# SHORT GRBS FROM STRUCTURED JETS AS EM COUNTERPARTS

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#### Binary NS merger

#### Gravitational waves



Short GRBs



## Prospects of detecting a sGRB associated with GWs

- Rate of observing an on-axis sGRB within LIGO detectability volume (~ 200 Mpc) is ~ I per decade (Wanderman & Piran 2015, Fong et al. 2015).
- sGRB within LIGO volume will be extremely bright (~10<sup>6</sup> cts/s!).

 At these distances, we are more sensitive to off-axis emission and structure of the jet.

## A structured jet

• Numerical and theoretical studies find that a jet develops structure once it breaks out of the collimating medium. (Tchekhovskoy et al. 2010; Komissarov et al. 2010; Sapountzis et al. 2014)



## Simulating the jet

- We perform relativistic MHD simulations using the HARM code (Gammie et al. 2003).
- Simulations are run in 2D (1024x384), axisymmetric, initiated by rotating compact object with magnetization ~ 25.



Emission profile of structured jet

Extracting Lorentz factor and power profile

Calculate observed (radiated) Iuminosity from jet



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#### Emission profile of structured jet



- From the range of  $\Gamma$  and P, can estimate a delay of a few to few tens of seconds (for substantially misaligned  $\theta_{obs}$ ).
- With a coincident LIGO trigger, the count rate required for robust detection ~10<sup>3</sup> cts/s (Connaughton et al. 2016)

#### Structured vs. uniform jet



• When compared to a uniform jet, emission from structured jet detectable for much larger range of  $heta_{
m obs}$  .

## Conclusions

- Prompt emission from structured jet detectable by substantially misaligned observers, with the aid of a coincident LIGO trigger.
- Consequently, rates of observing sGRB within LIGO volume can be ~ few per year.
- Prompt detection can localize the source and aid in the follow-up detection of delayed signals.

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