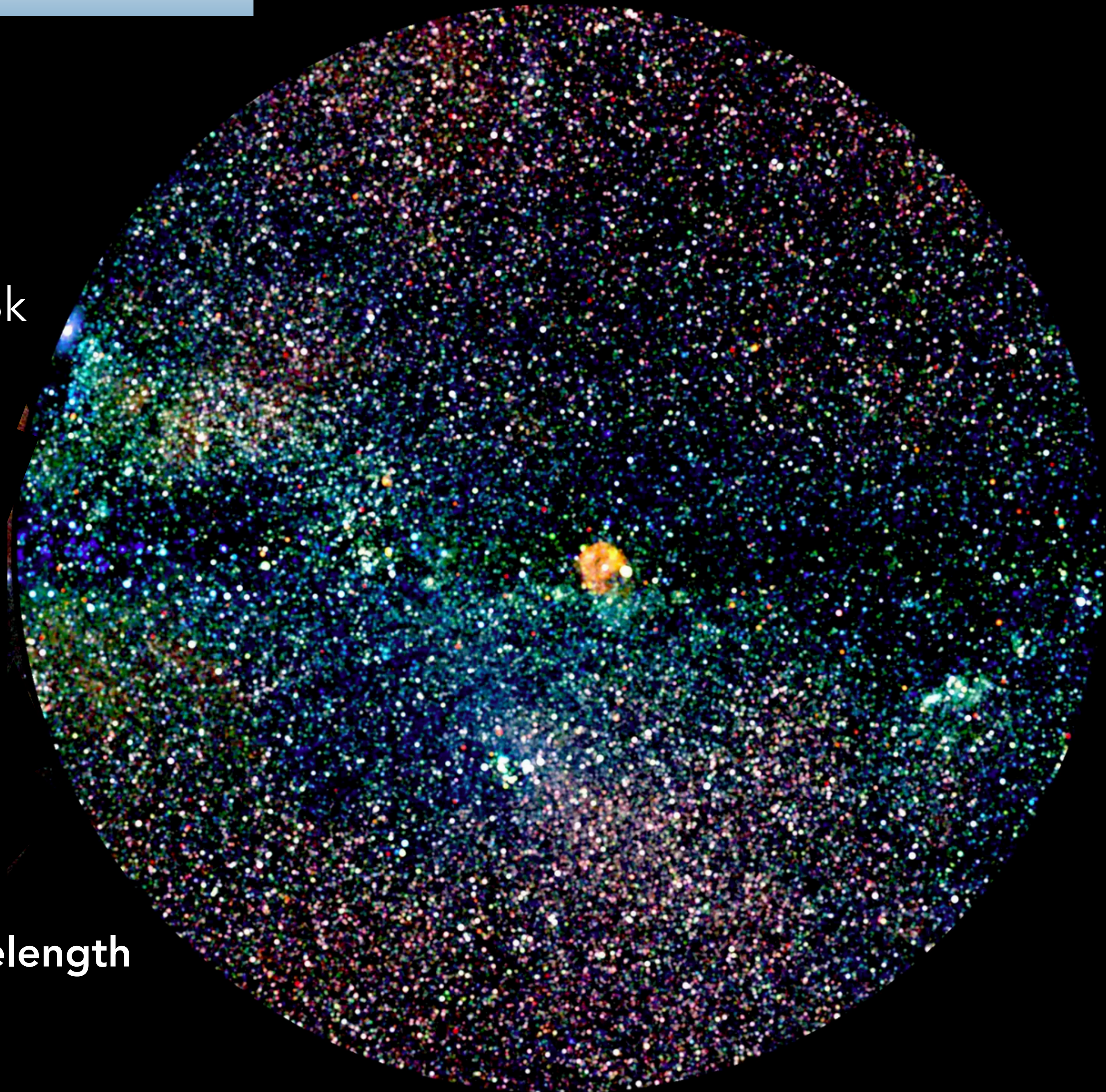
# Talk outlook

From Merloni et al 2024:

1. Soft band 0.2-2.3 keV, Point sources: 903k
2. Hard band 2.3-5 keV, Point Sources: 5.5k  
(see S. Waddell talk)

## In this talk:

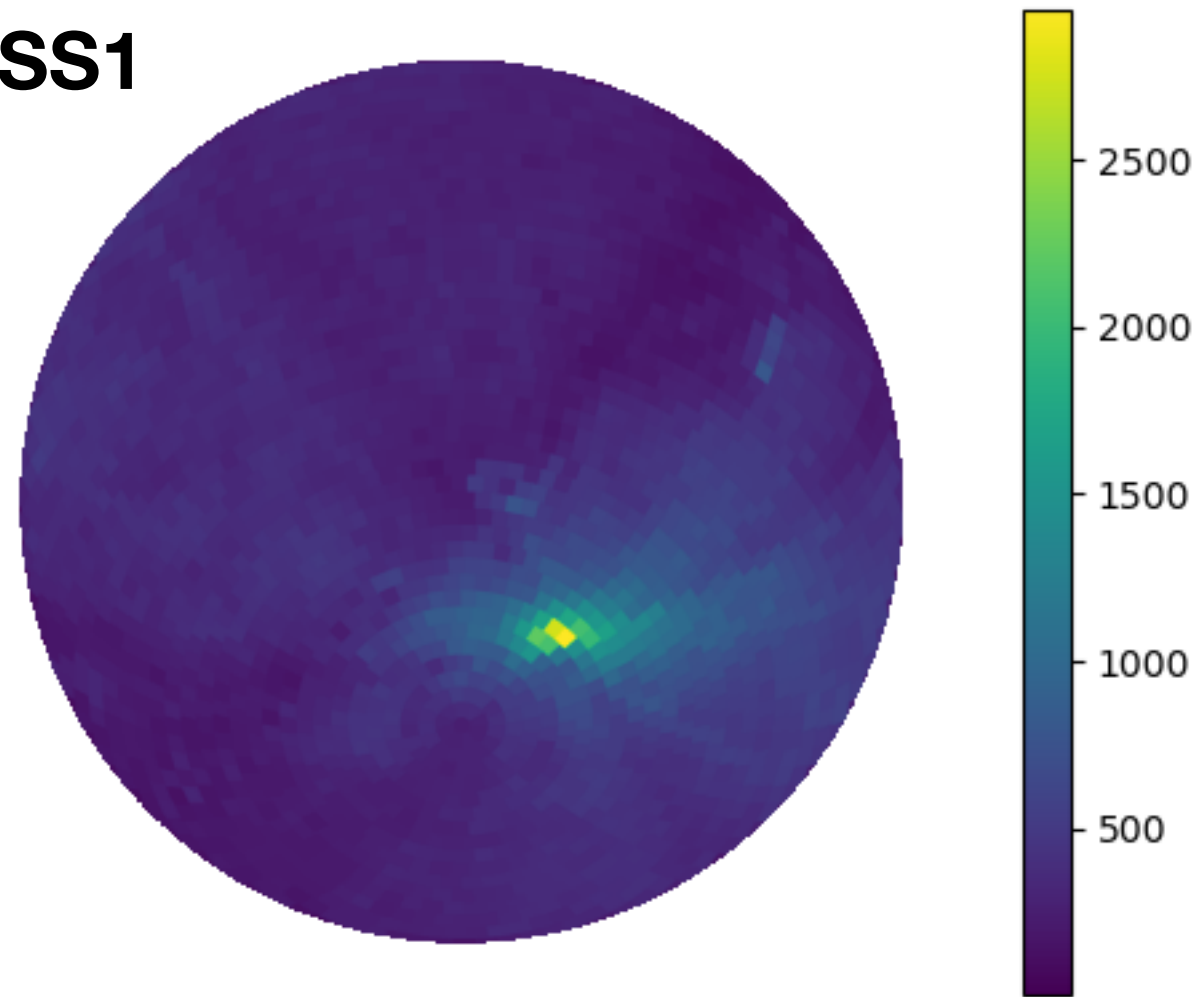
- Short overview on associations
- classification of sources
- comparison with AGN selected at other wavelength
- remarks





# Counterparts identification depends on ancillary data

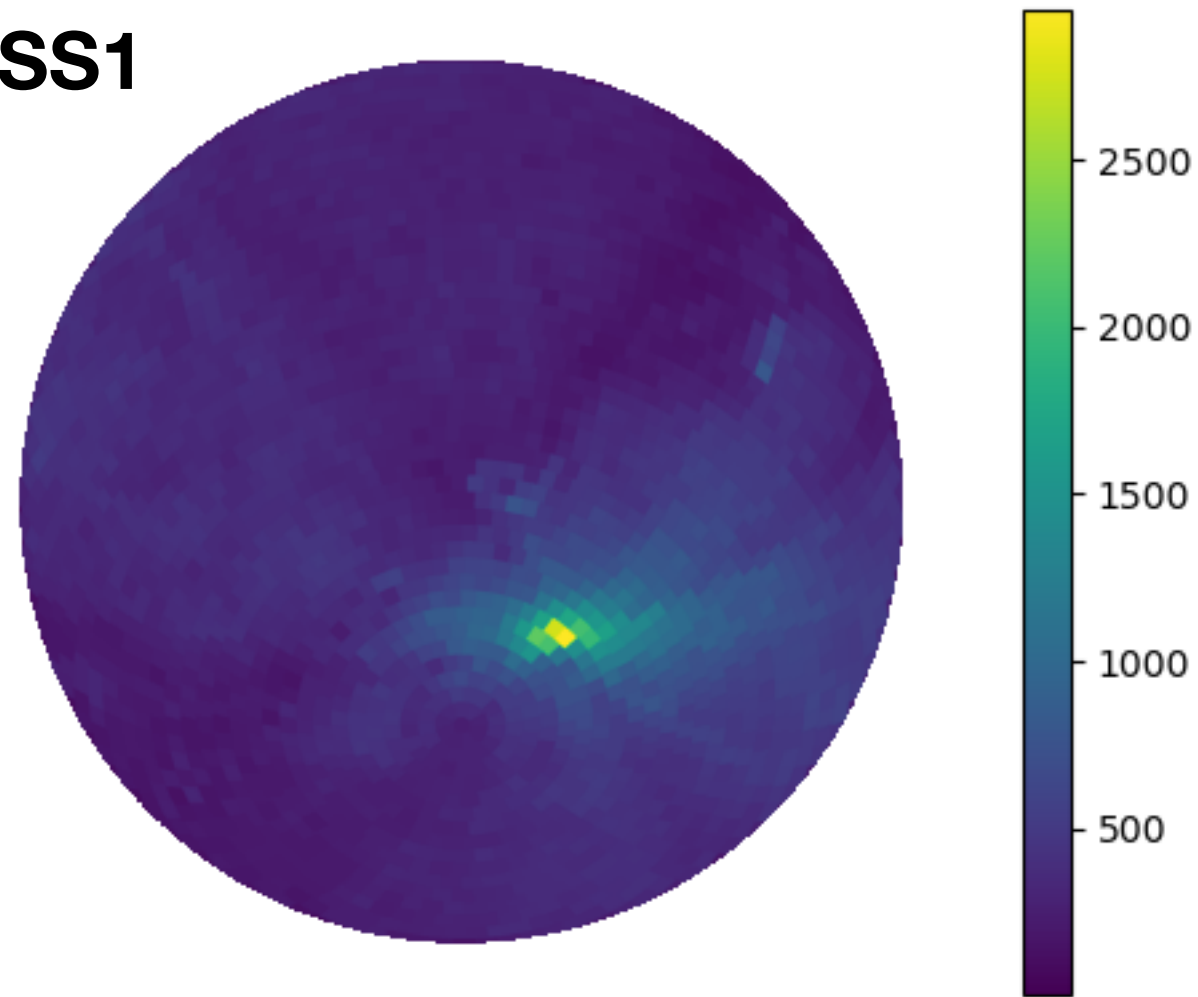
eRASS1



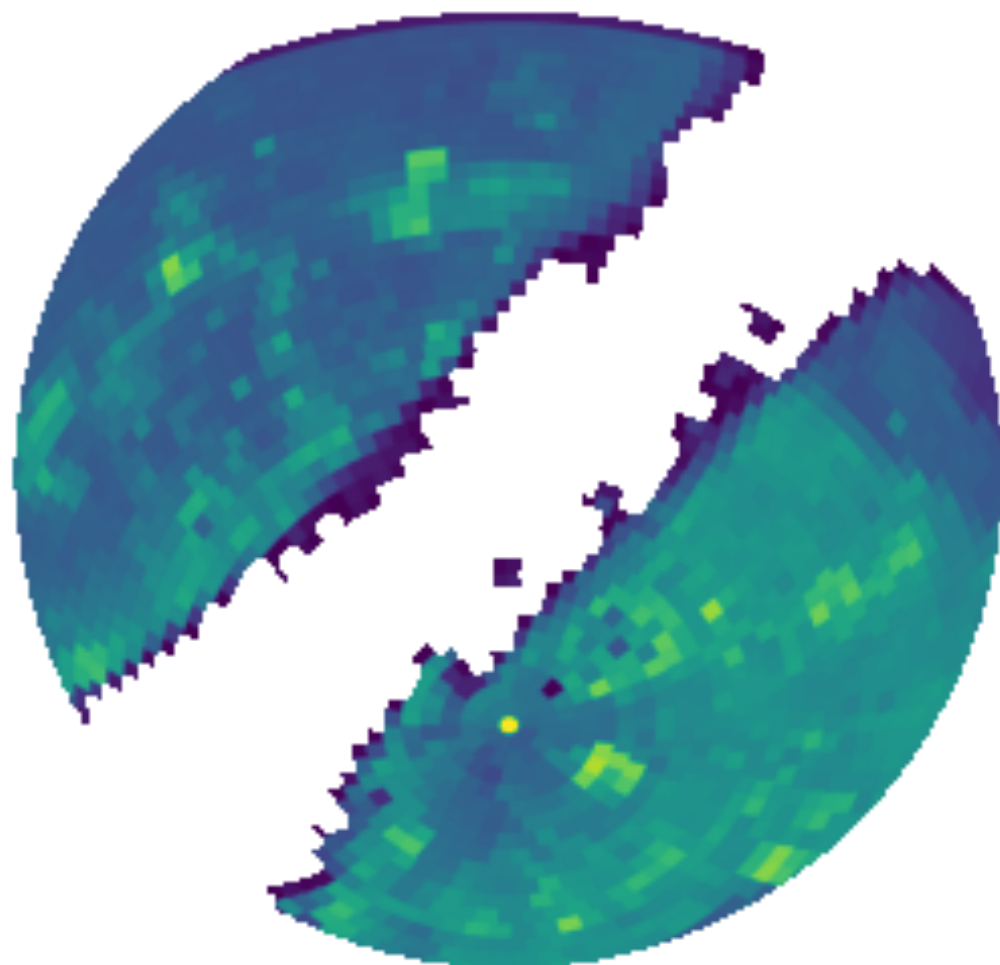
Survey	Depth (AB)	Bands	Area	Population
LS10	~24	griz(W1-4)	14k deg <sup>2</sup>	all, including clusters
Gaia	20	Gr,Gb,G	all-sky	stars, compact objects,qso
CW2020	20.4, 20.8	W1,W2	all-sky	cold stars, QSO, AGN

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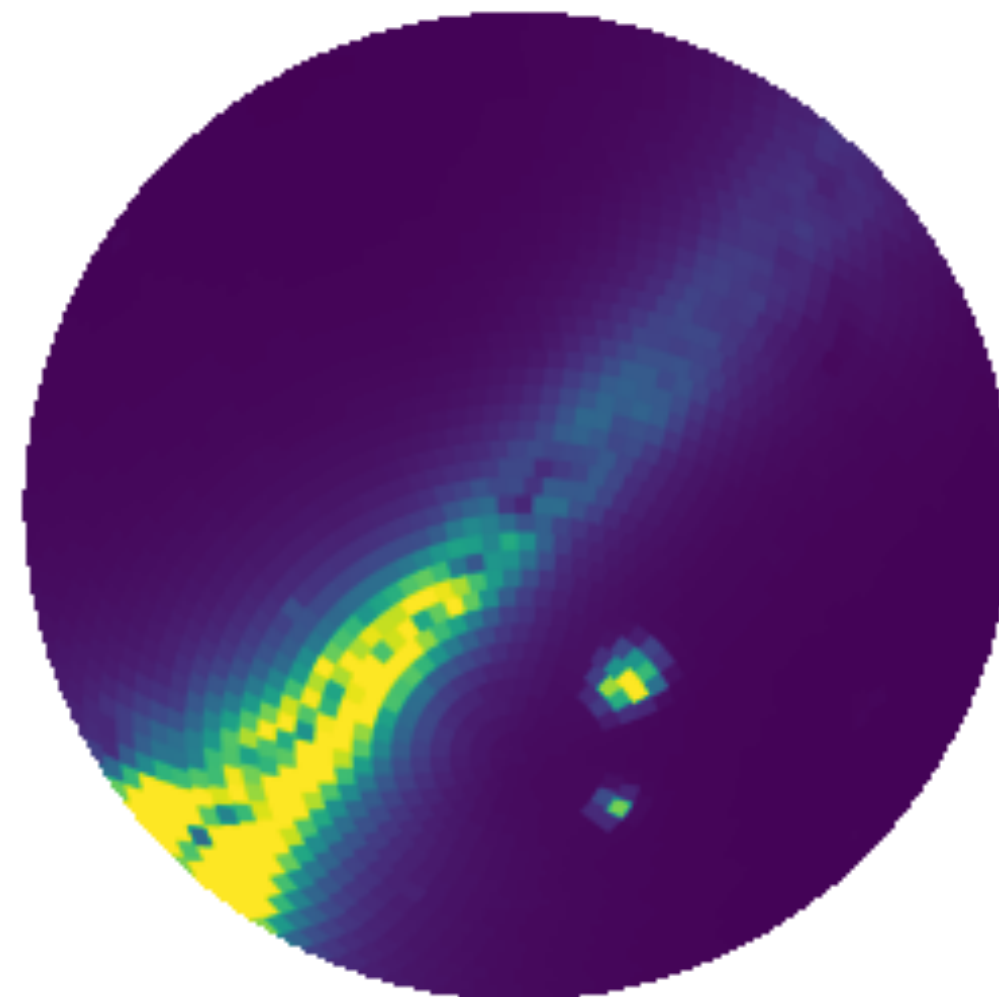
eRASS1



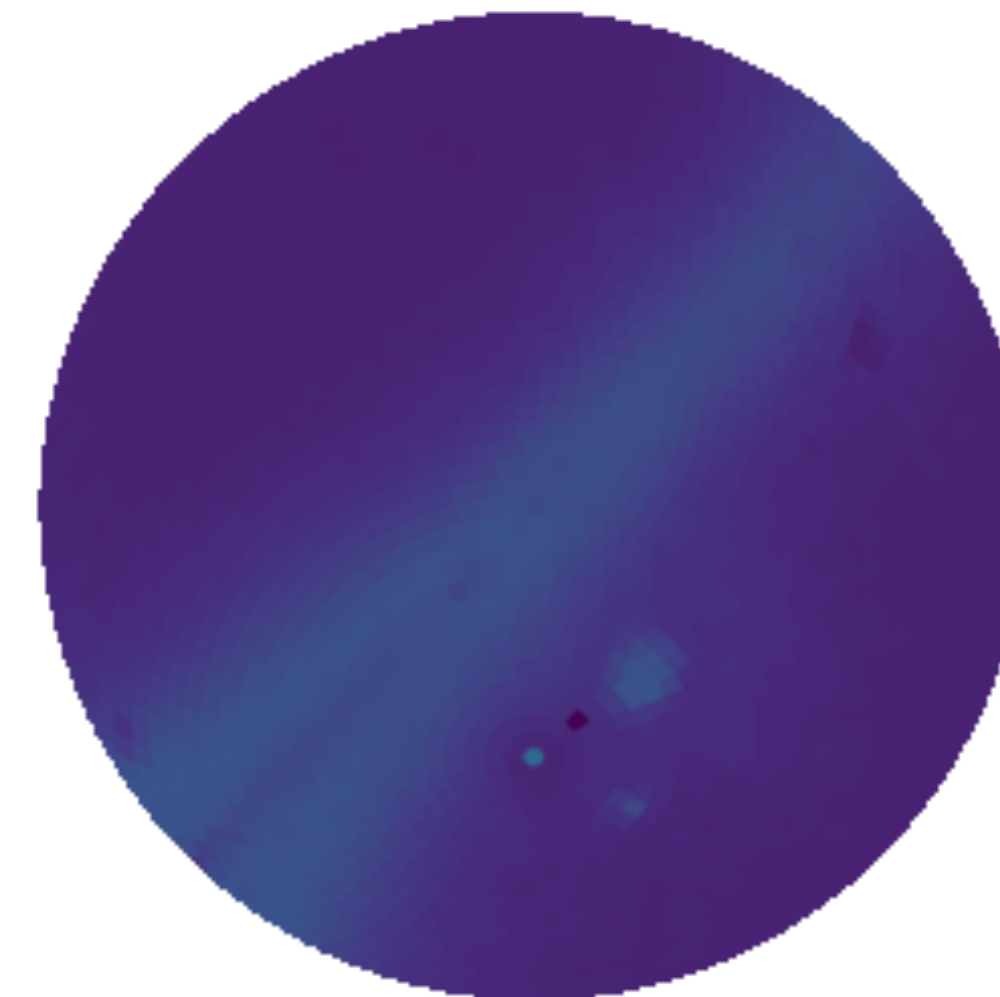
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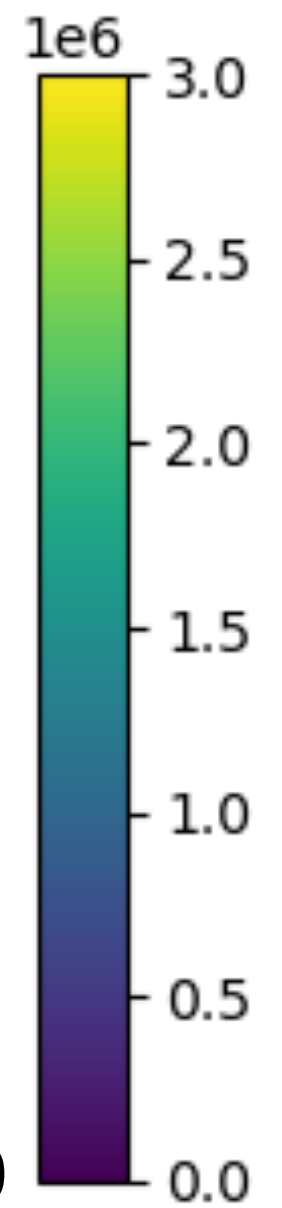
LS10



Gaia DR3



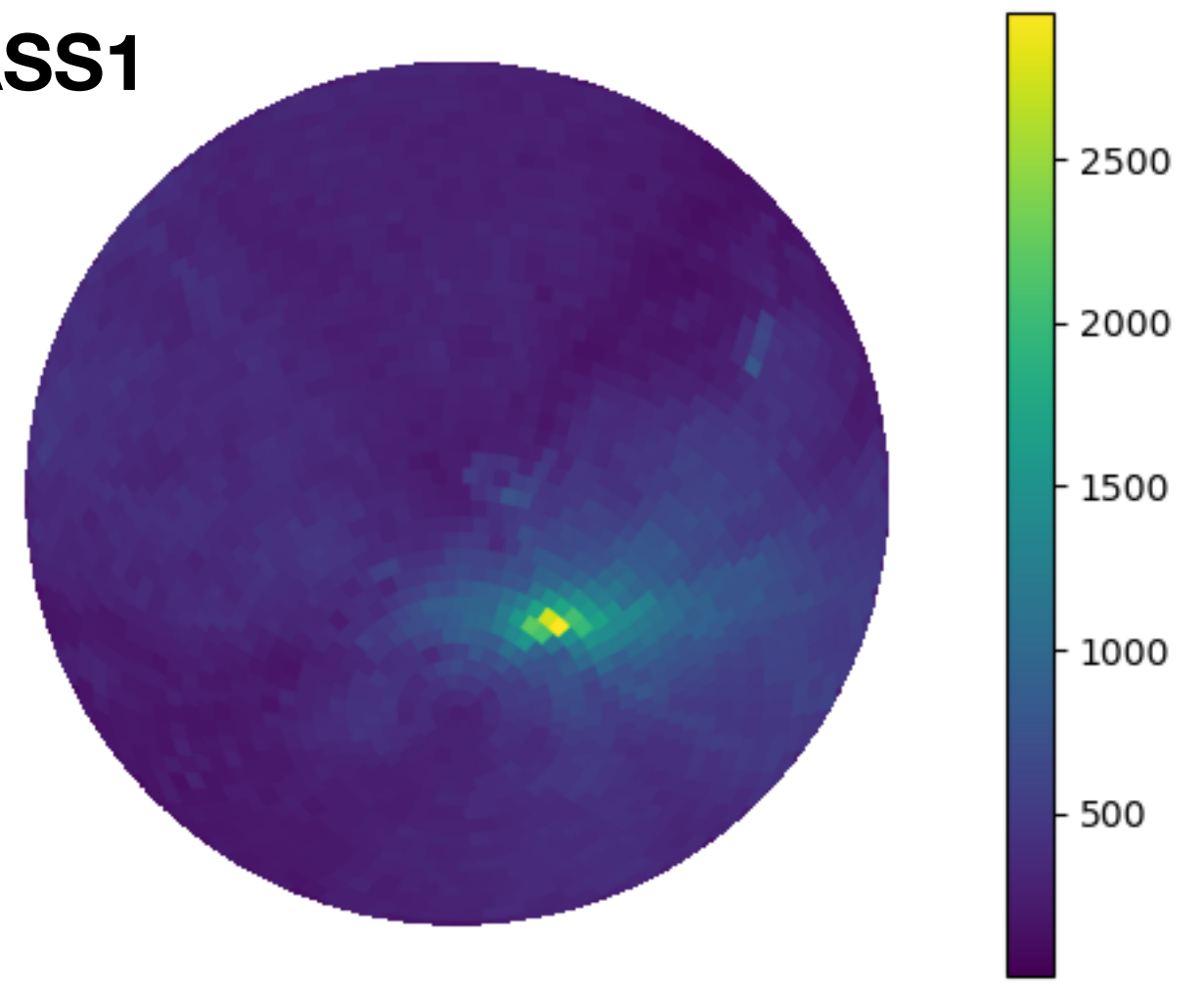
CW2020





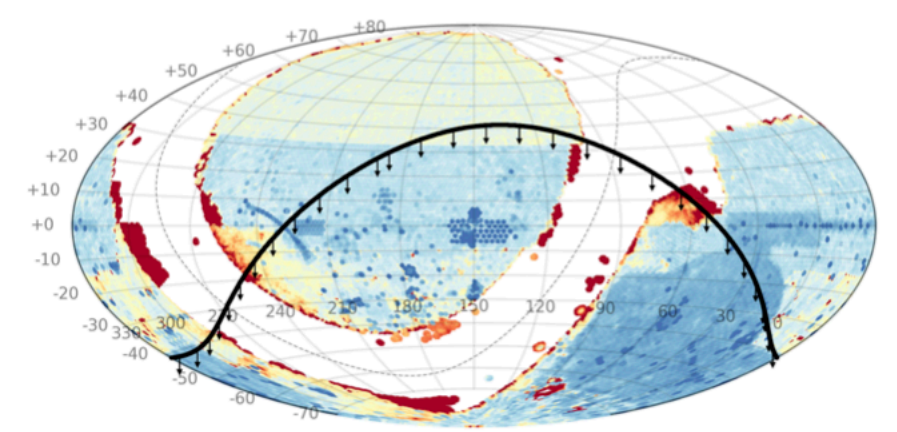
# Counterparts identification depends on ancillary data

eRASS1

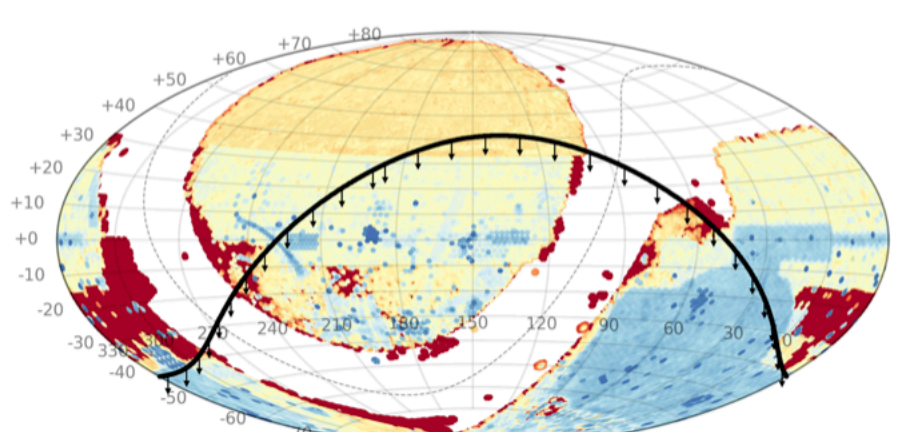


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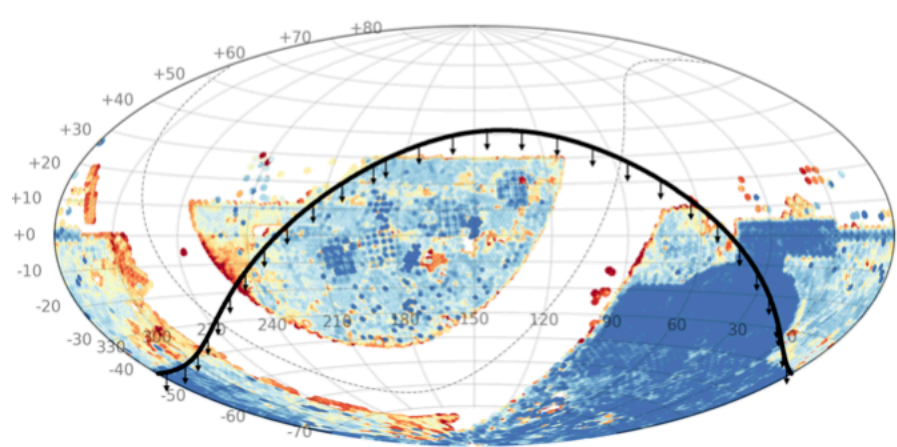
g-band depth (mag)



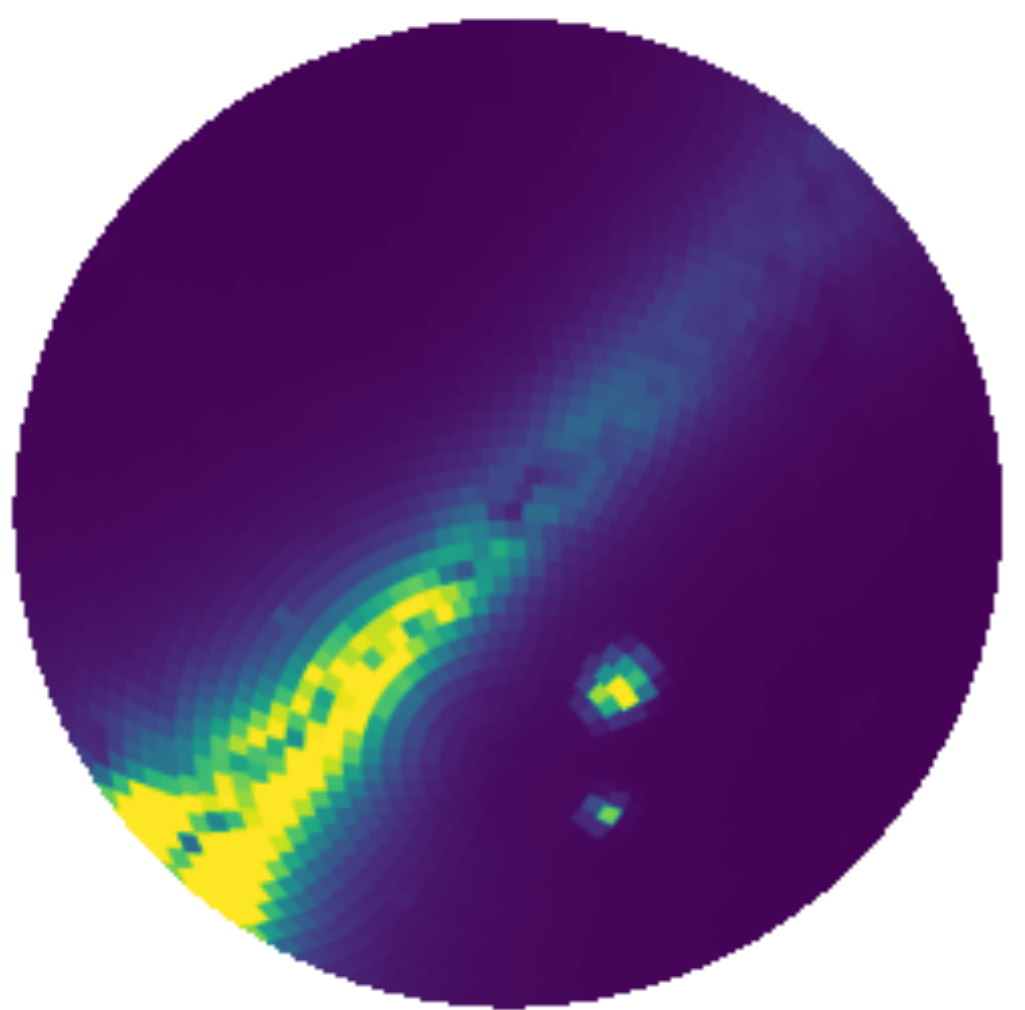
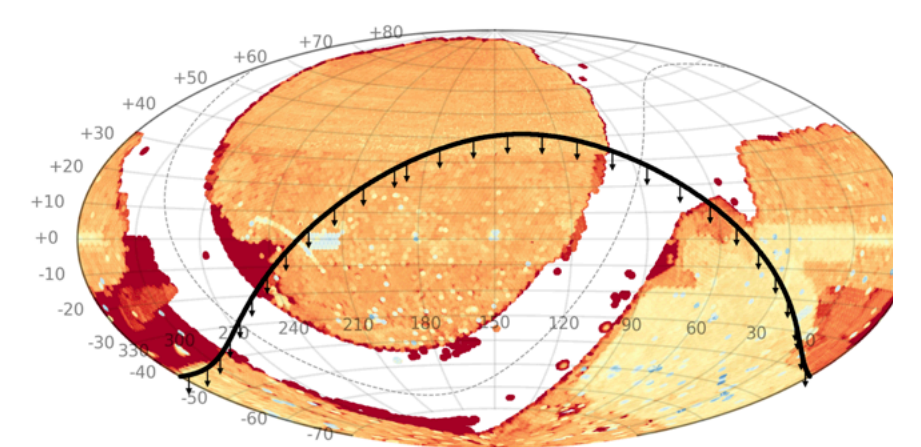
r-band depth (mag)



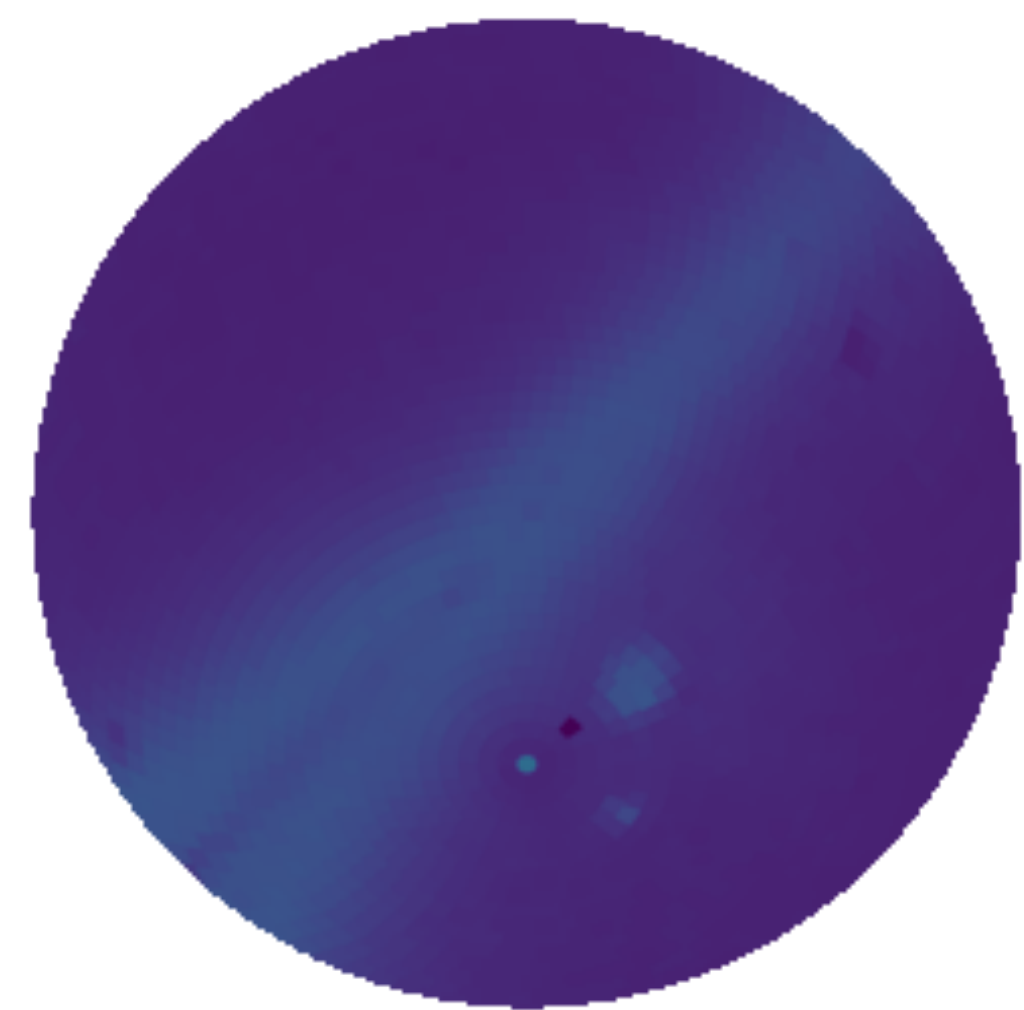
i-band depth (mag)



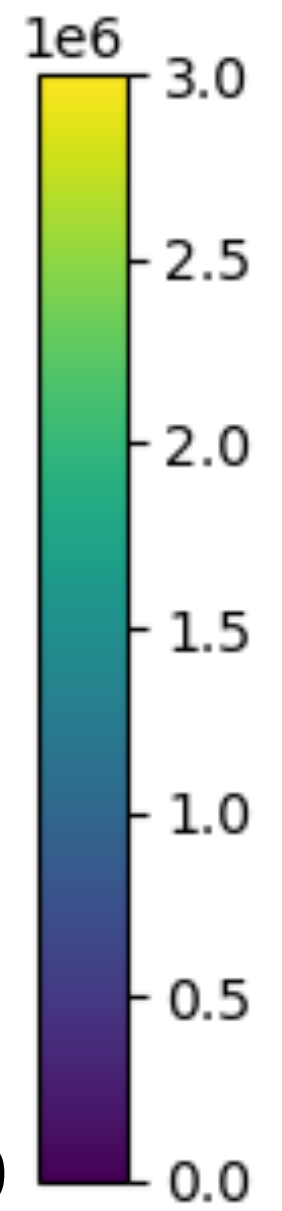
z-band depth (mag)



Gaia DR3



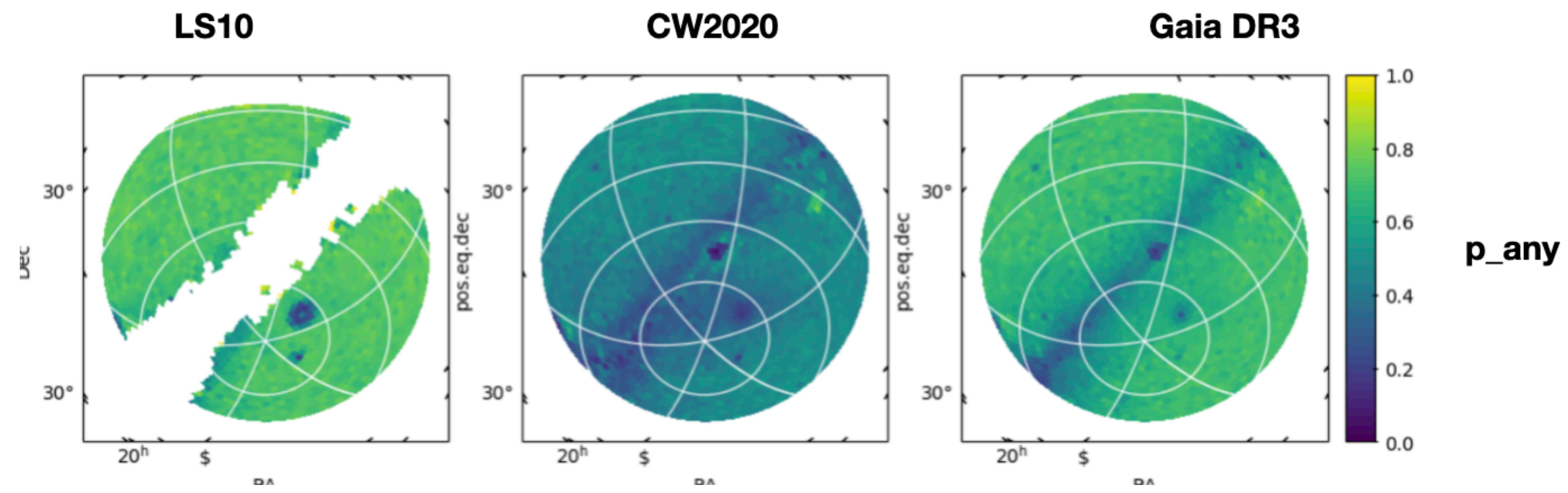
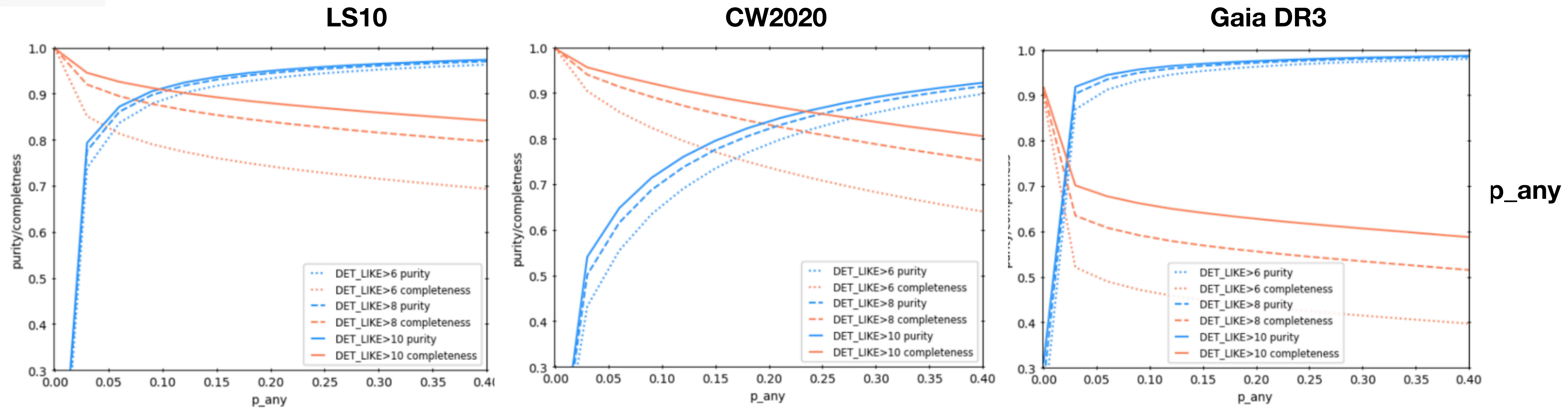
CW2020







# eRASS1 counterparts (NWAY, Salvato+2018)



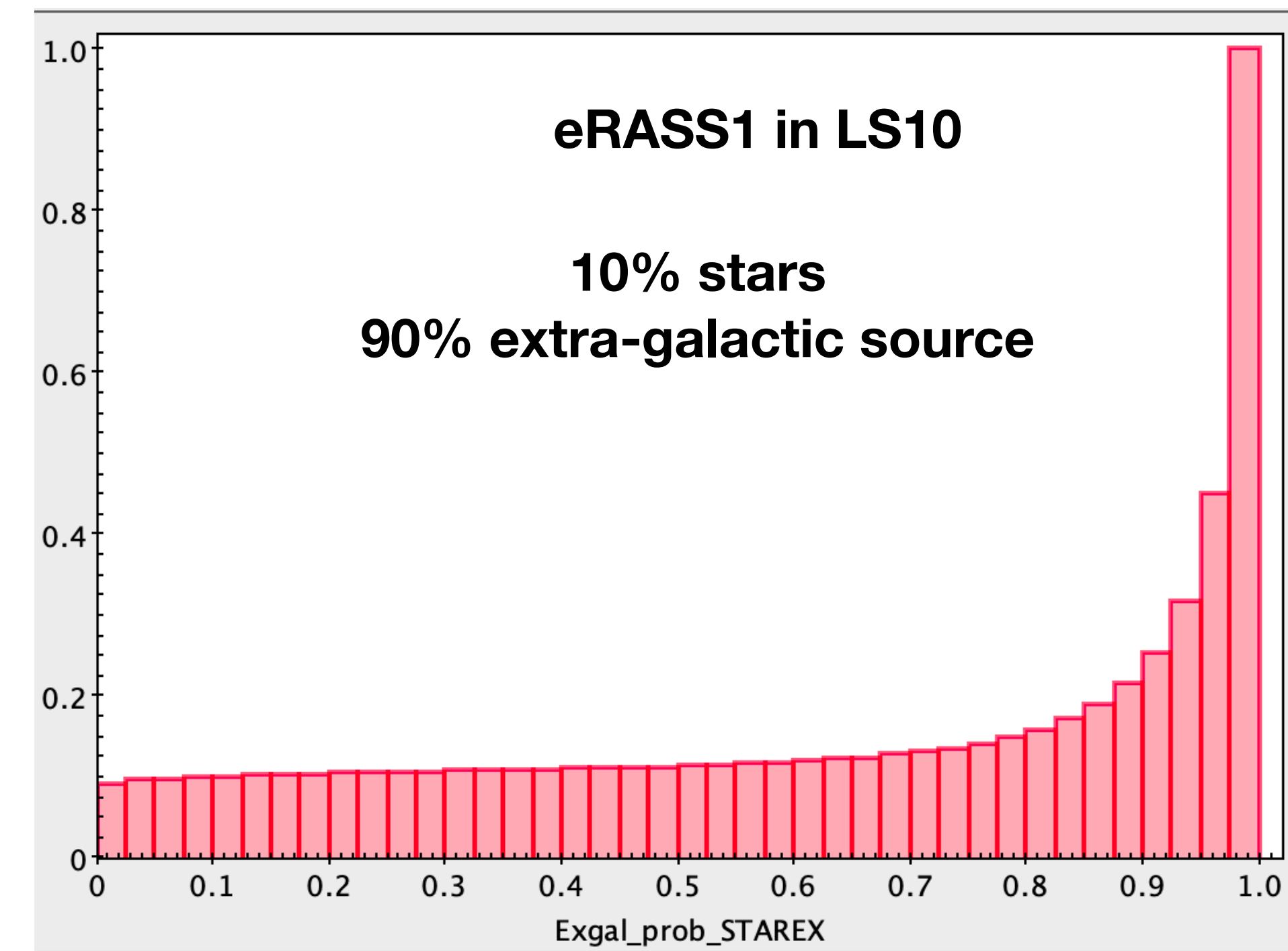
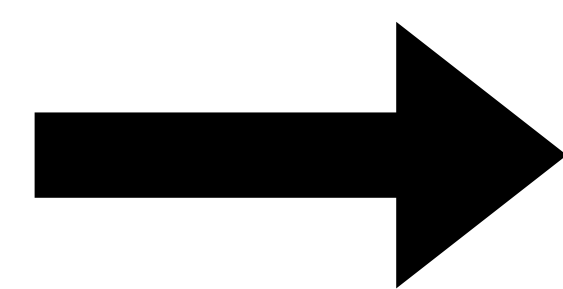
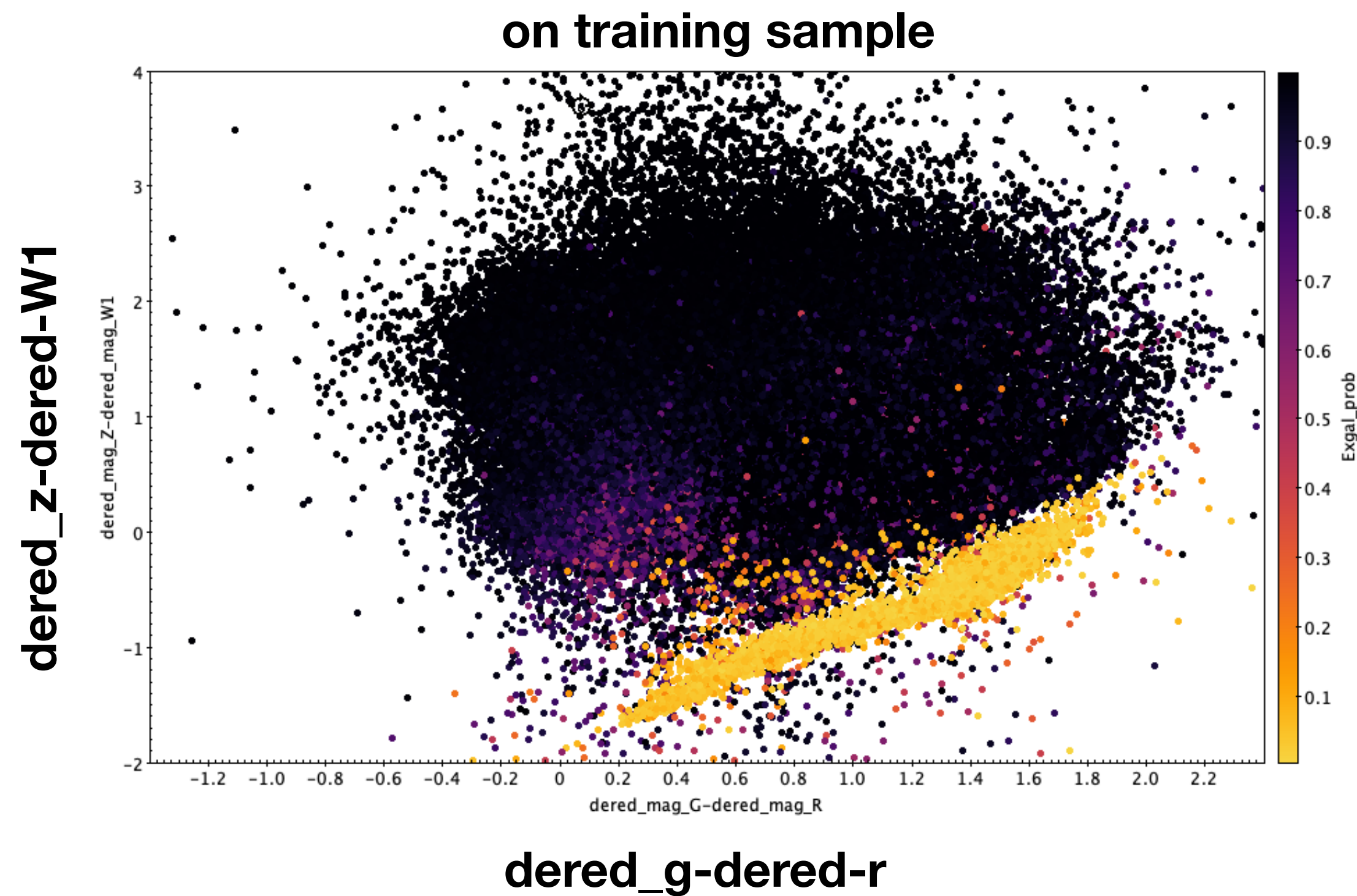




# Classification via STAREX (Shashwat Tiwari)



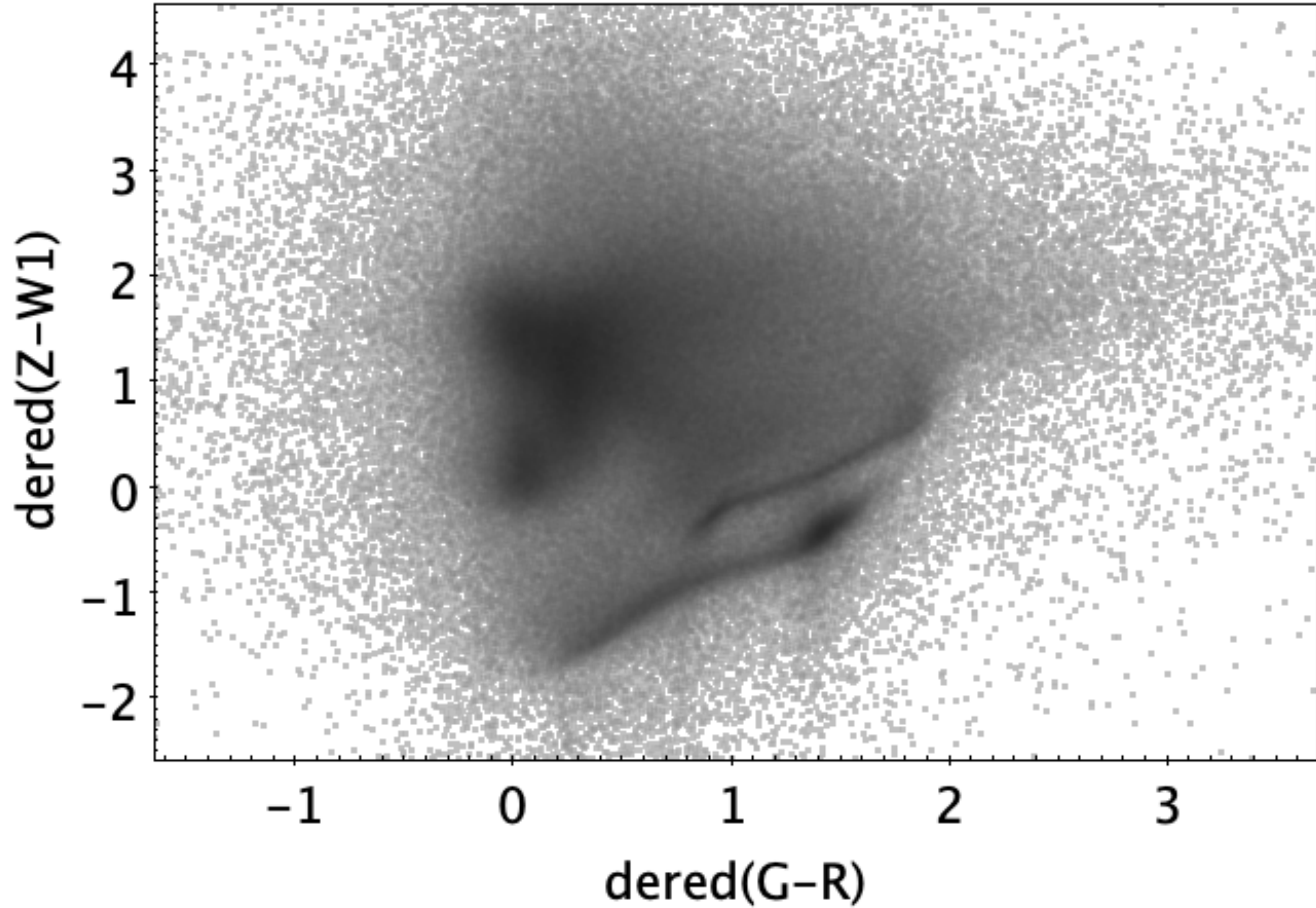
Training sample of 22252 correctly classified sources (7735 galactic and 14517 extragalactic)  
6 features from eROSITA and LS10 (X-ray flux, g,r,z,w1, TYPE ) + 4 features from Gaia (mag and parallax), when available  
Accuracy = >98%







750 000 sources in the LS10 area

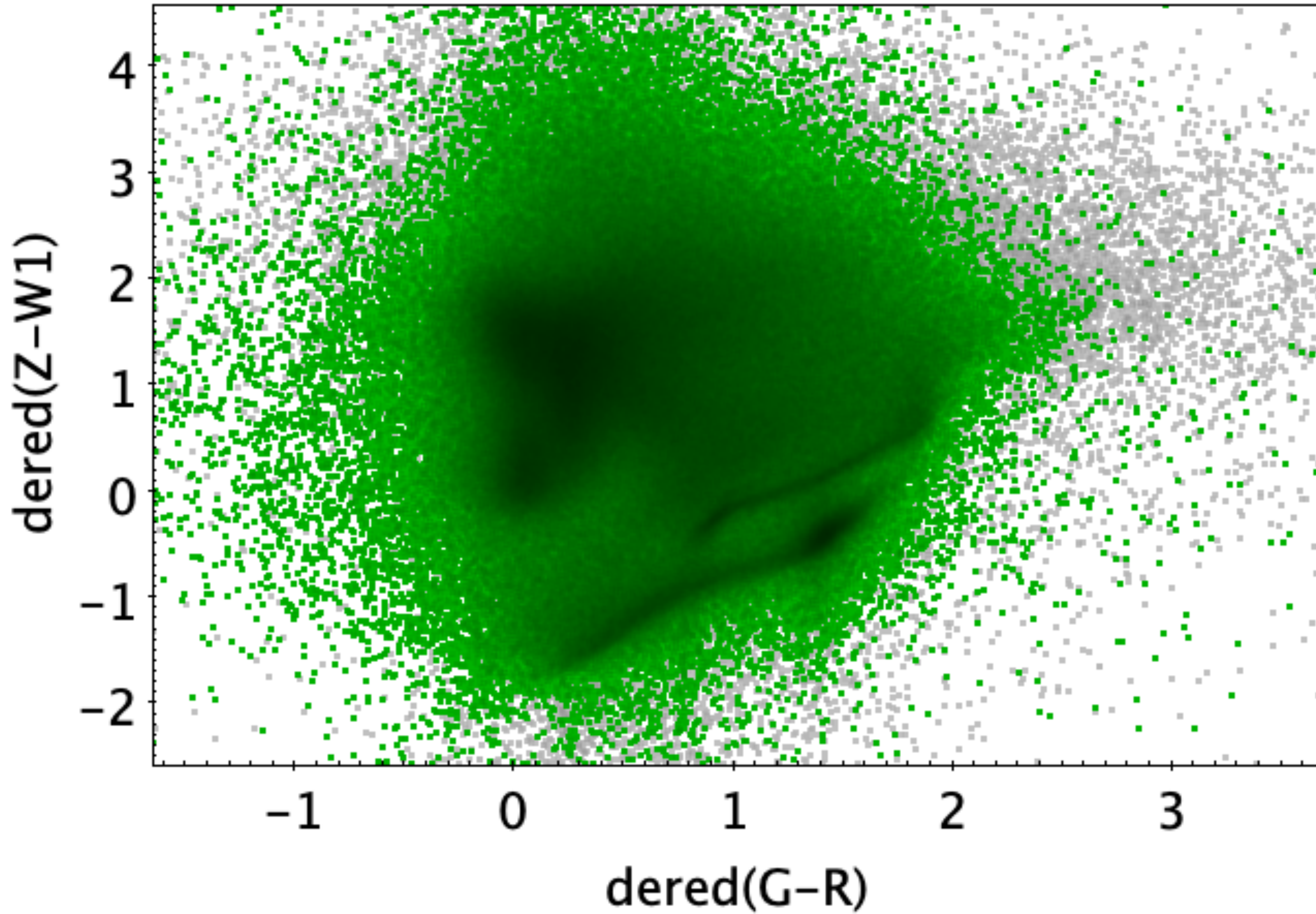


Salvato et al 2024





750 000 sources in the LS10 area



Salvato et al 2024

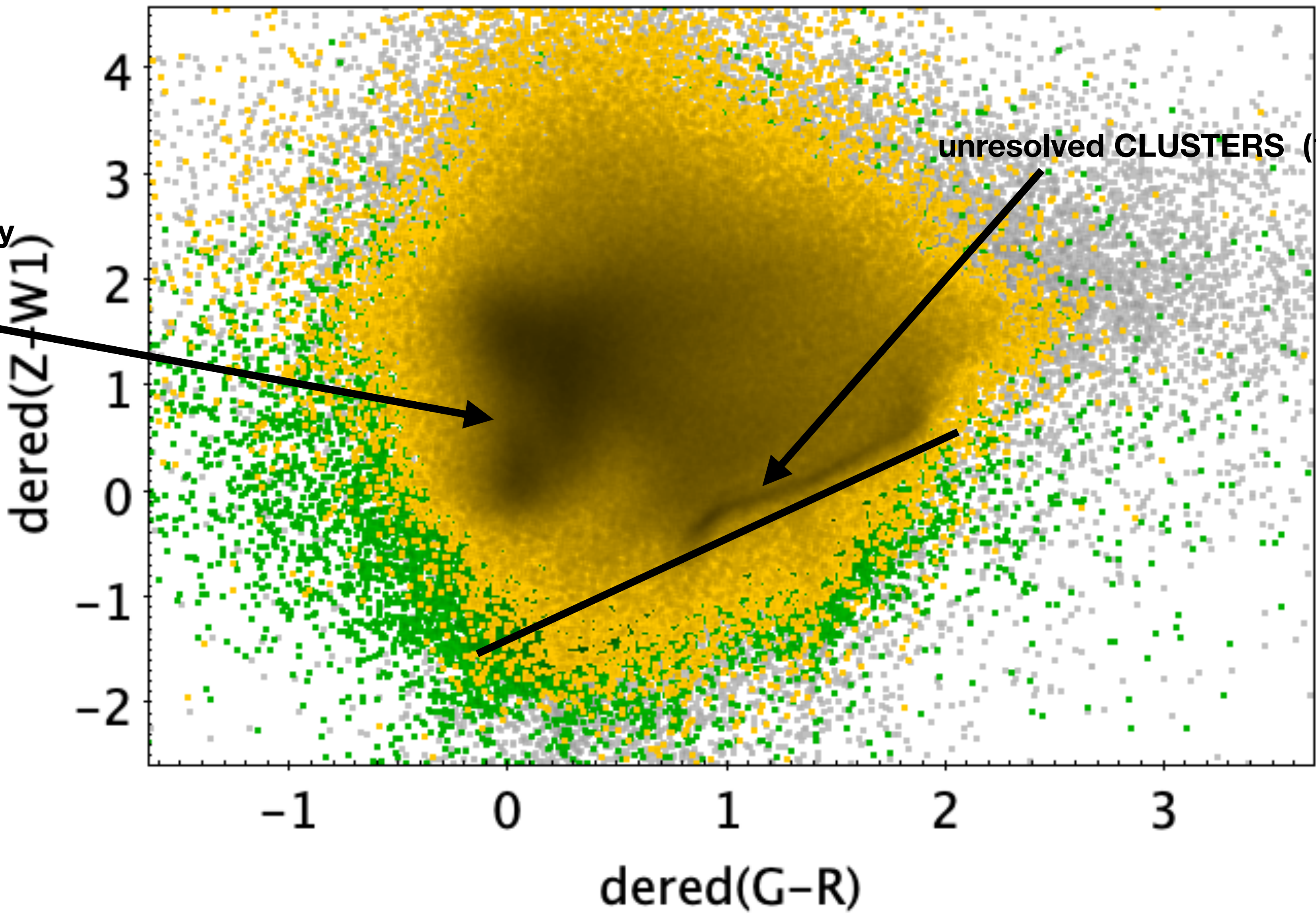




750 000 sources in the LS10 area



about 550k AGN  
with good photometry  
(SNR>3 in 6 bands)



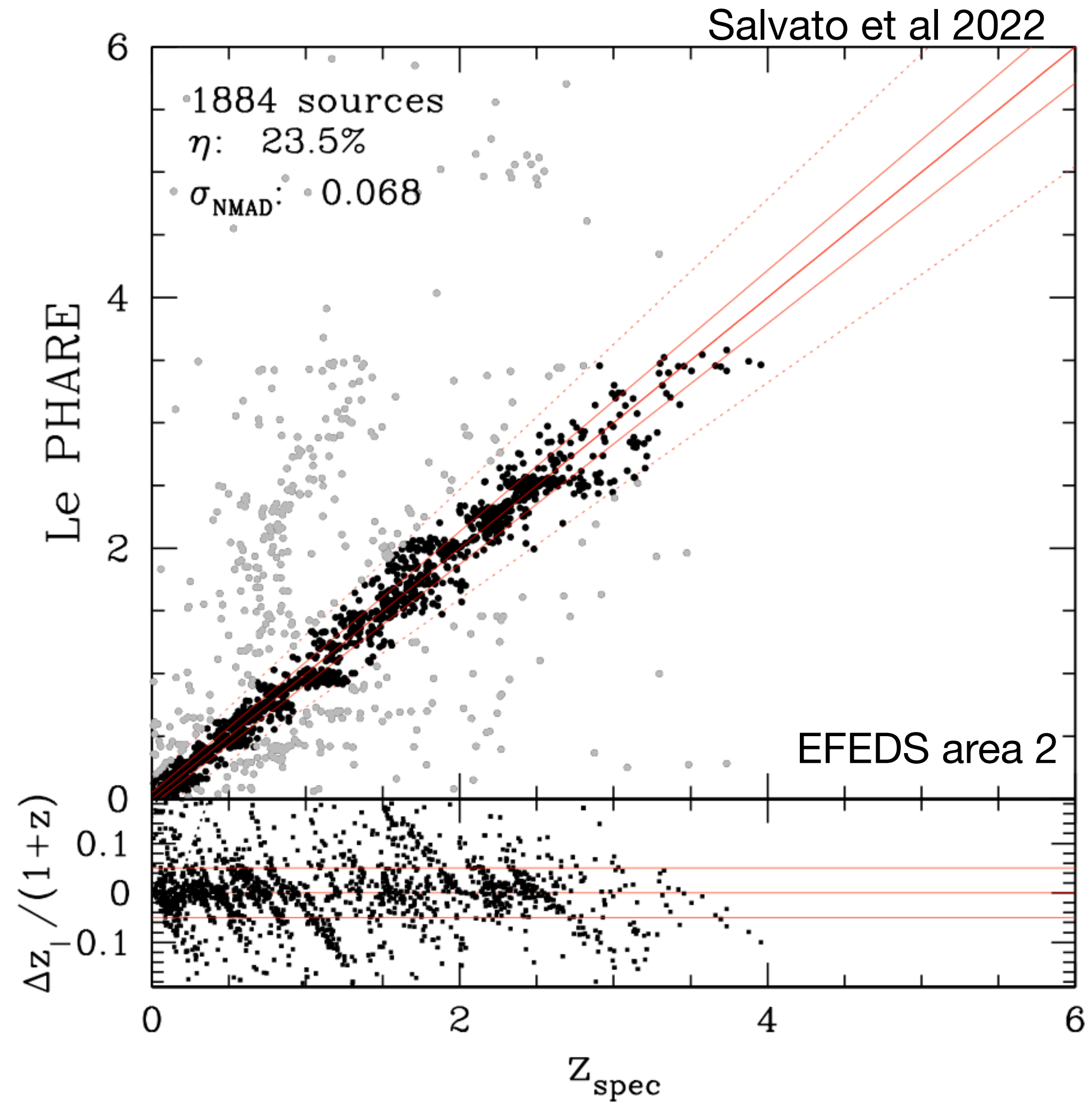
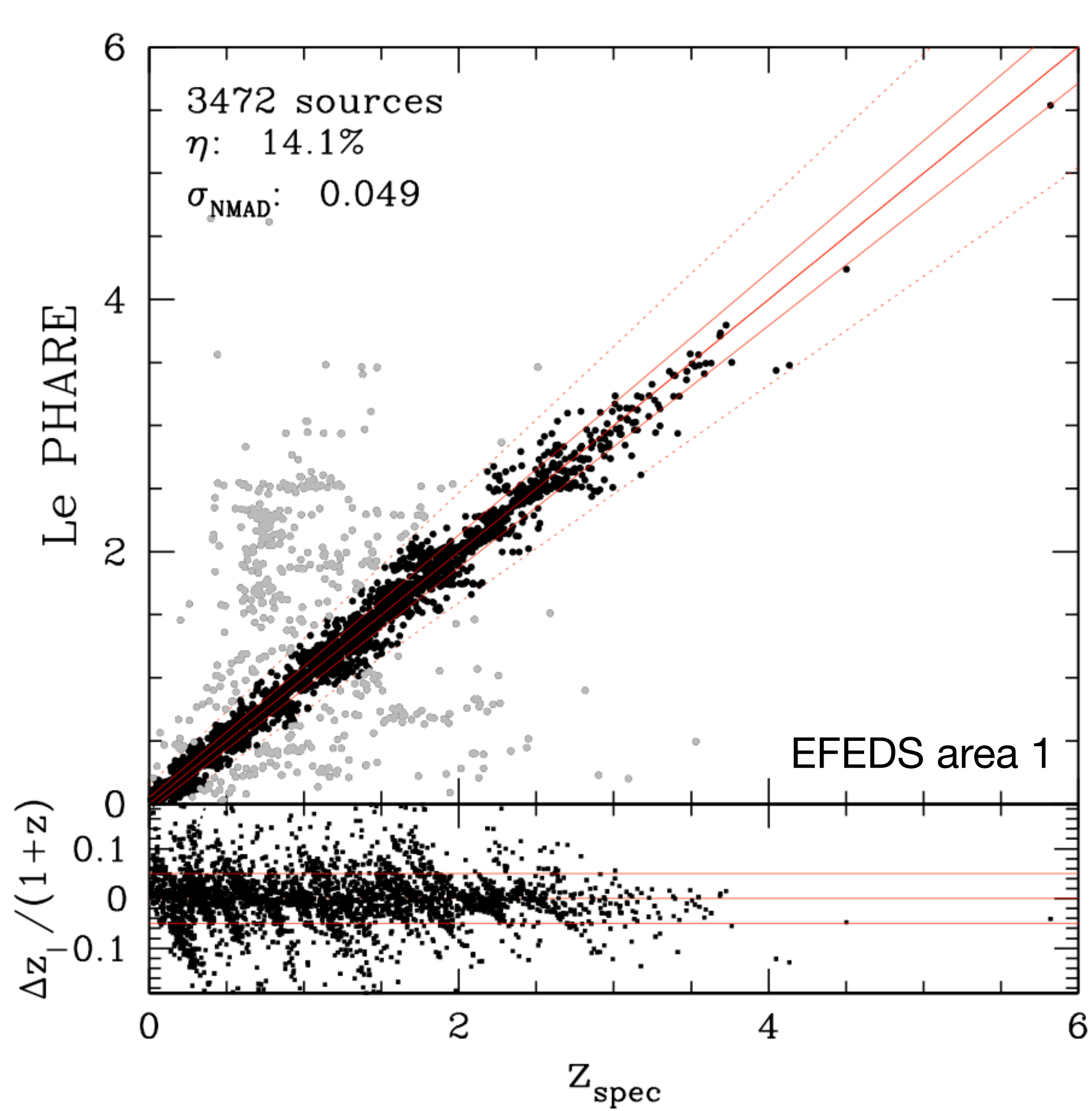
unresolved CLUSTERS (16k), Balzer et al in prep)

Salvato et al 2024





# eRASS1 photoz better than in eFEDS!



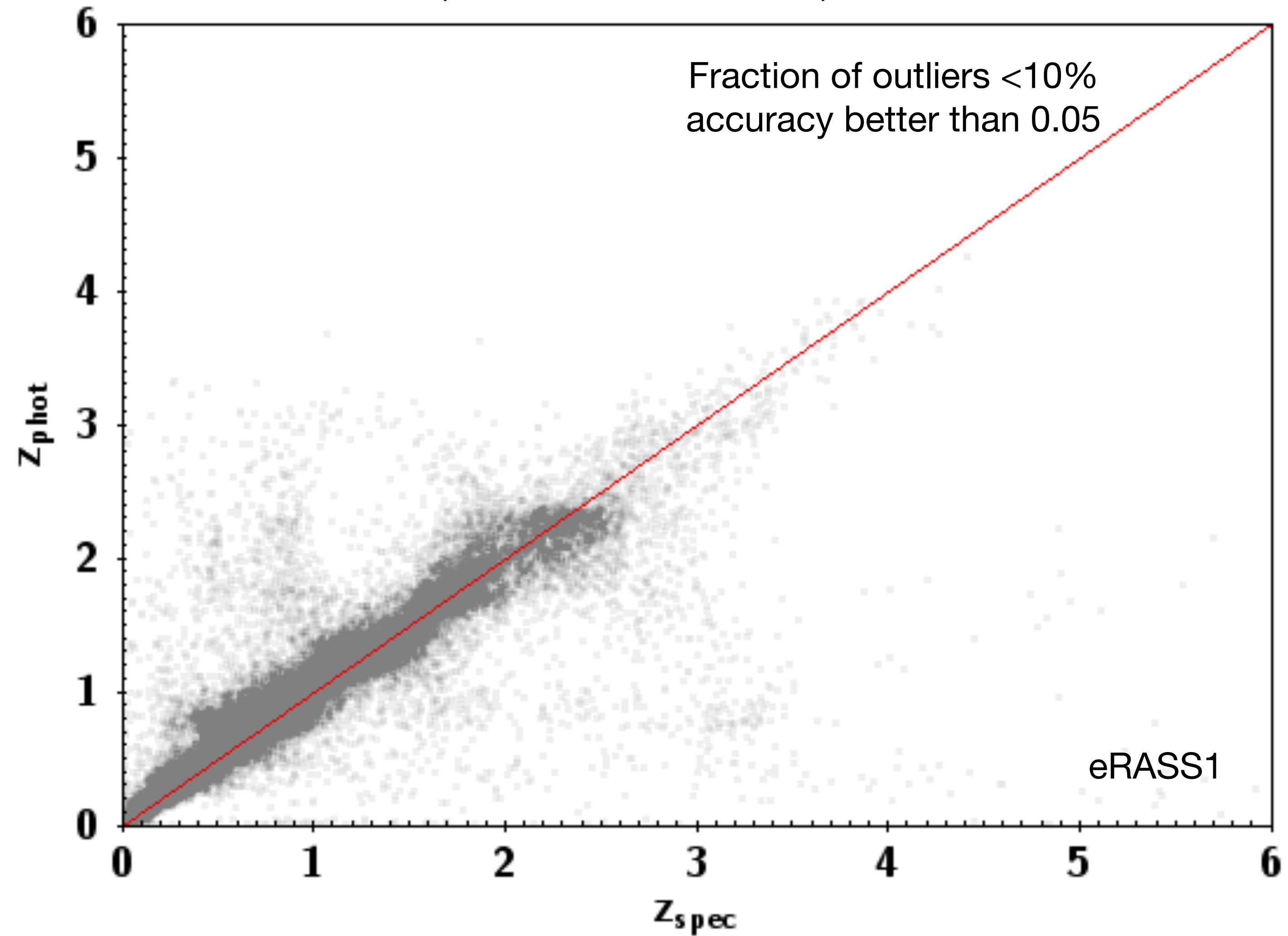
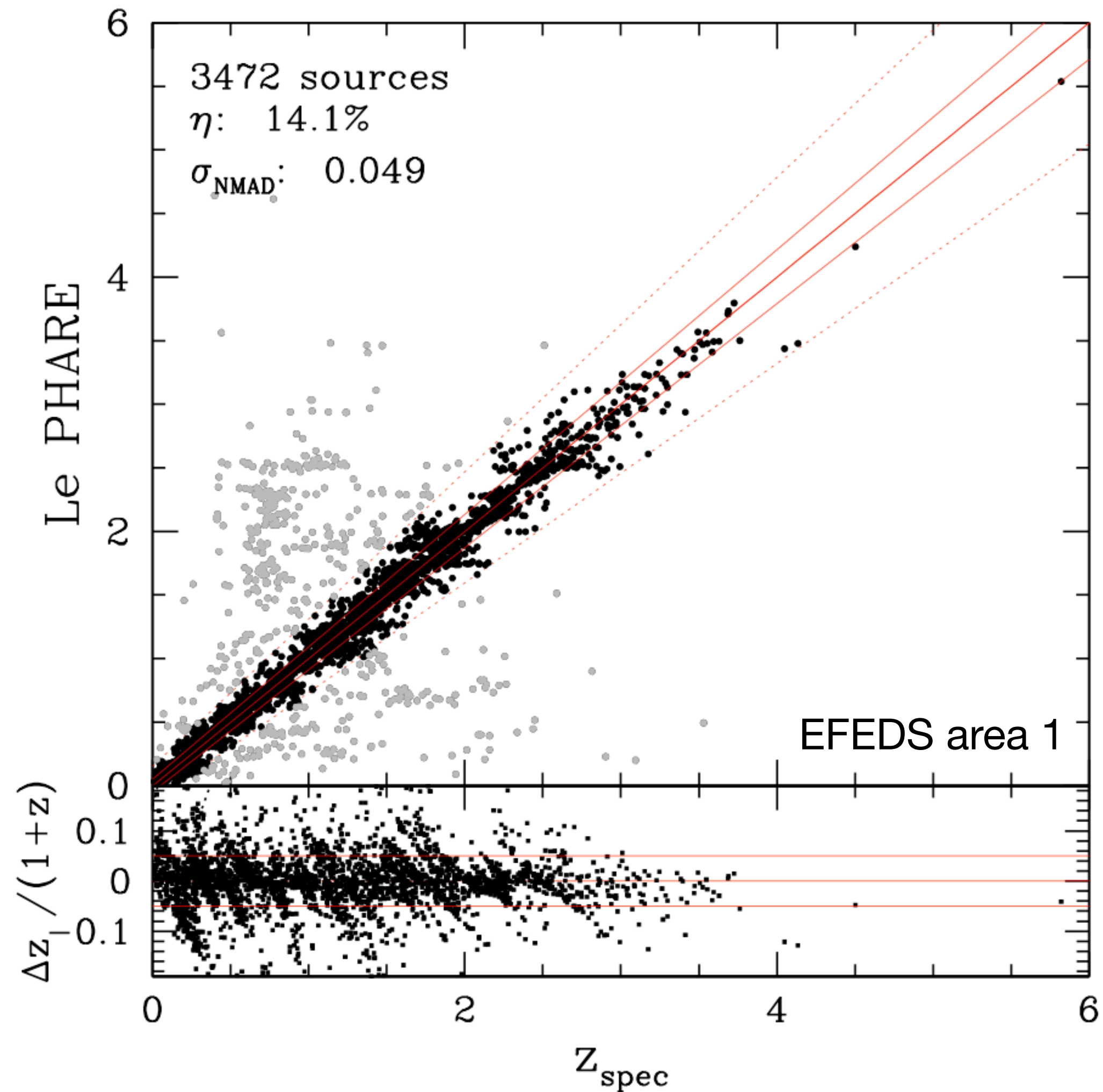




# eRASS1 photoz better than in eFEDS!



Using CIRCLEZ (ML on parametrized LS10 images)  
(Saxena, MS, et al 2024)



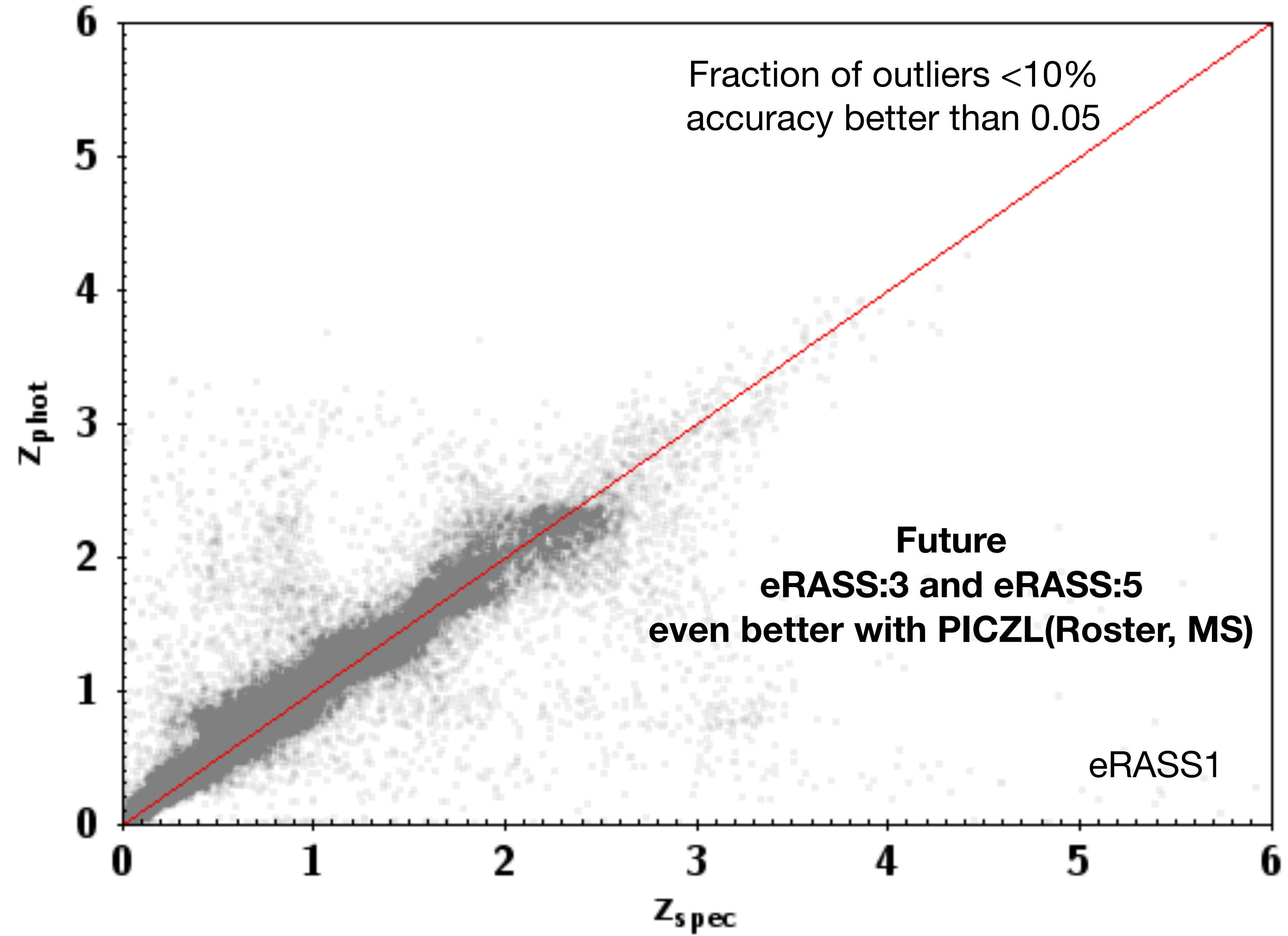
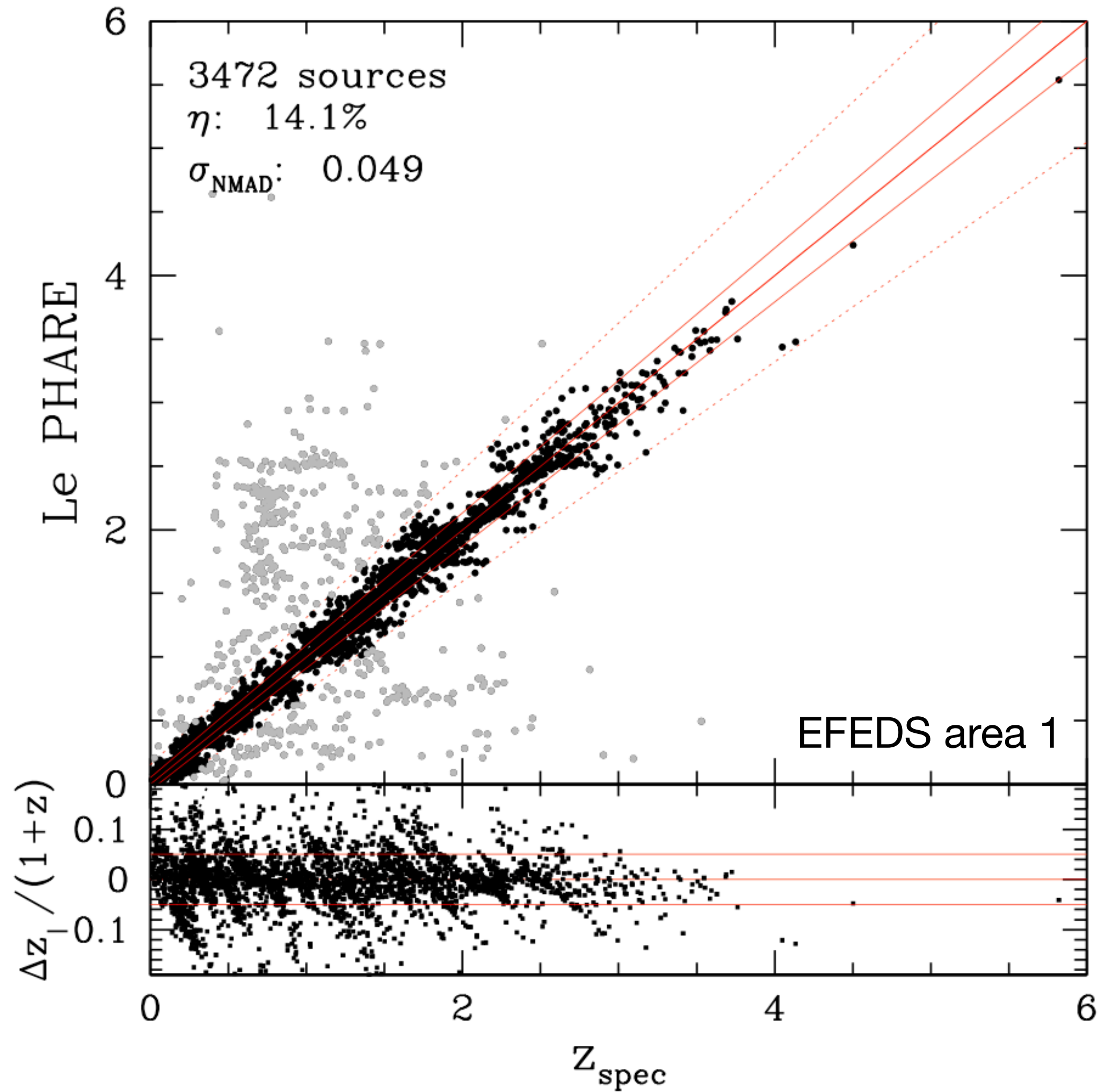




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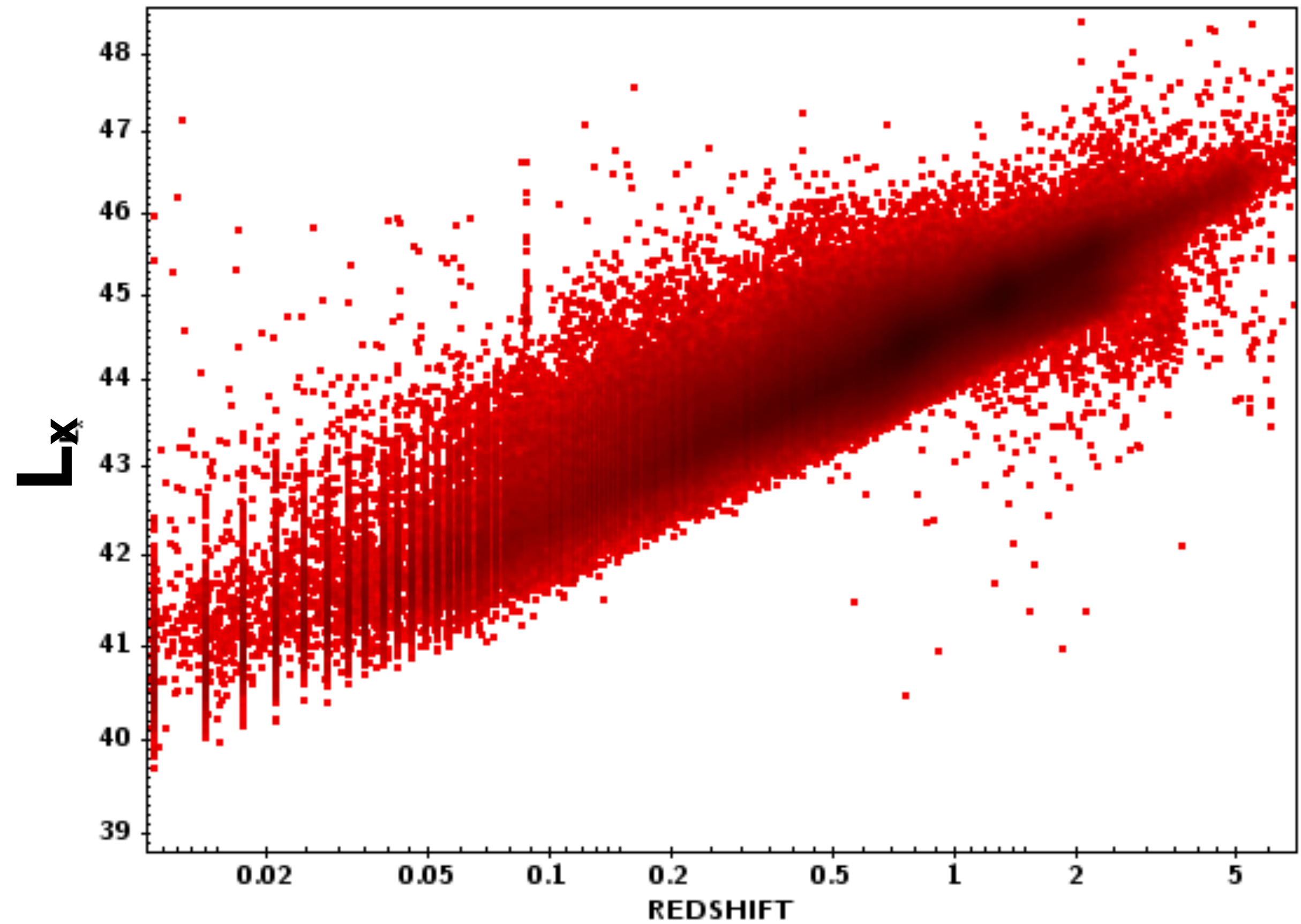
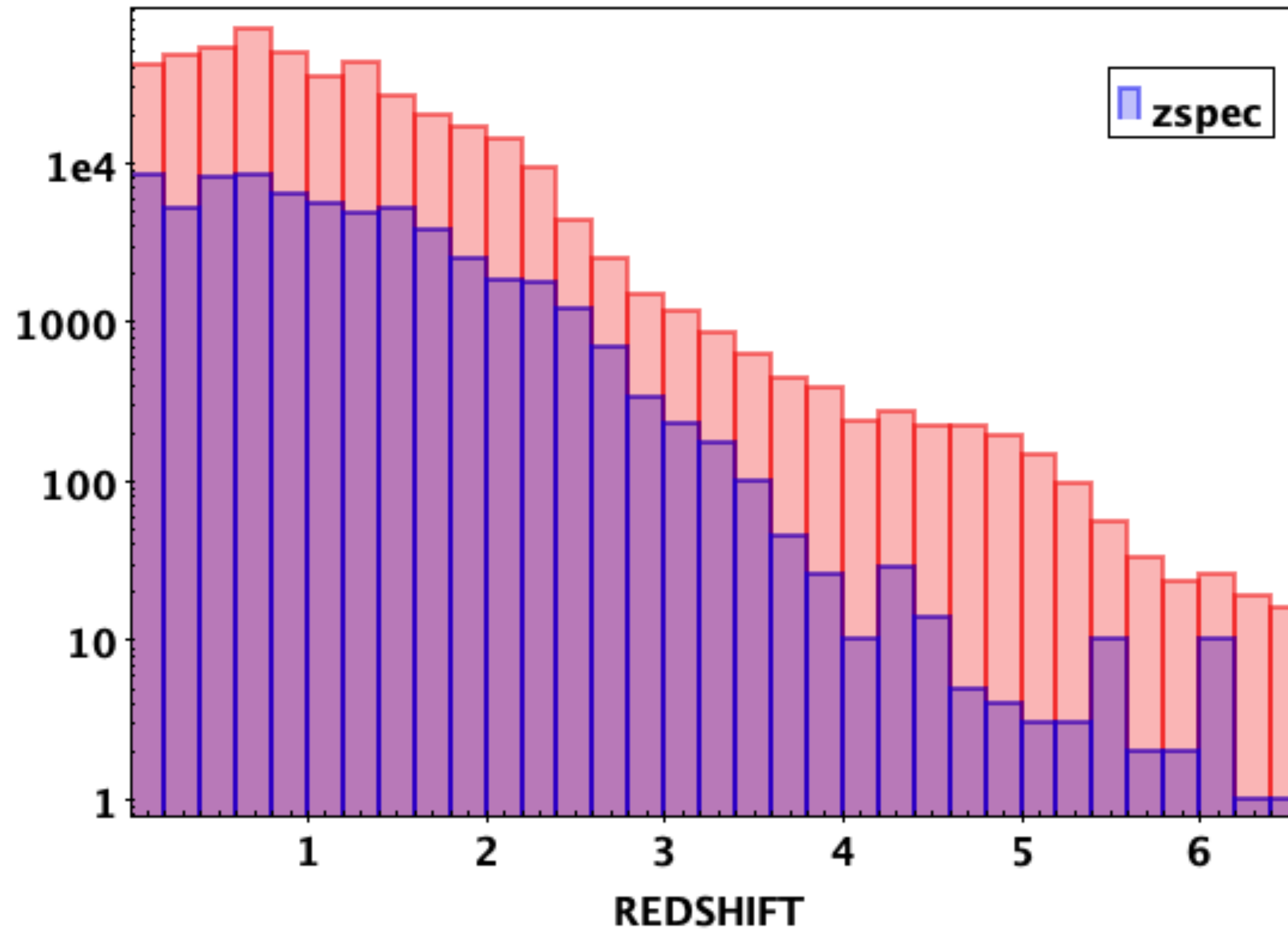


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# eRASS1/ redshift and luminosity distribution

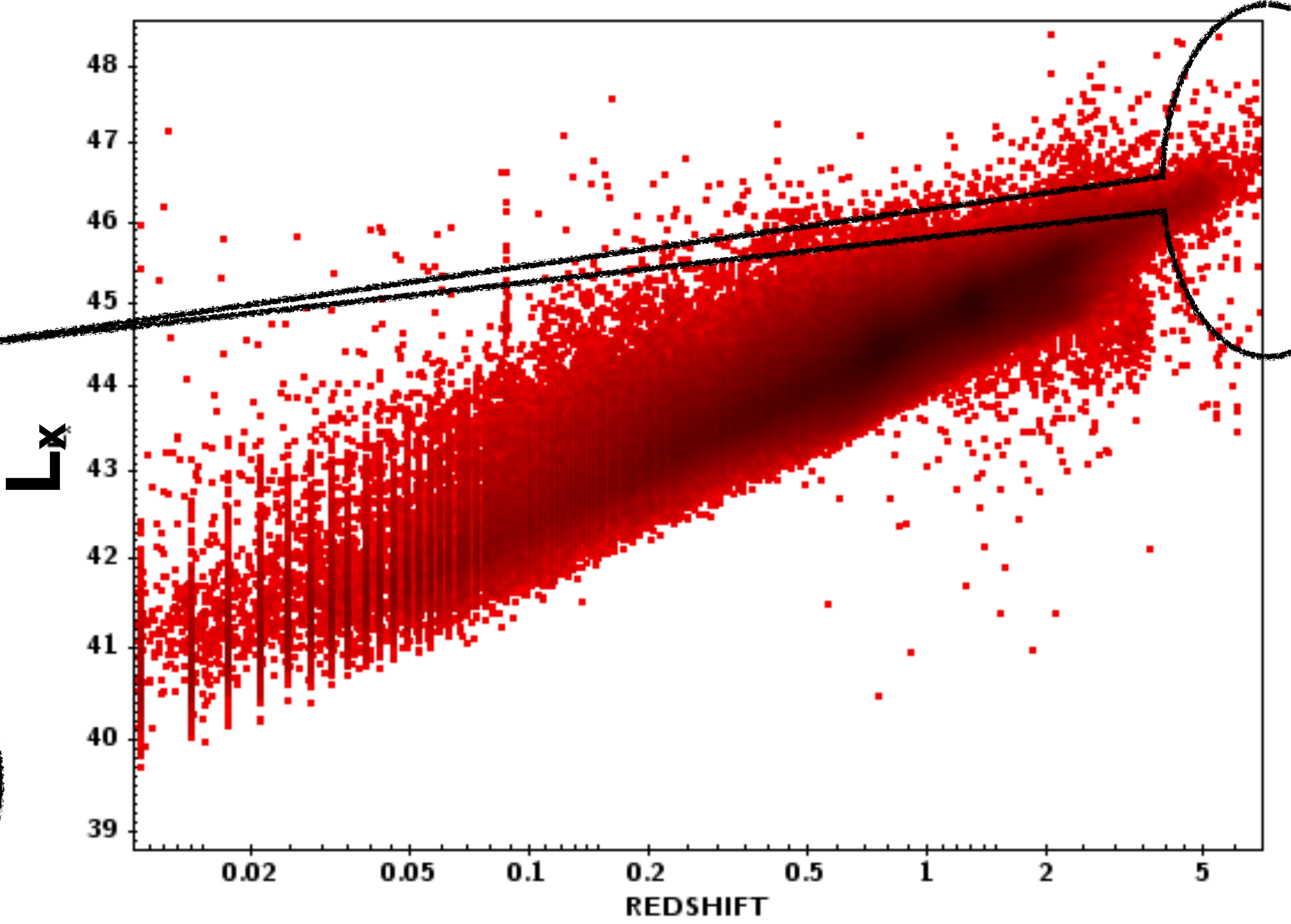
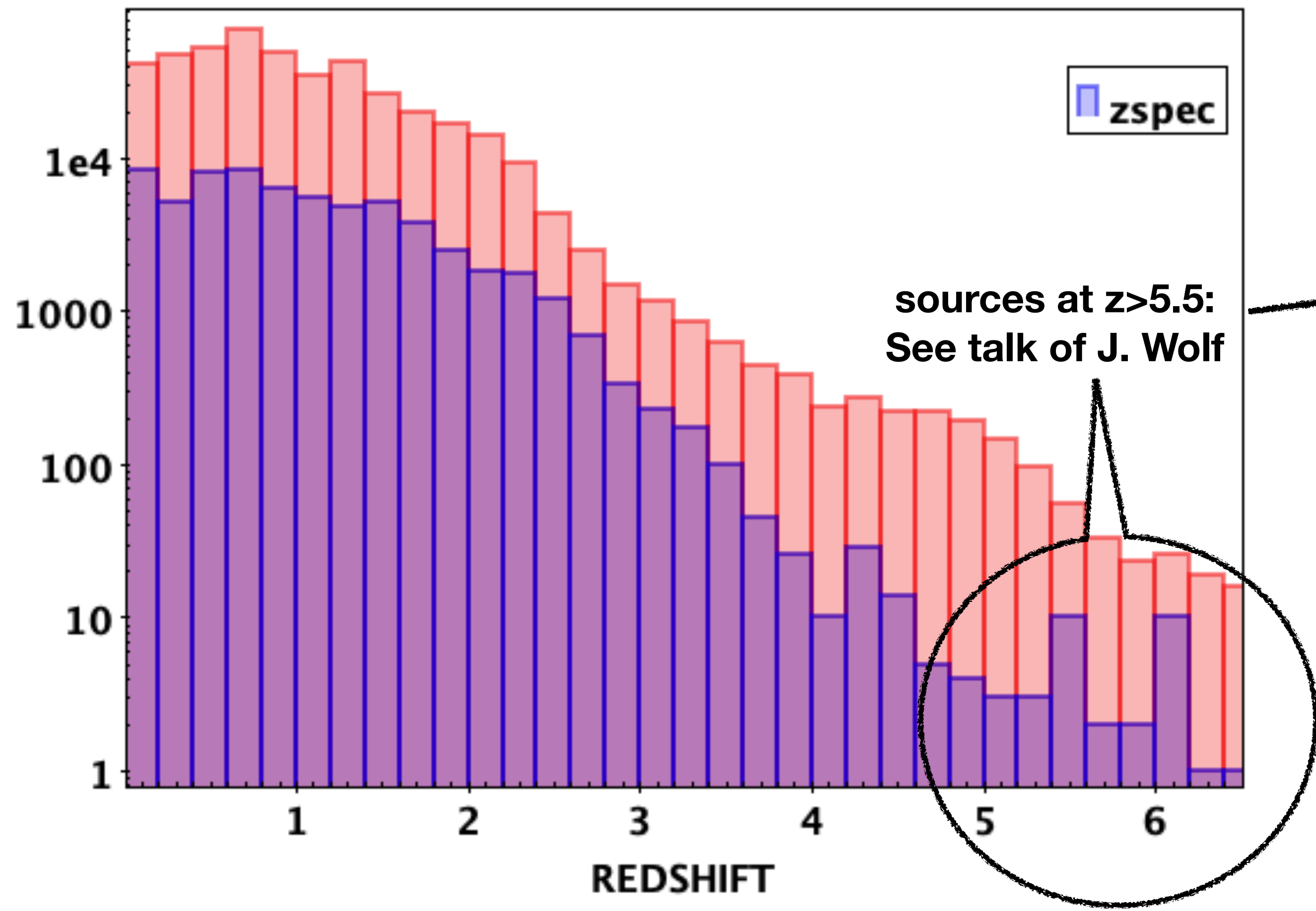


~200k redshifts from literature, mostly Quiaia in the plot

more to come from SDSS-V, DESI, 4MOST



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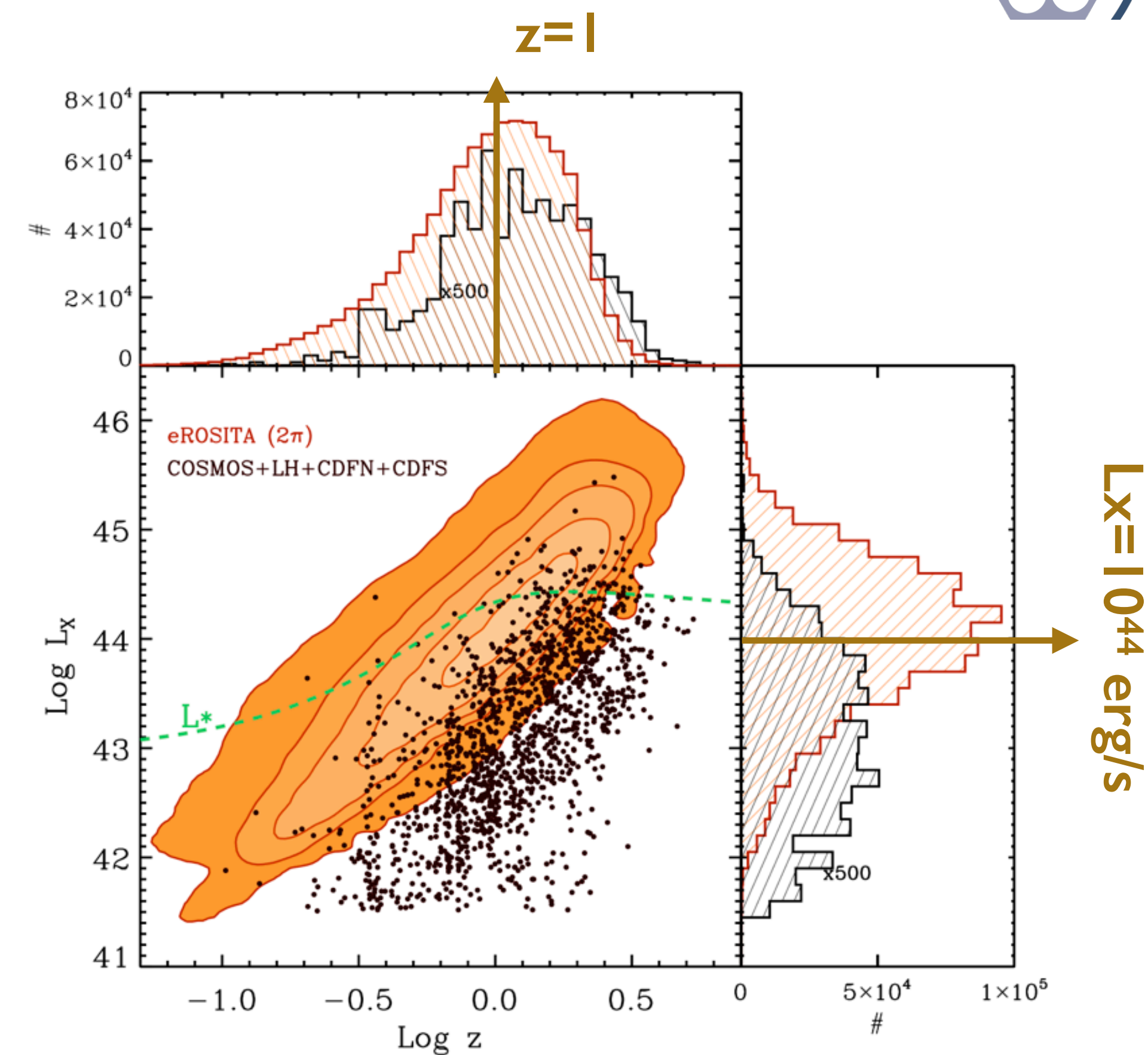
# eROSITA AGN surveys in contest







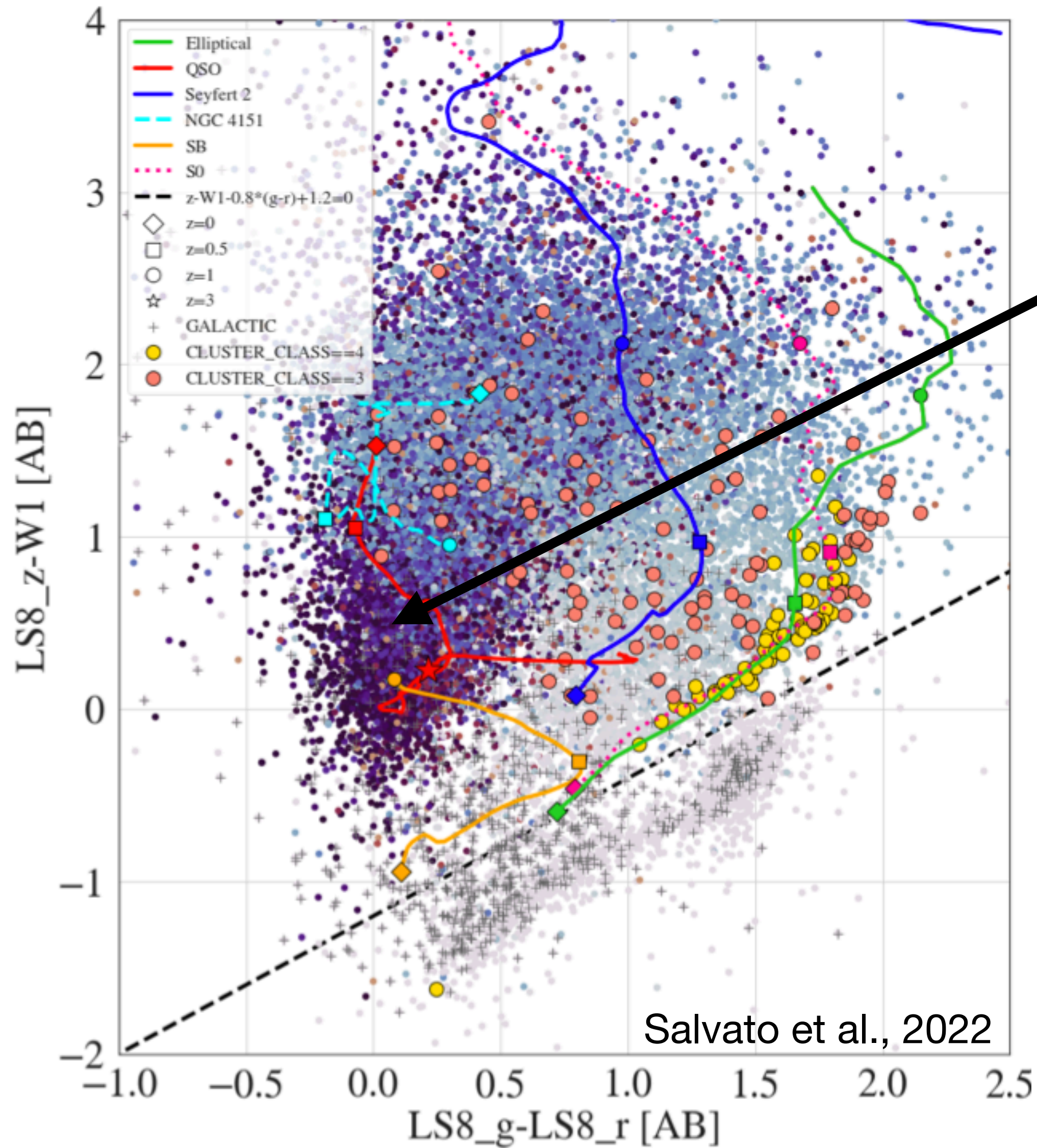
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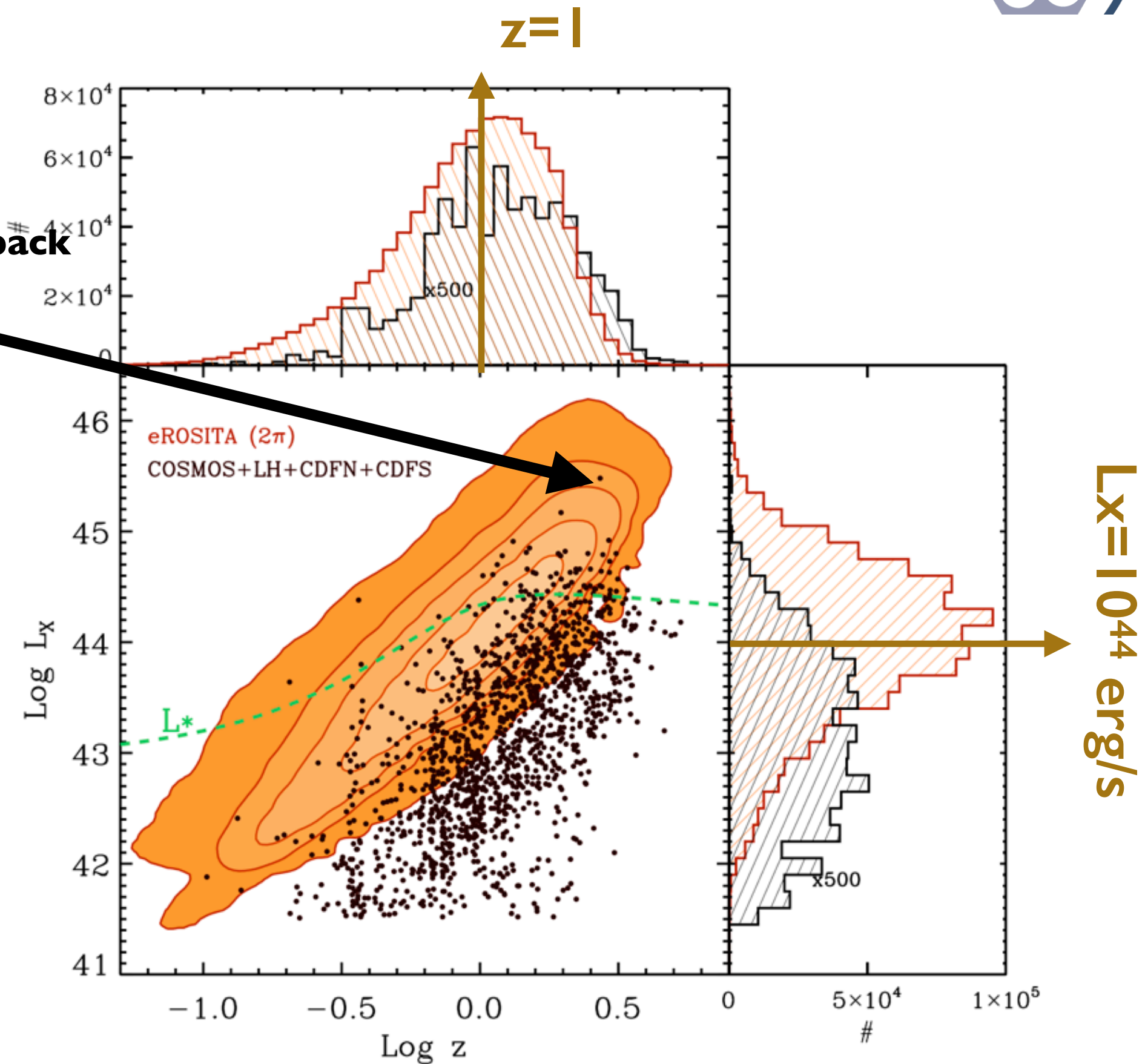




# eROSITA AGN surveys in contest



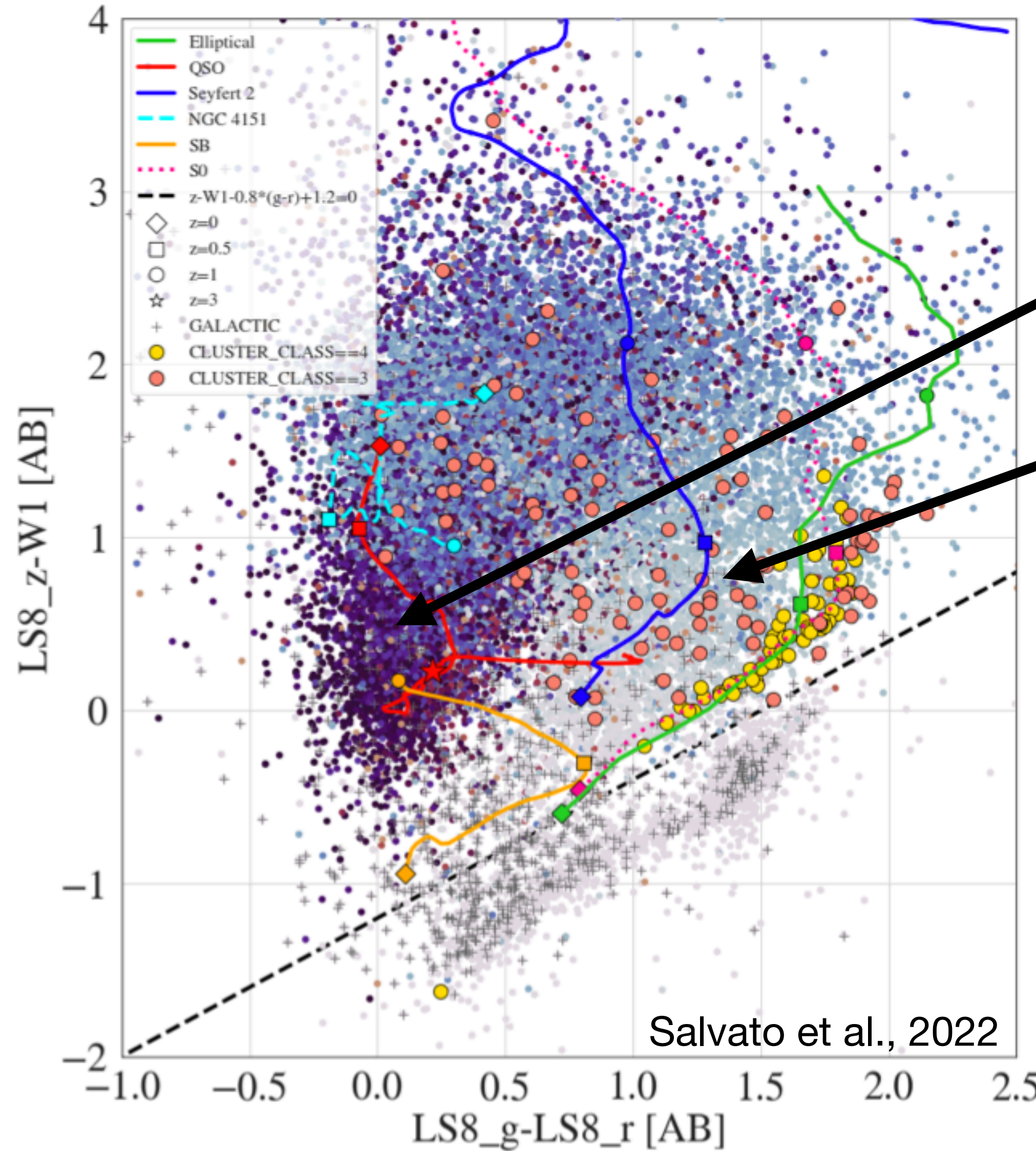
**The most luminous AGN,  
tracers of large-scale structure:  
the “quasar” mode of AGN feedback**





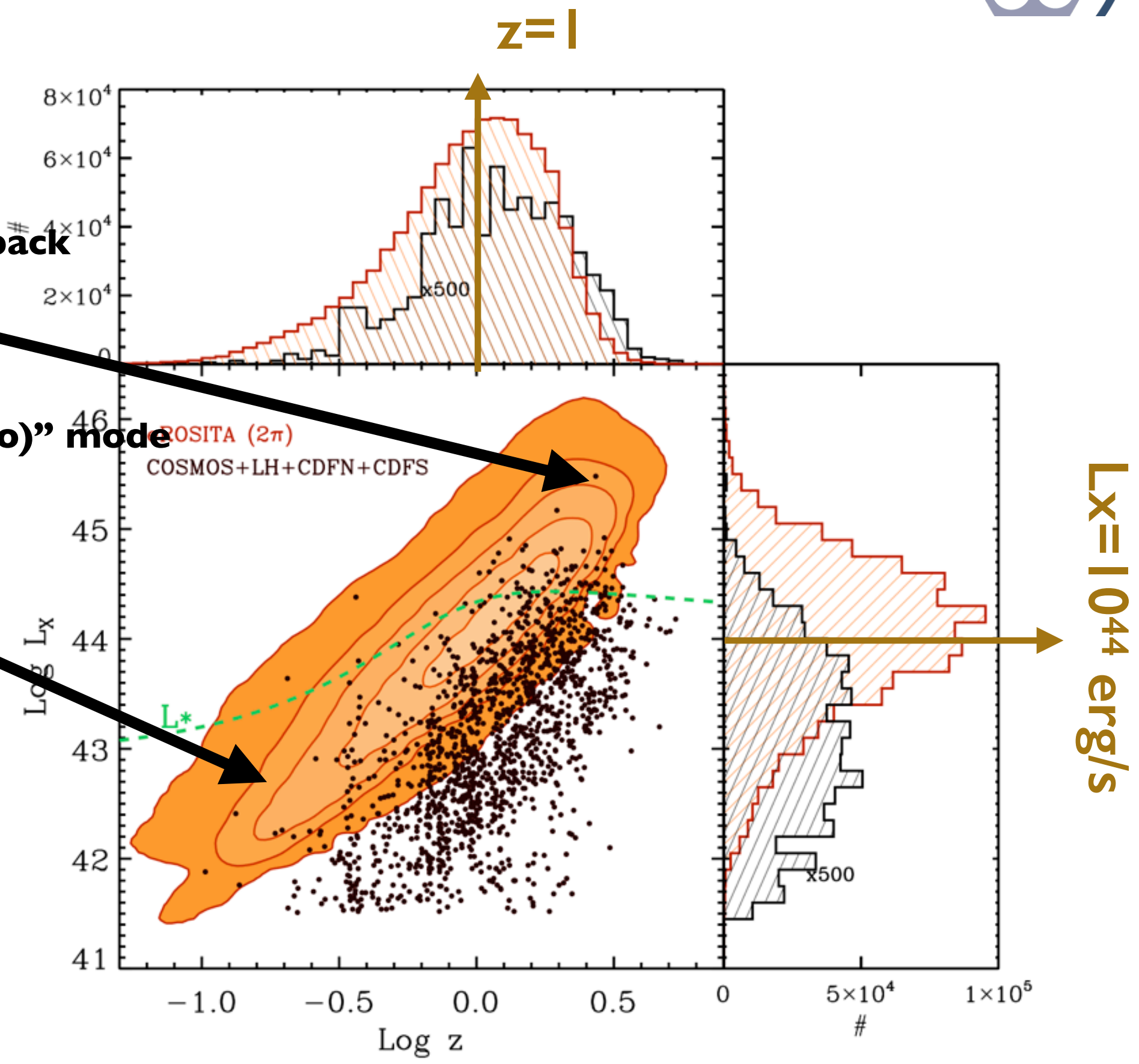


# eROSITA AGN surveys in contest



**The most luminous AGN, tracers of large-scale structure: the “quasar” mode of AGN feedback**

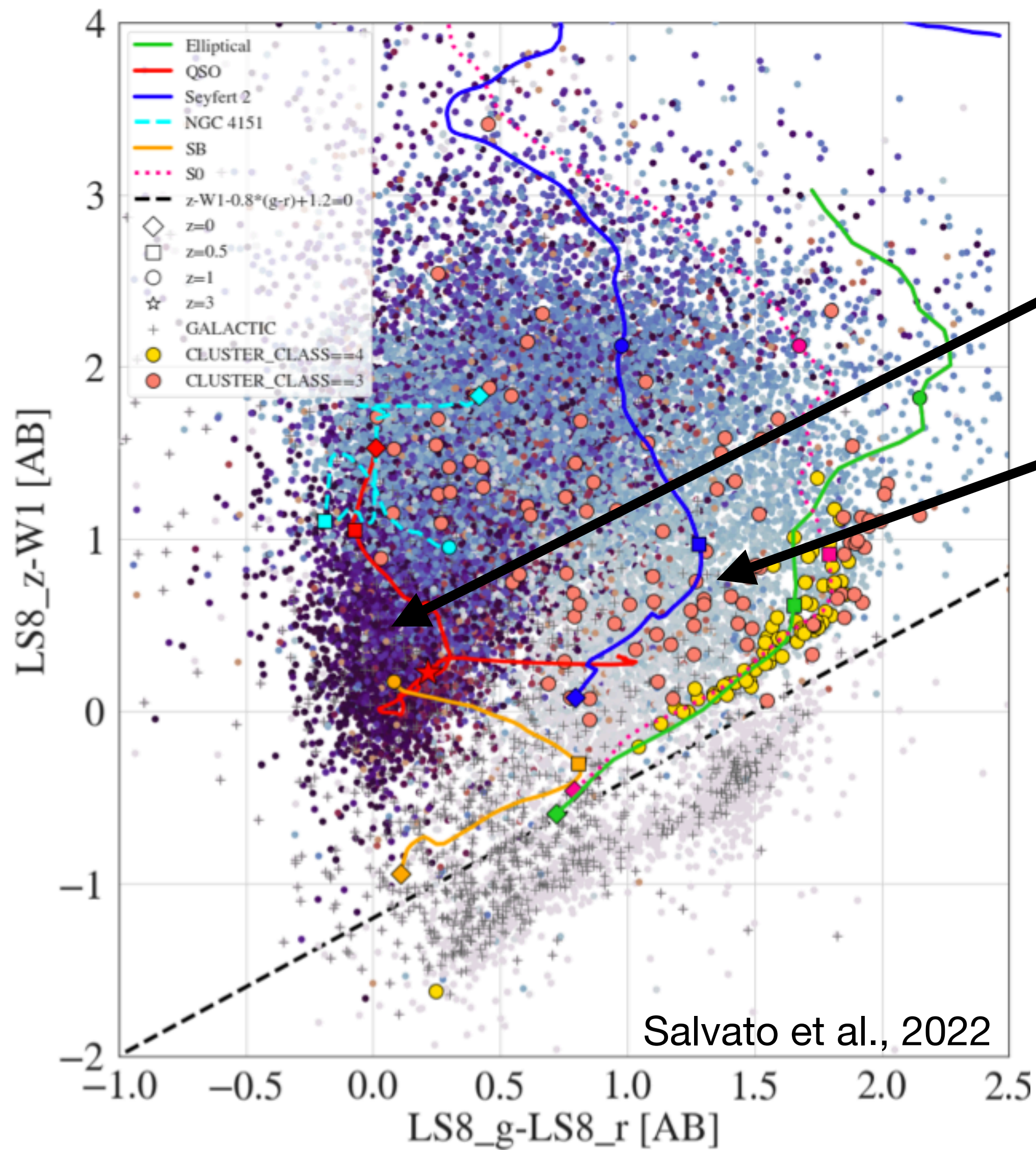
**Nearby LLAGN: the “kinetic (radio)” mode of AGN feedback**





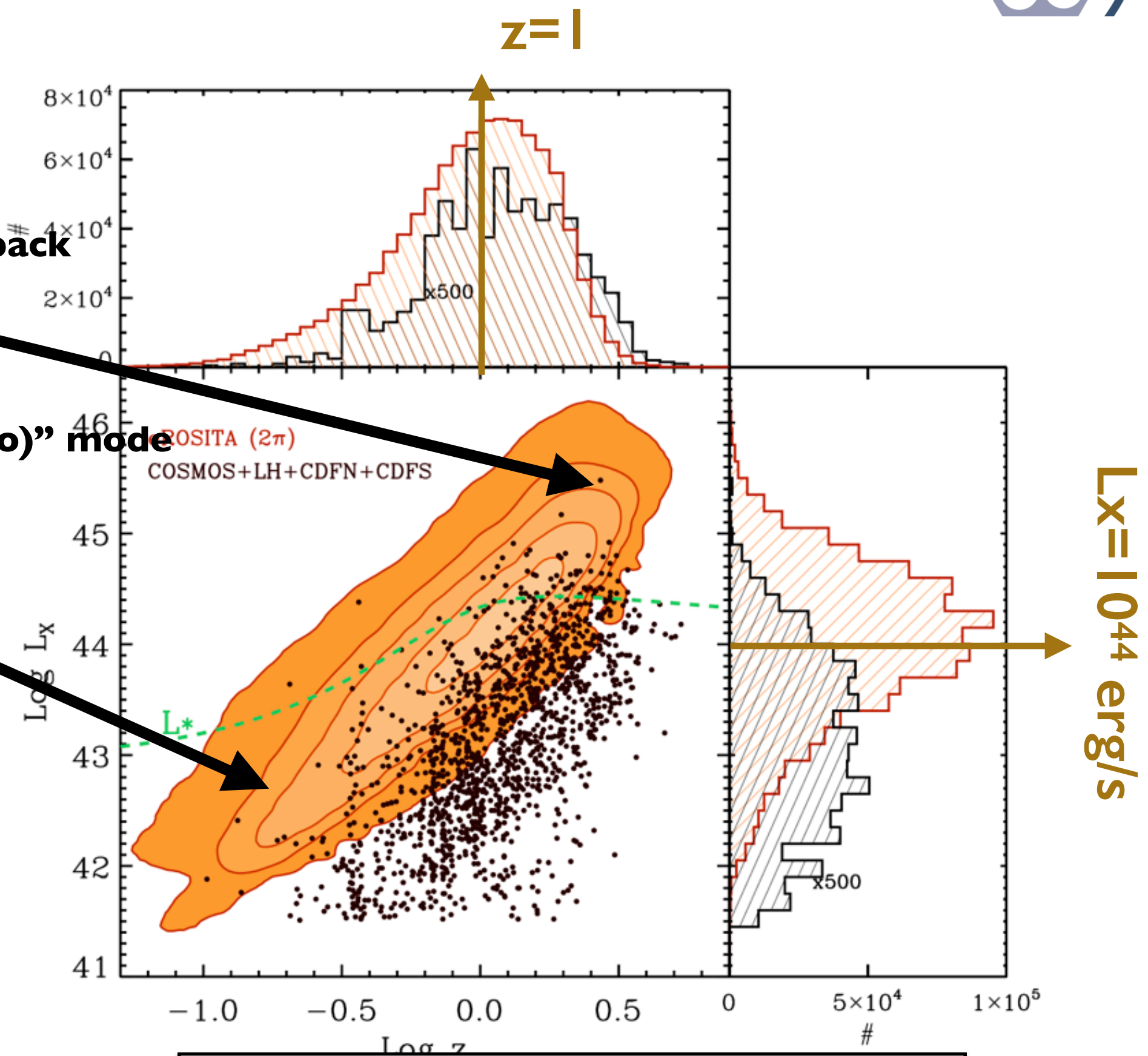


# eROSITA AGN surveys in contest



The most luminous AGN, tracers of large-scale structure: the “quasar” mode of AGN feedback

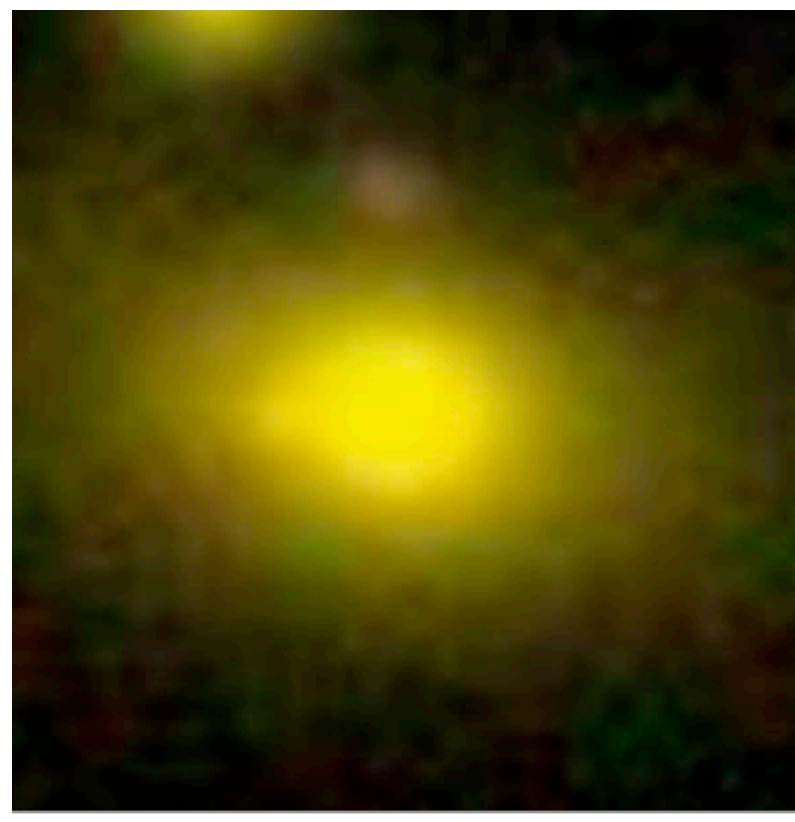
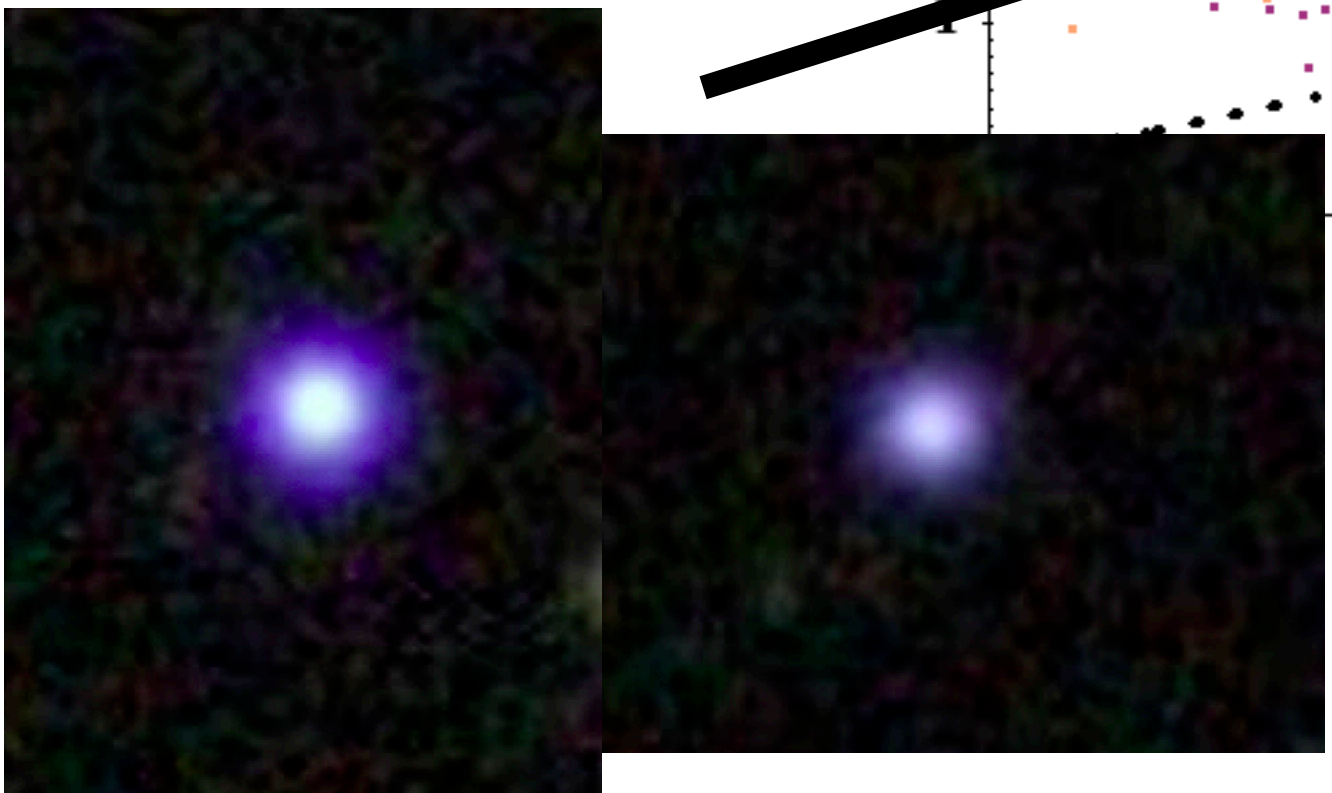
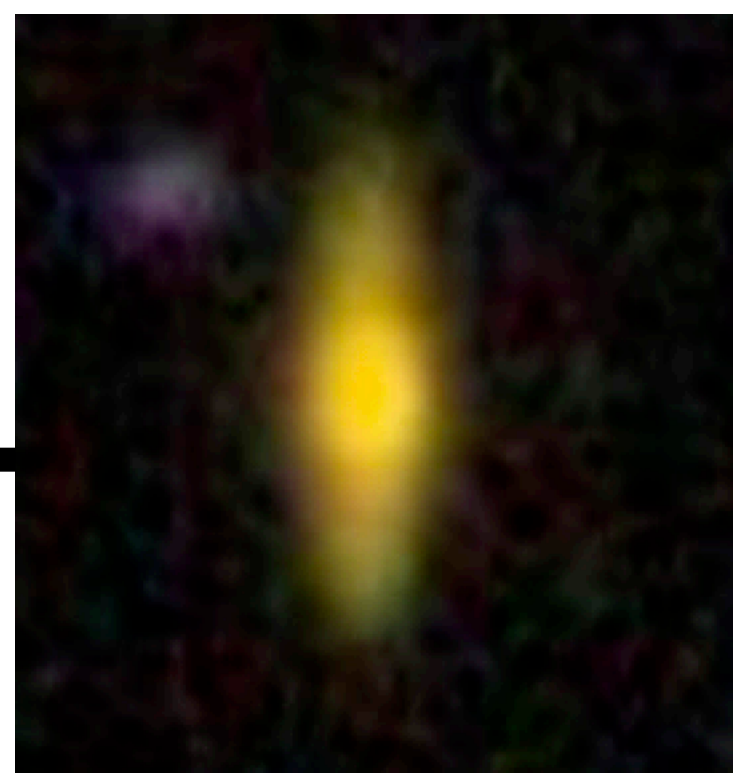
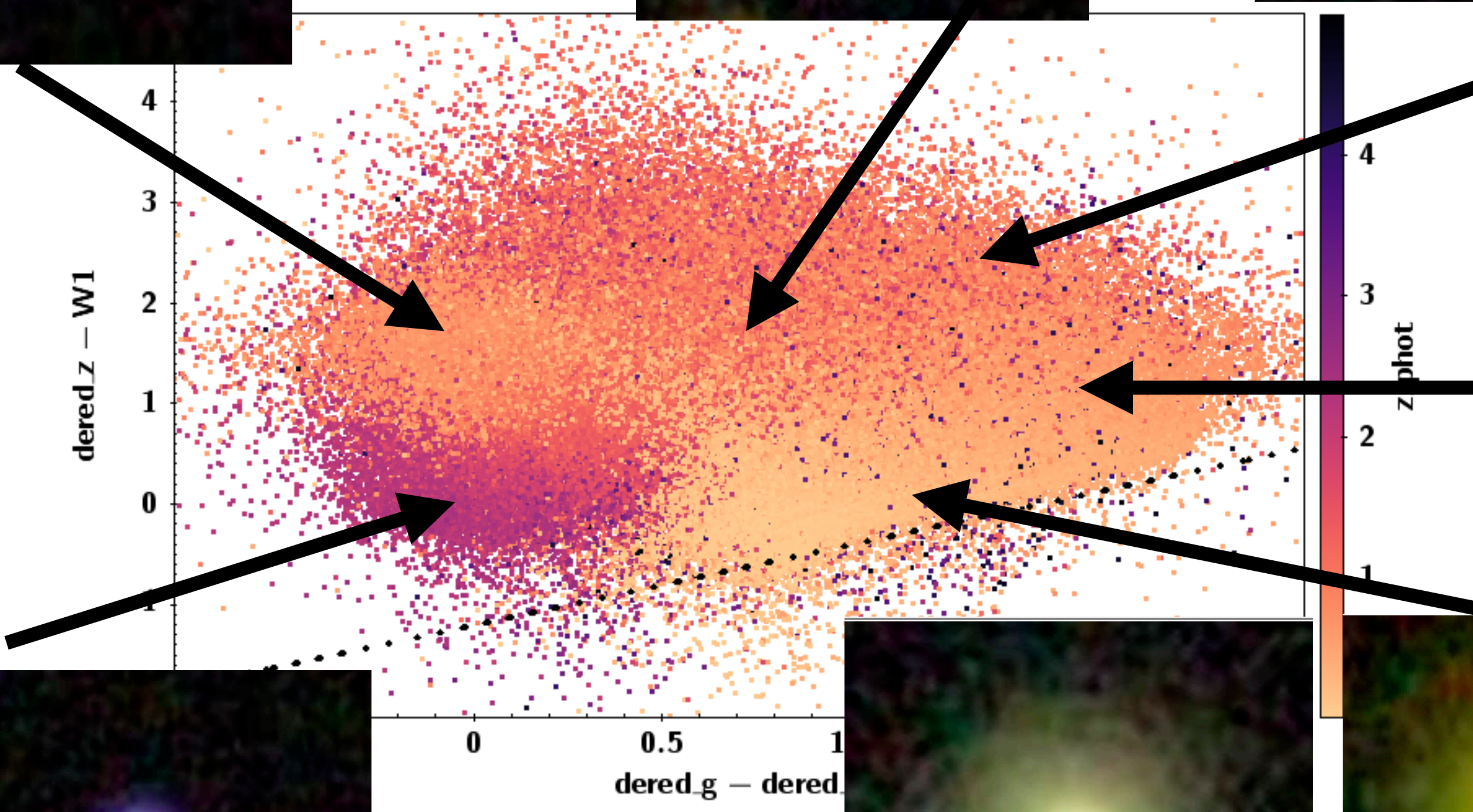
Nearby LLAGN: the “kinetic (radio)” mode of AGN feedback



**eROSITA will cover uniformly the redshift range  $0 < z < 3$**   
**Ideal! Large samples available to study AGN at different  $L$ ,  $z$ ,  $N_H$ ,  $M_*$ , SFR, Radio emission**



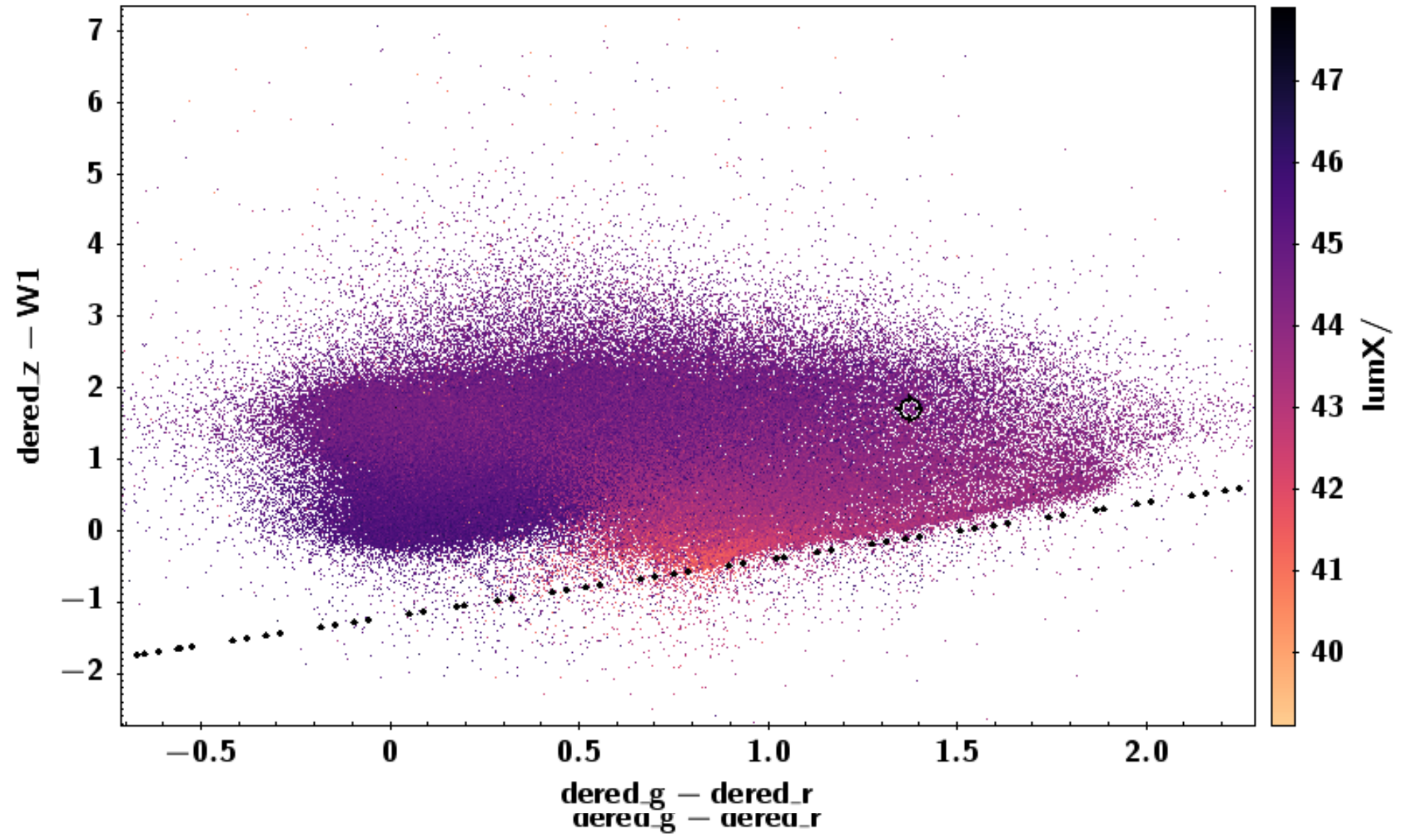
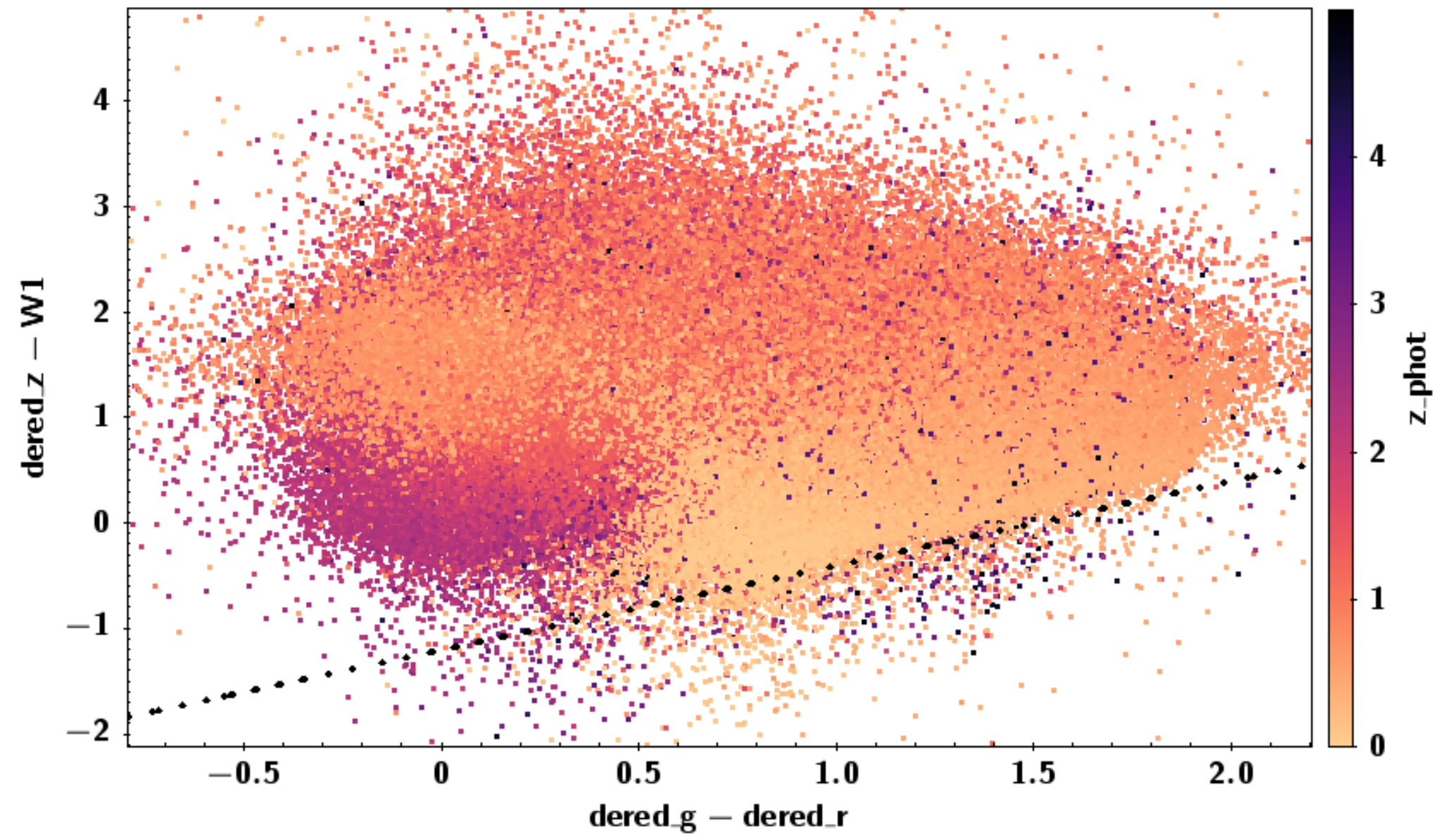
# what AGN and where







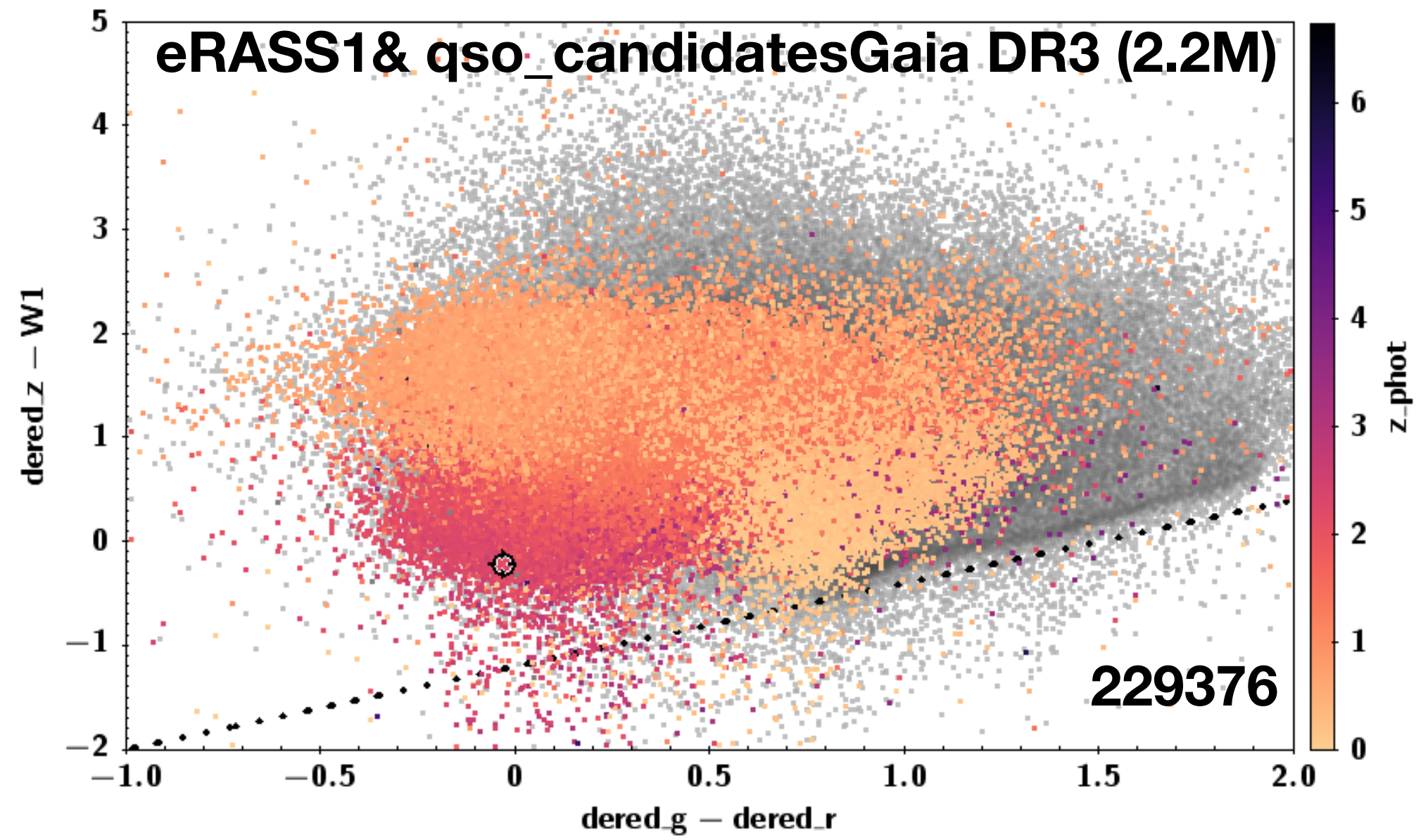
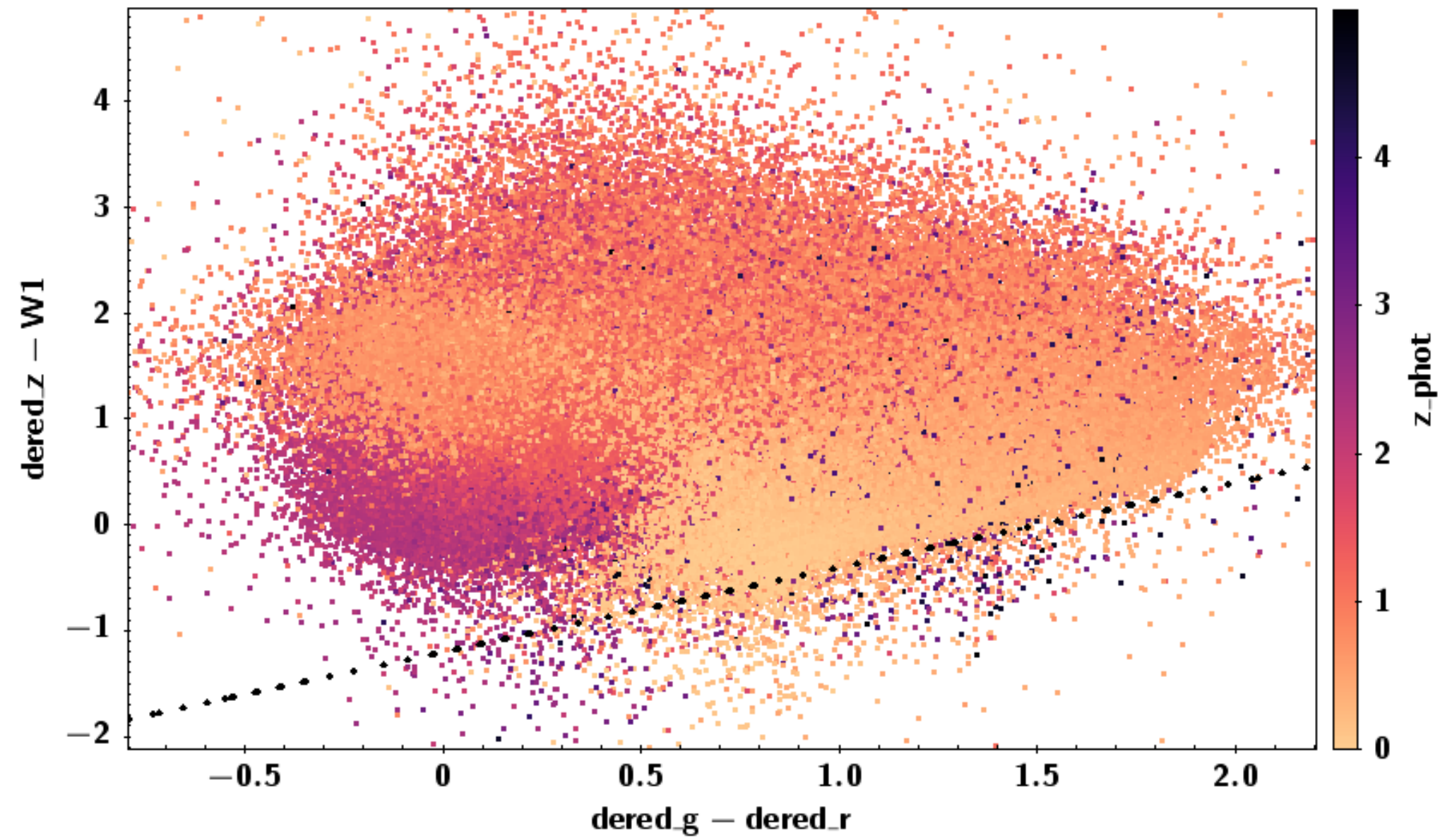
# eRASS1 “shared” AGN (inAllLS10==1)







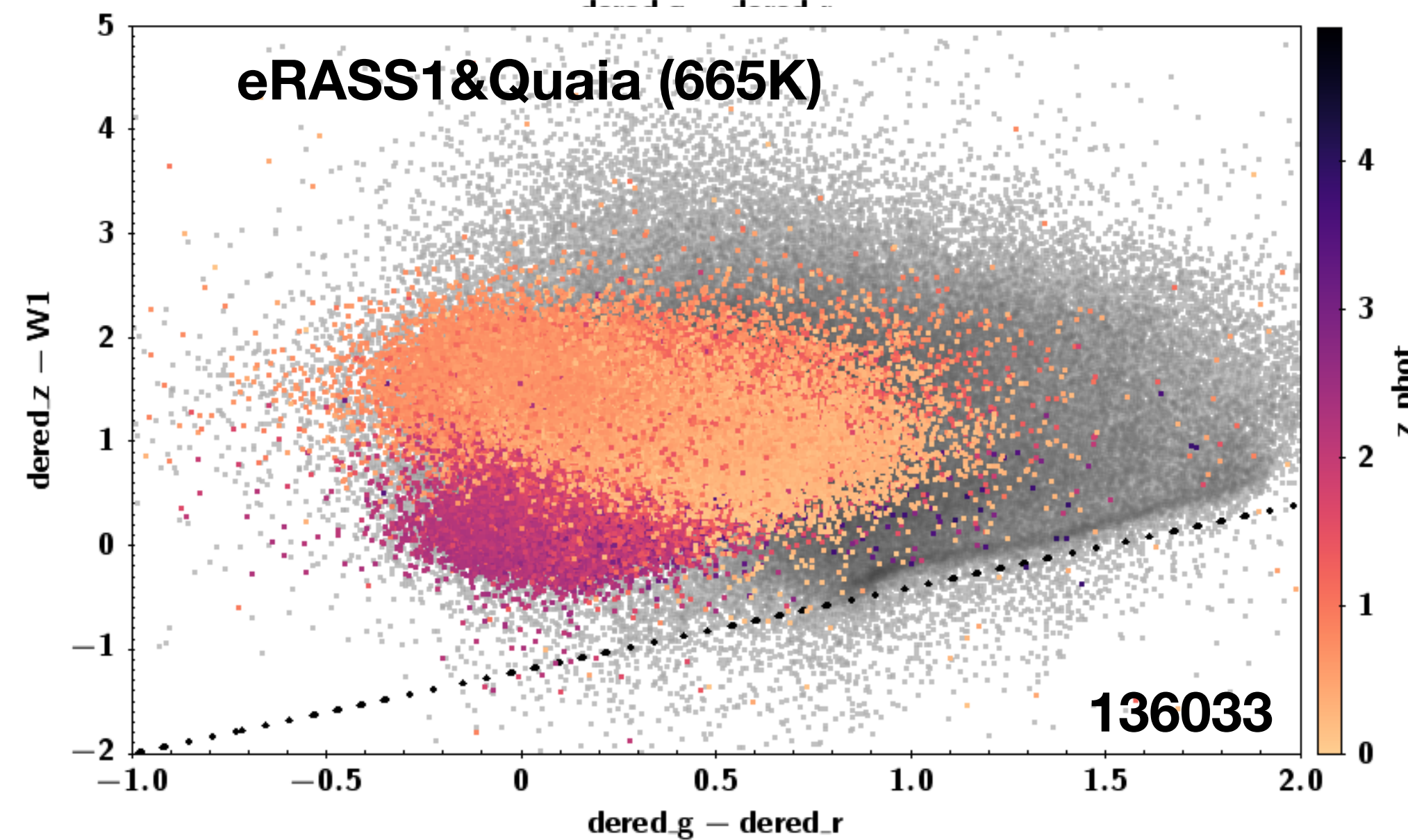
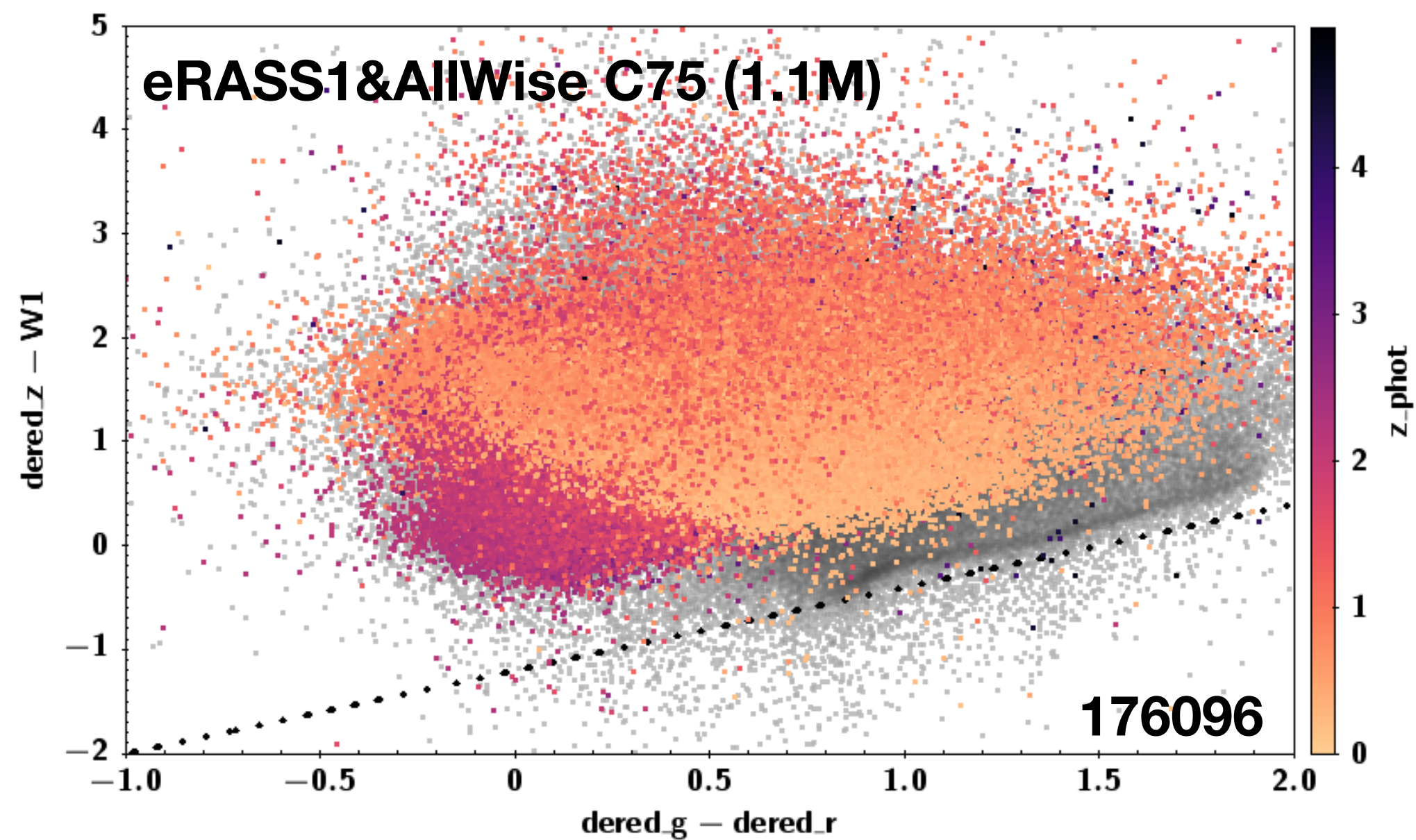
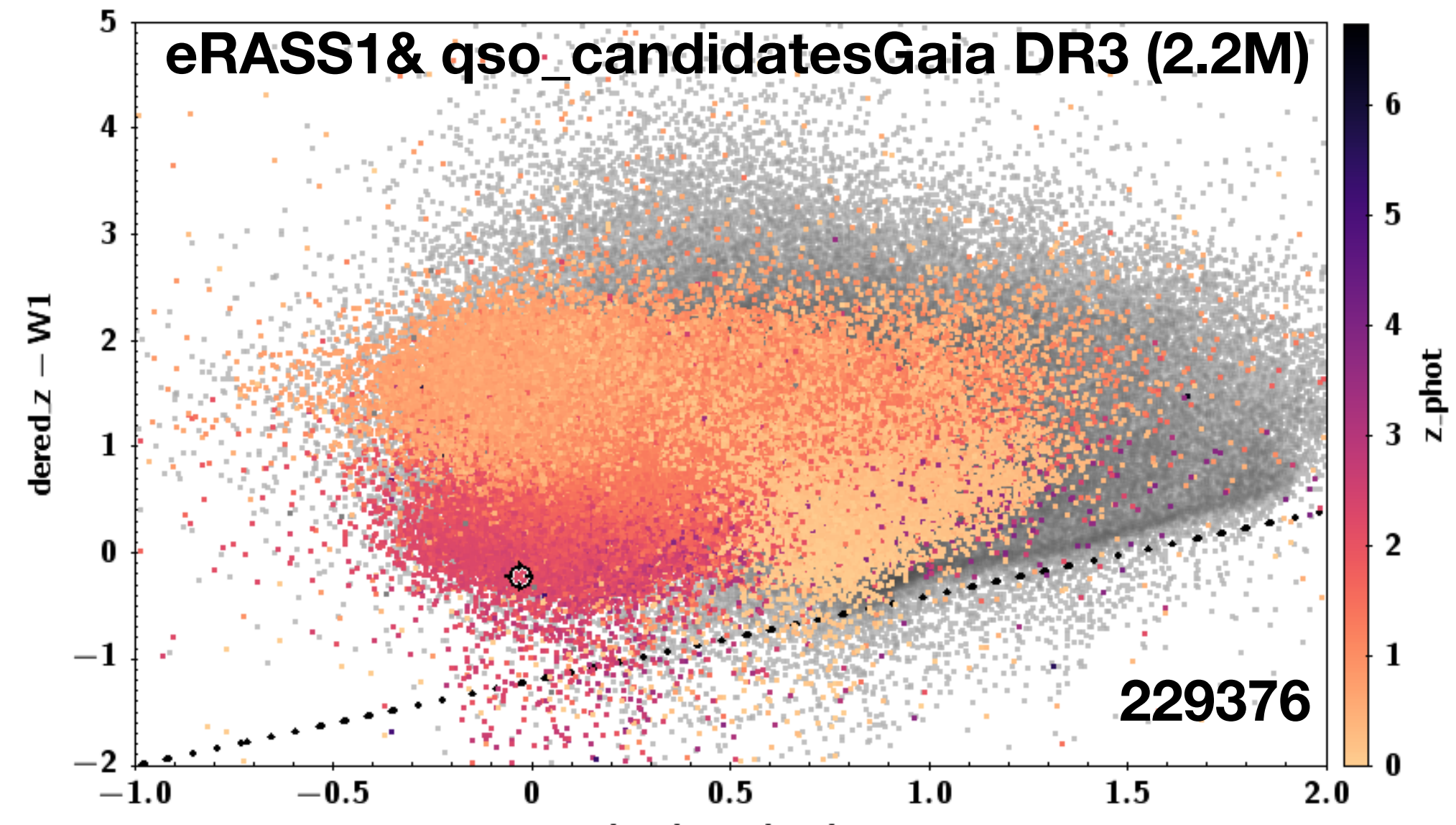
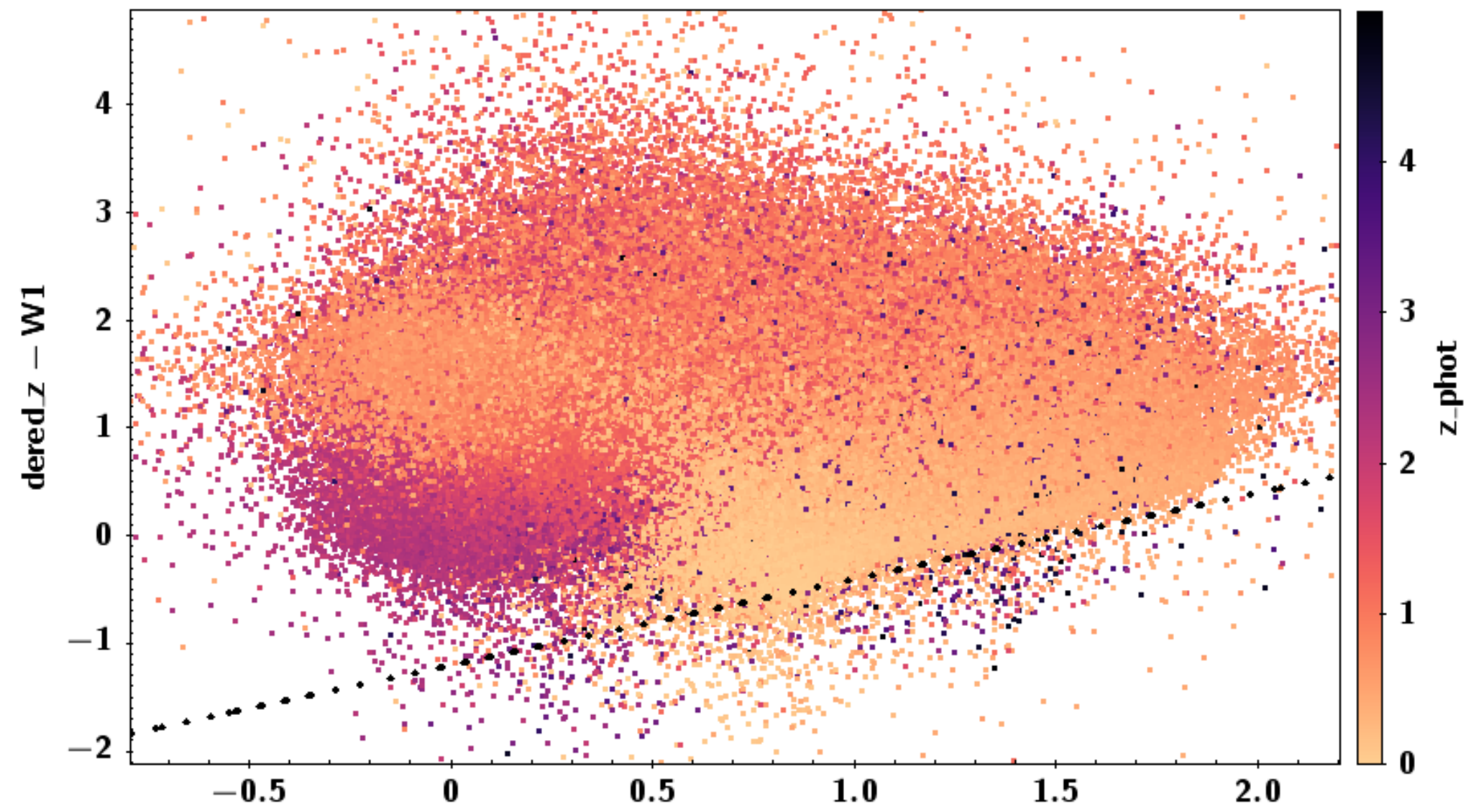
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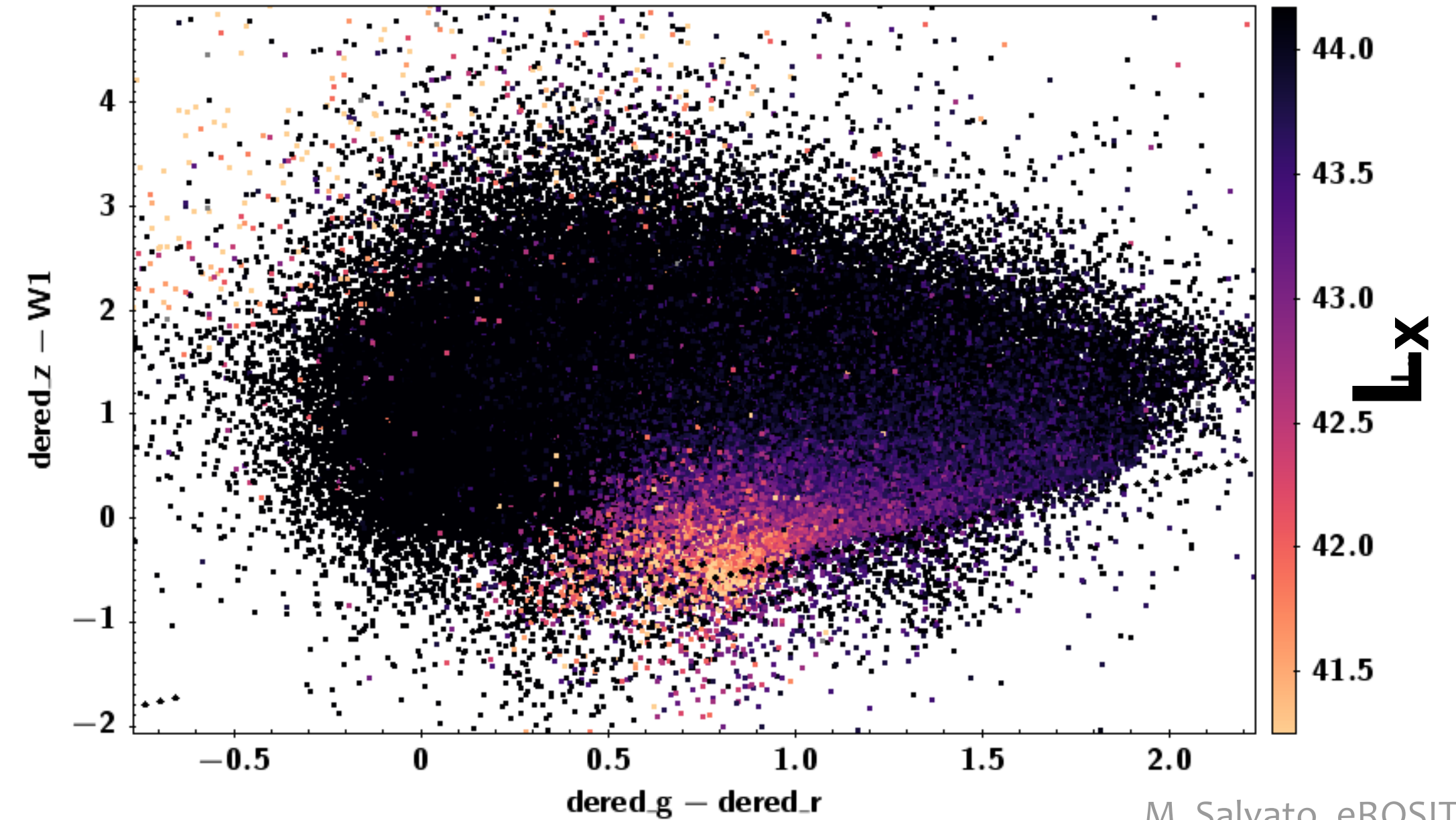
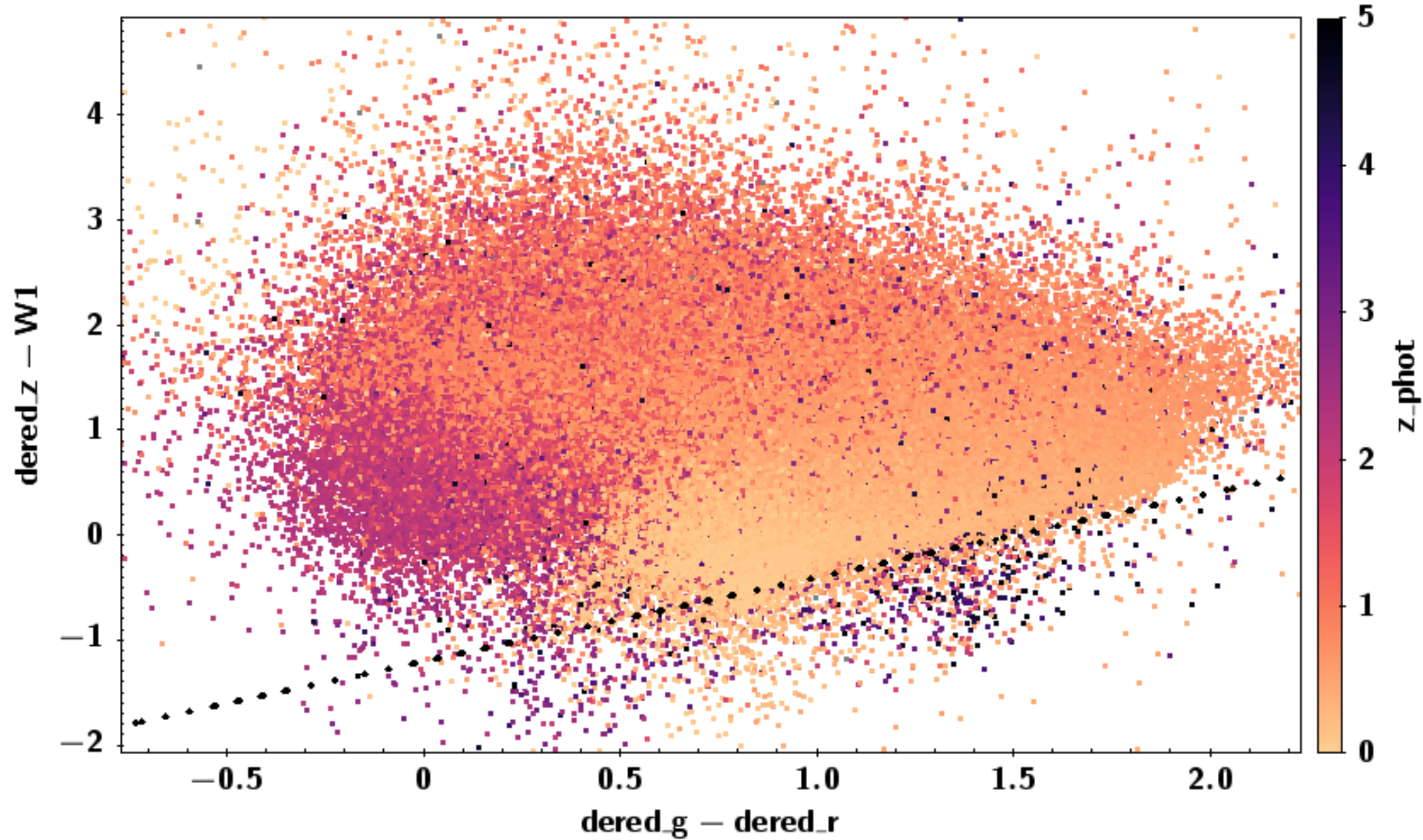
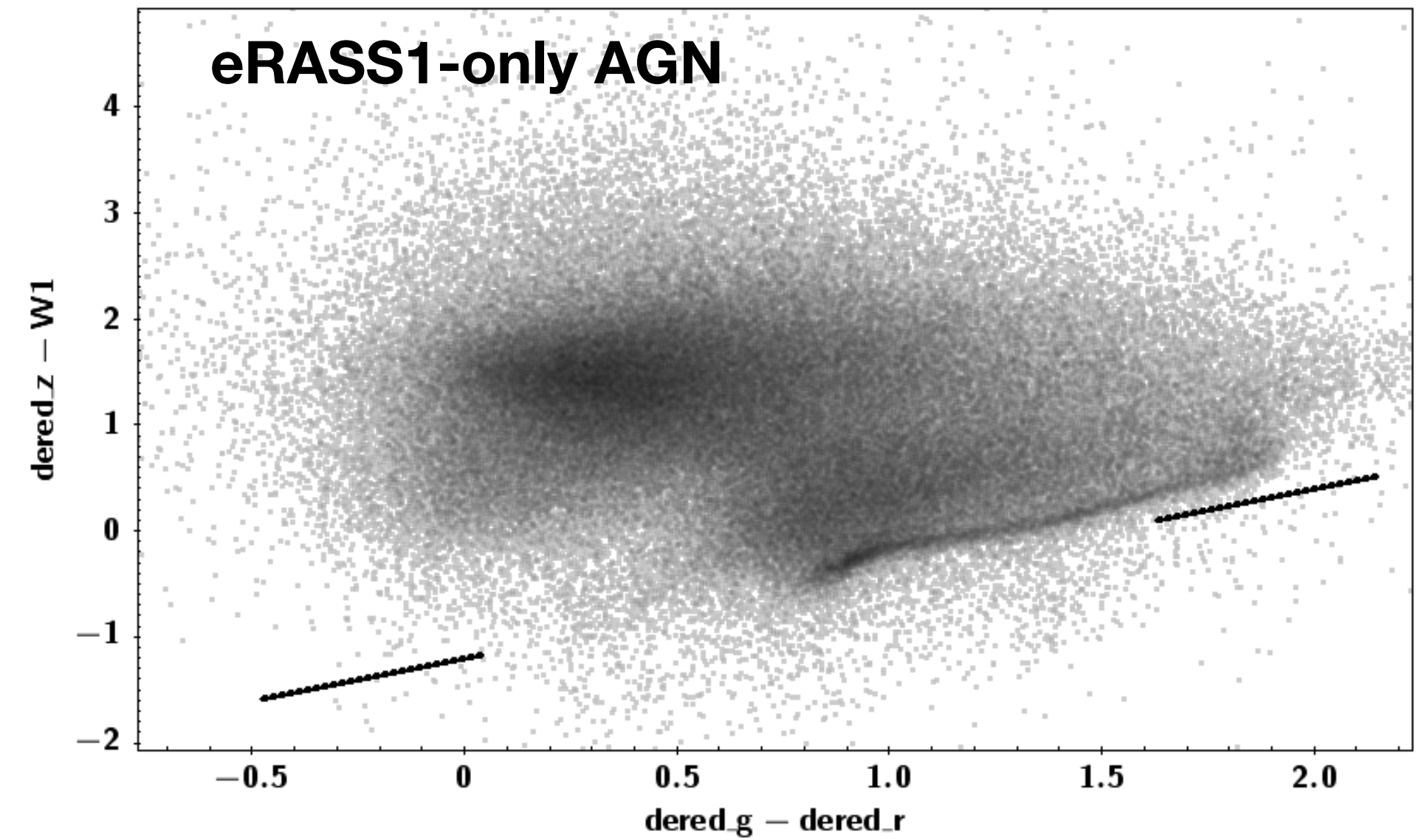
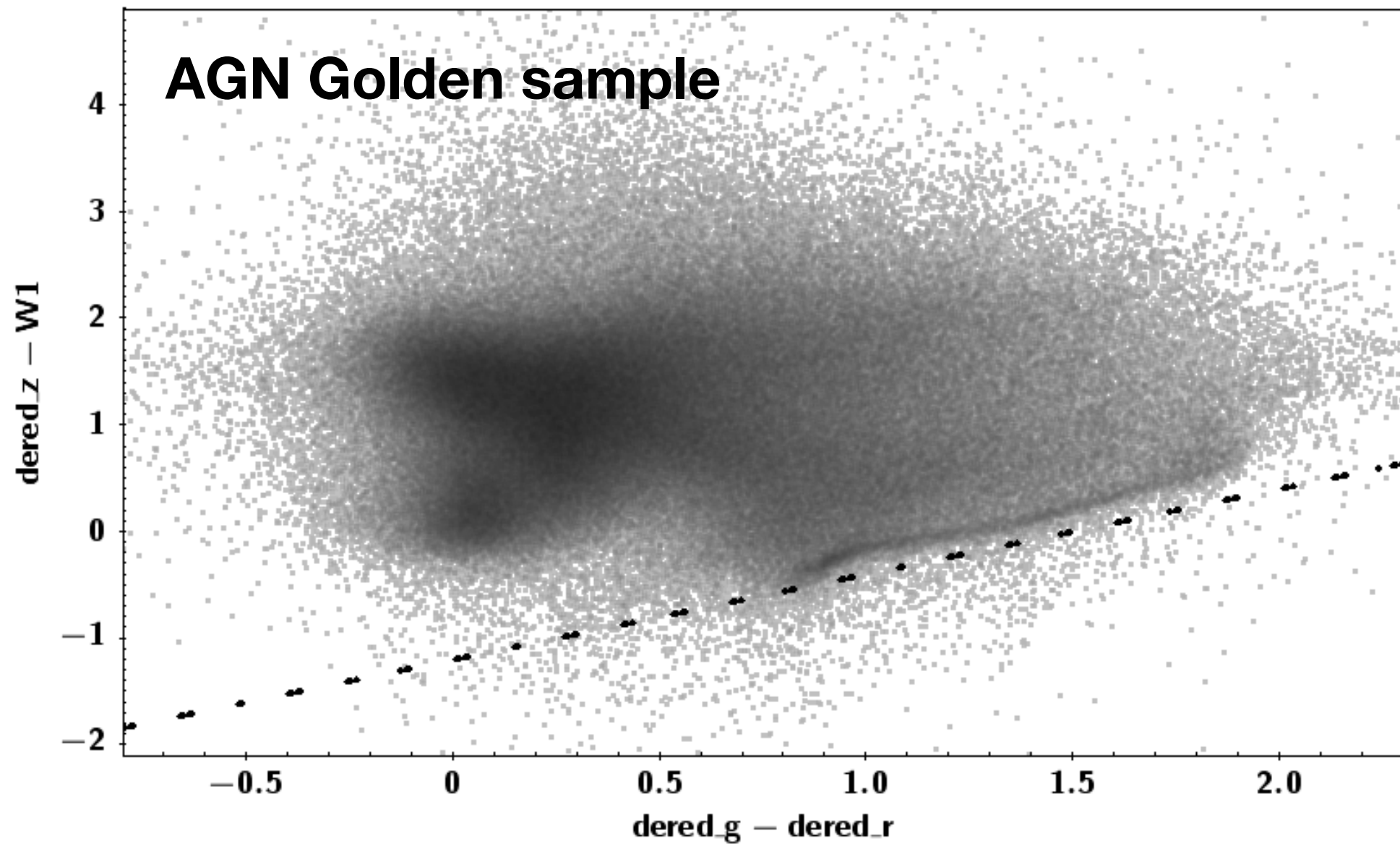


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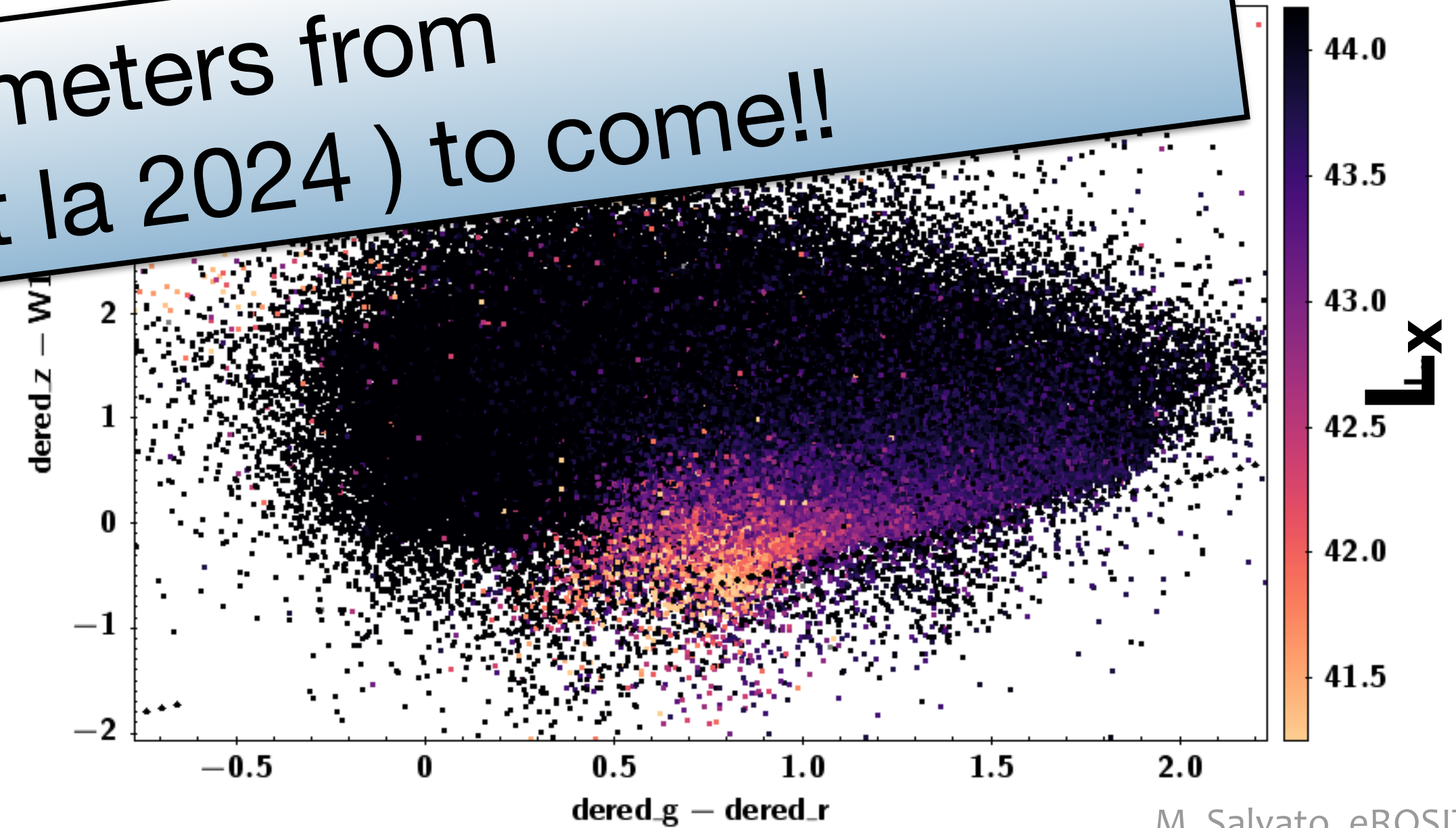
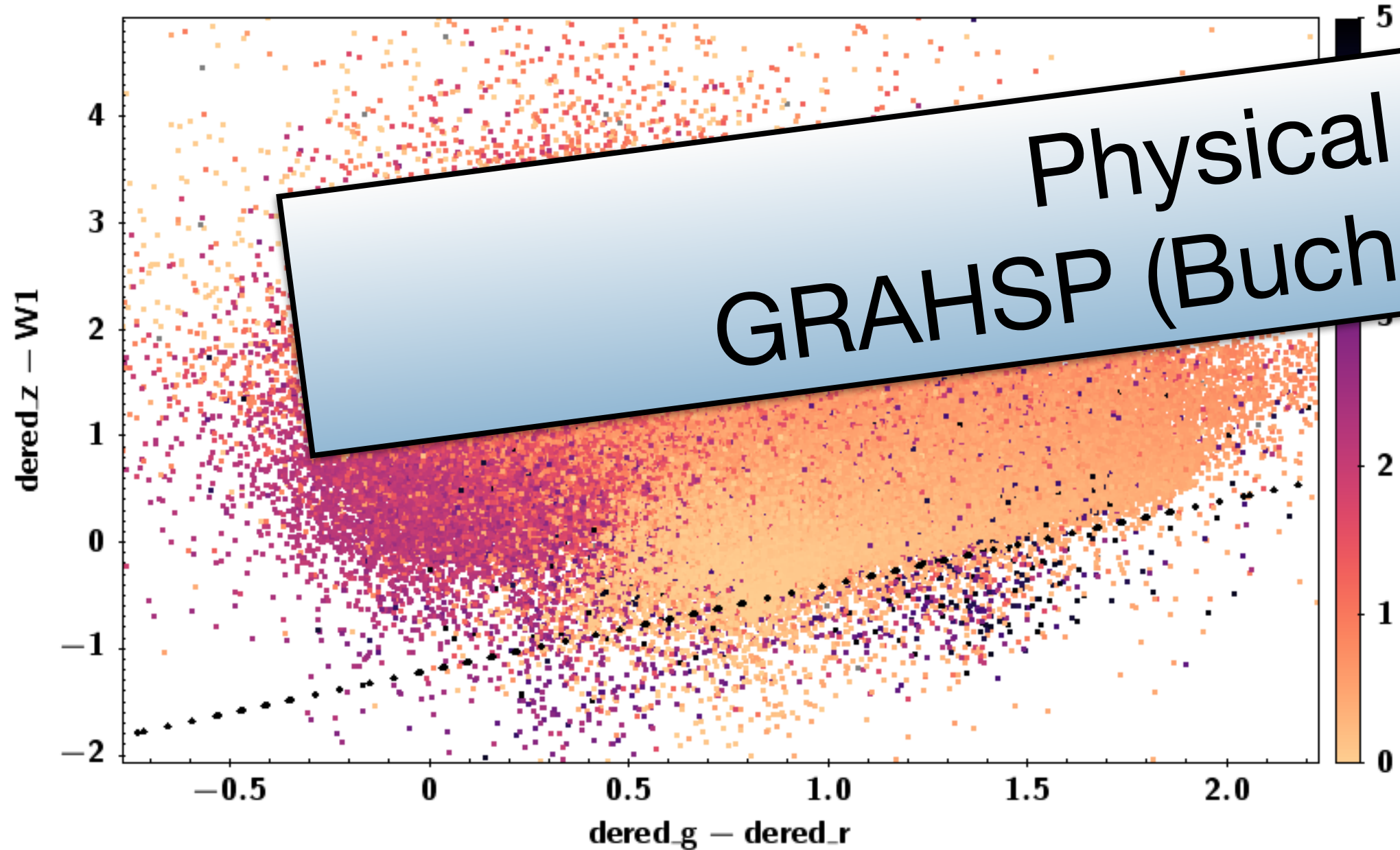
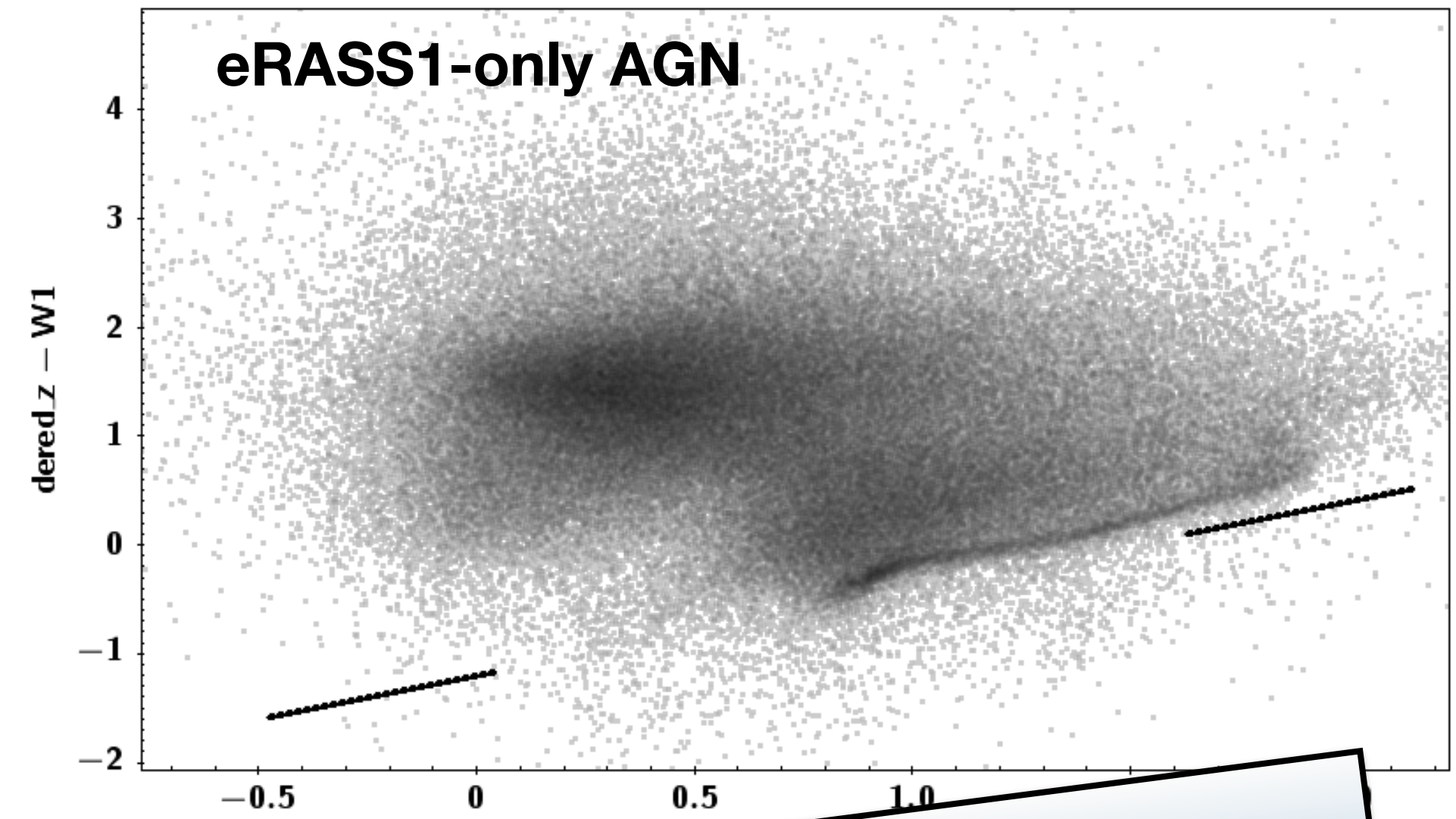
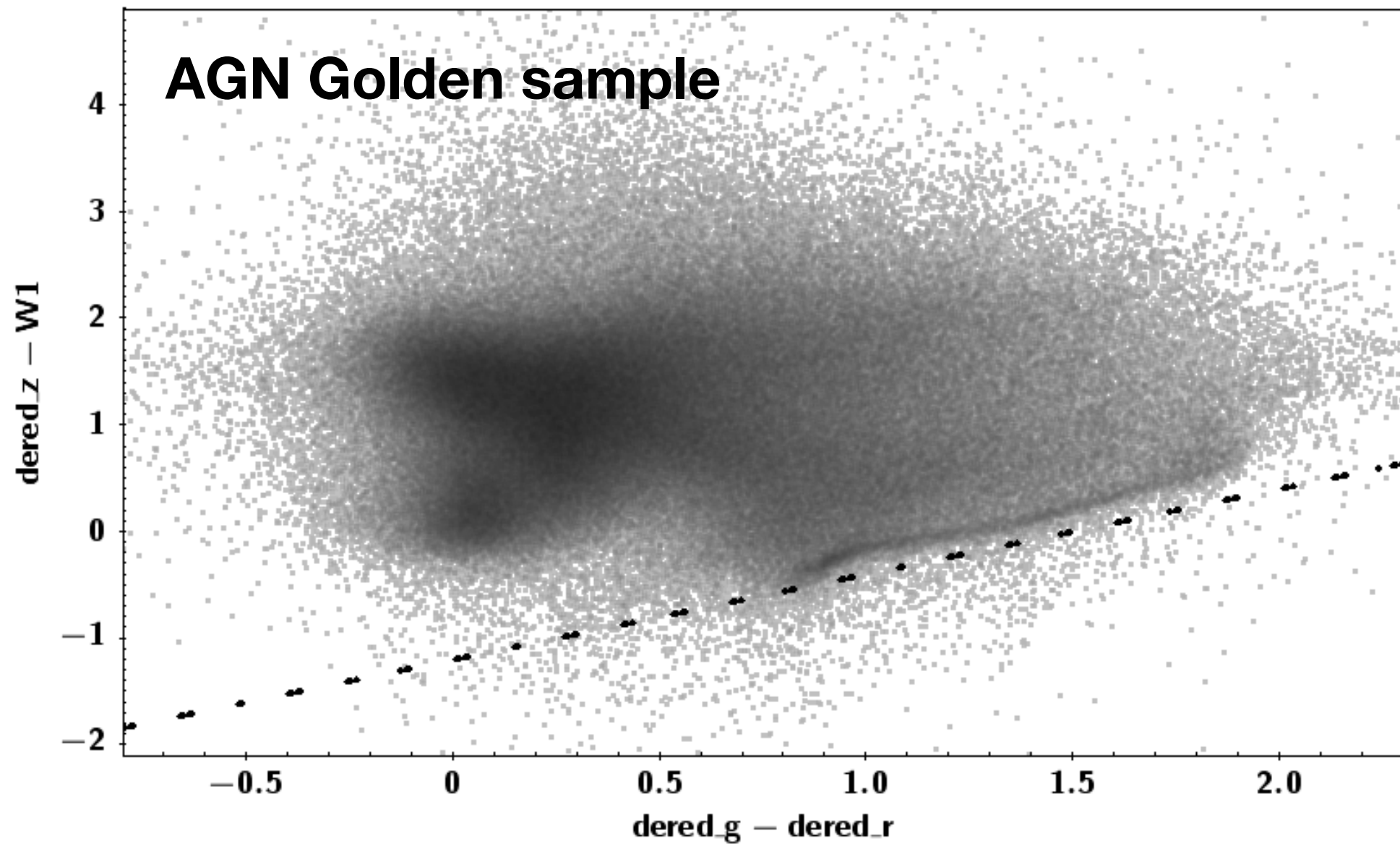


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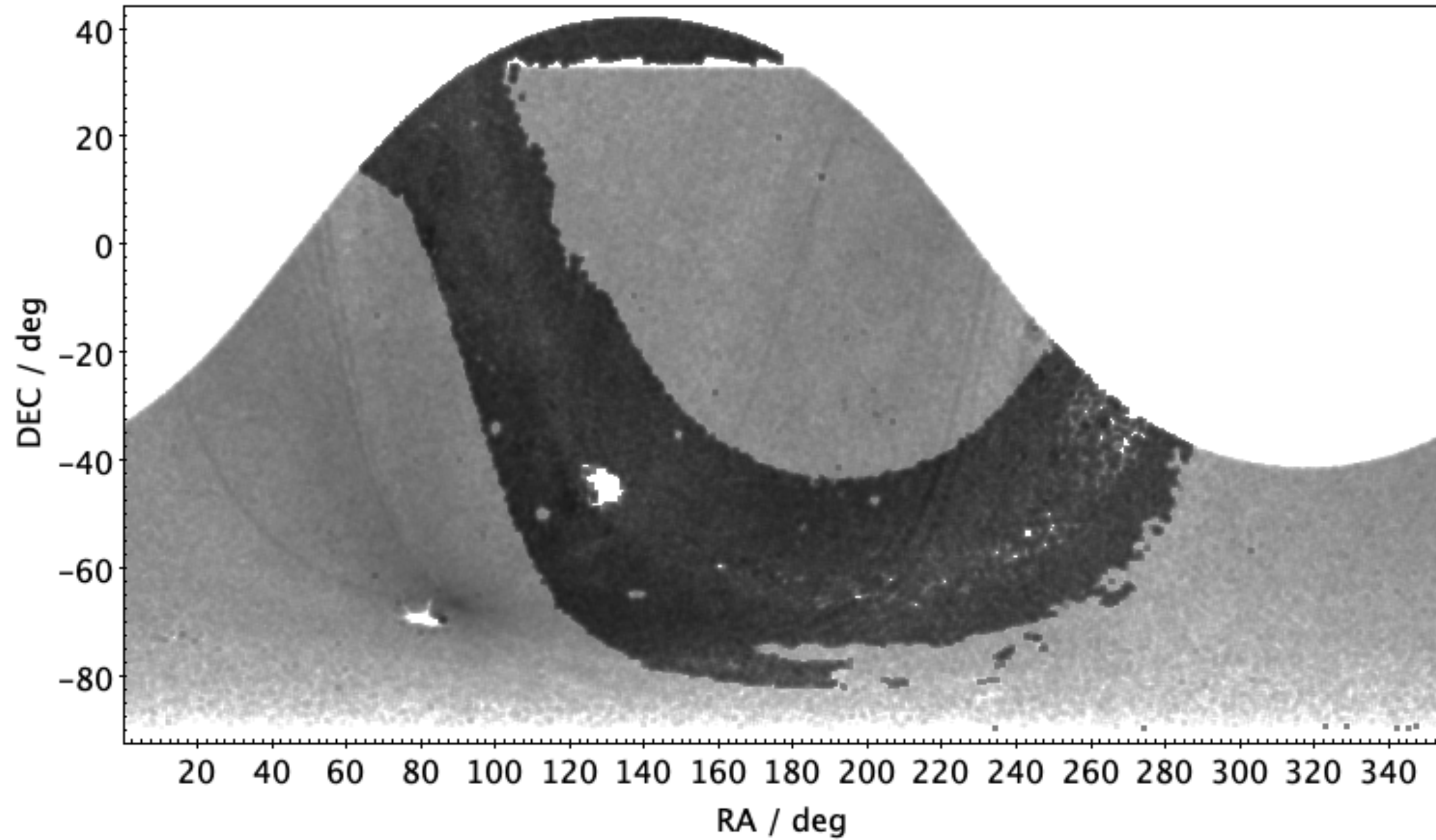


Physical parameters from GRAHSP (Buchner et al 2024) to come!!



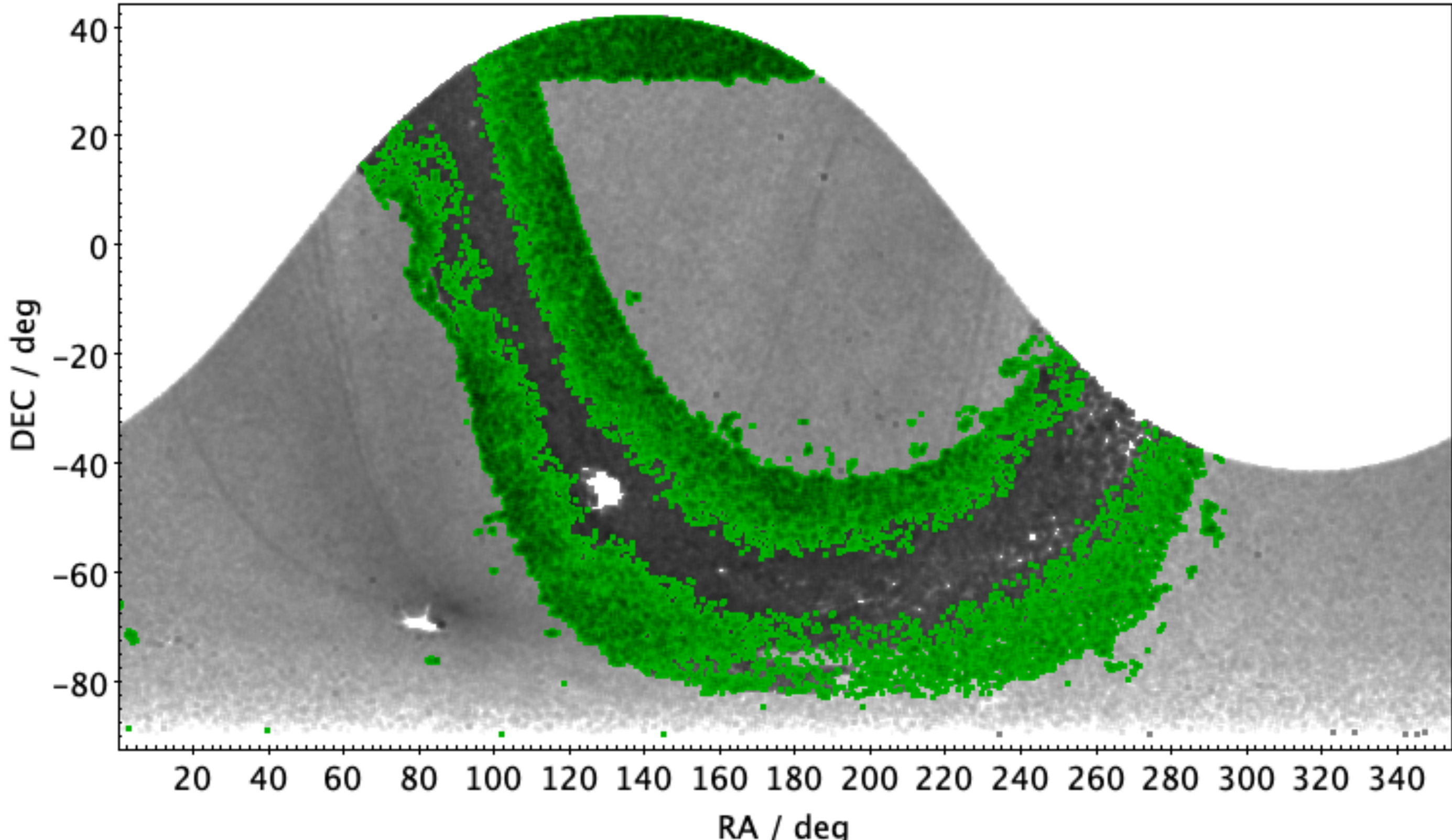


# eRASS1 AGN to be harvested also in the Galactic Plane





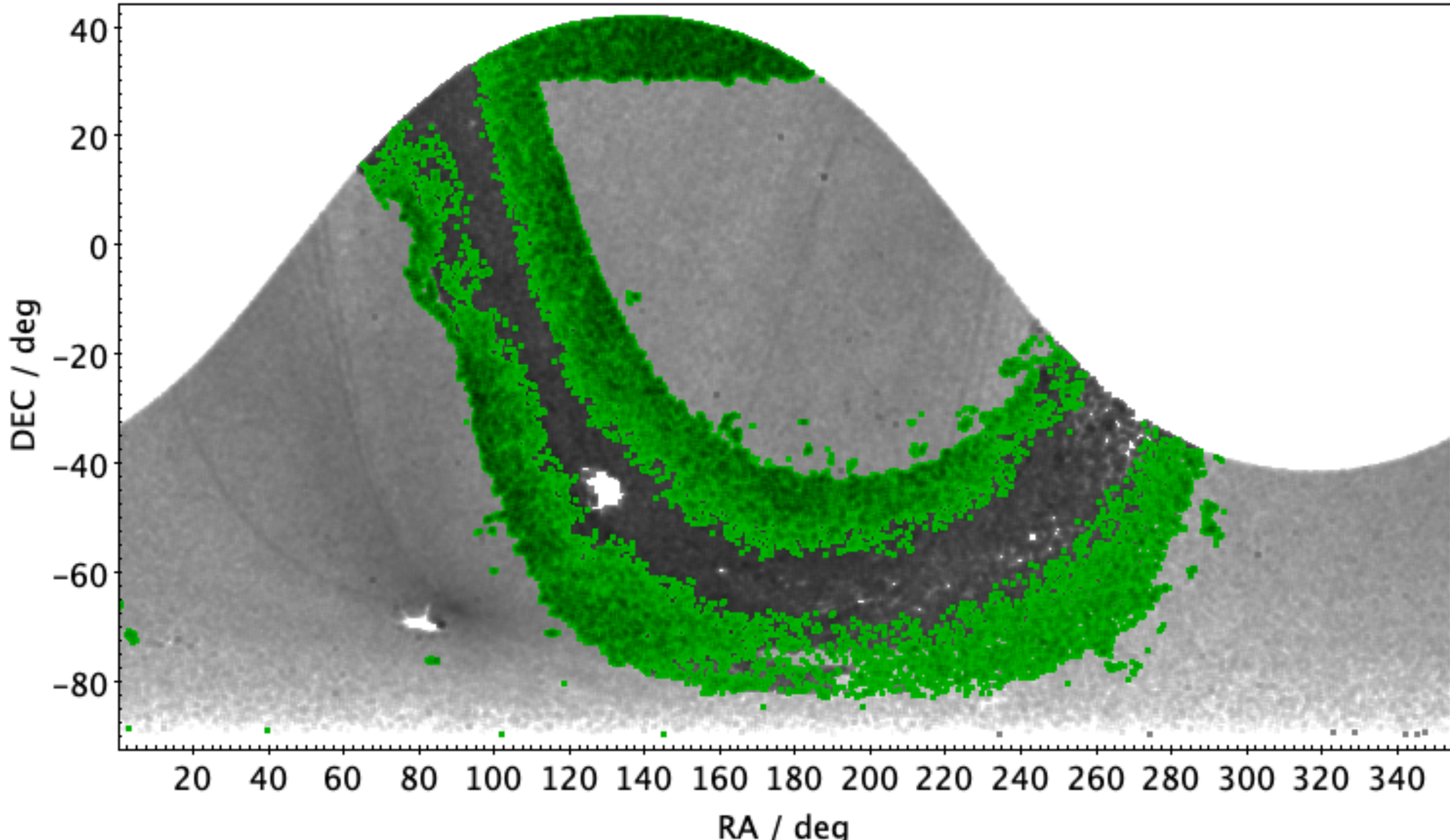
eRASS1 AGN to be harvested also  
in the Galactic Plane



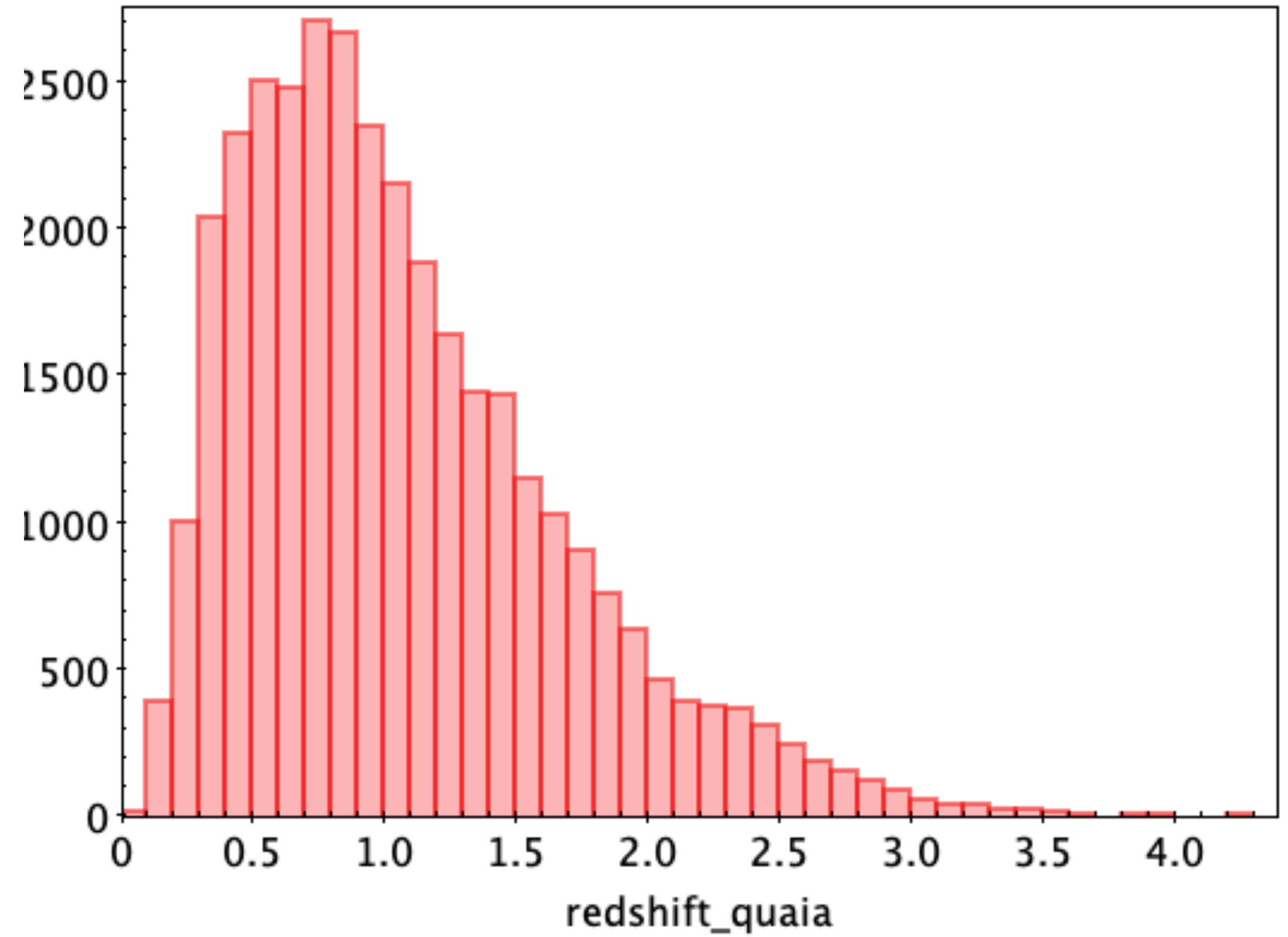
**100k eRASS1 sources with the same CTP in AllWISE\_AGN (C75)  
and qso\_candidates from GDR3**



# eRASS1 AGN to be harvested also in the Galactic Plane



**100k eRASS1 sources with the same CTP in AllWISE\_AGN (C75) and qso\_candidates from GDR3**

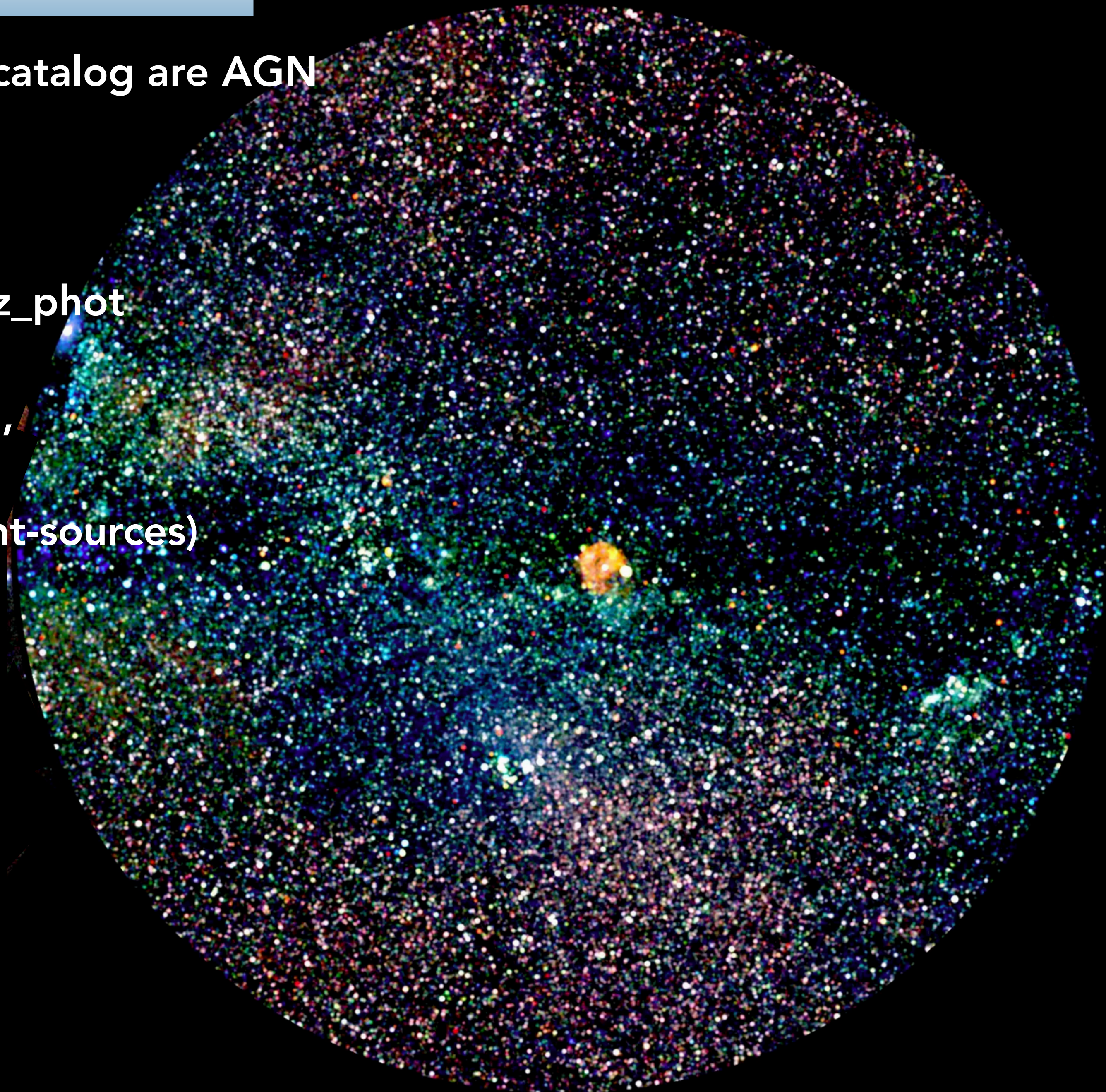


**~36k AGN with redshift from Quiaia**



# Summary

- about 900k point-sources in the eRASS1 main catalog are AGN
- 750k sources are on the footprint of LS10
- 550k are AGN with excellent photometry and  $z_{\text{phot}}$
- about 50% are new AGN non classified as such, by AllWise or Gaia or Quiaia (but AllWISE is shallow and Gaia sees mostly point-sources)
- Additional ~100k are AGN classified as such also by Allwise and Gaia in the Galactic plane (36K with zspec)
- Catalogues are ready and we are finalisizing the paper (Current goal is end of October)





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